Atomic Spectra

IB PHYSICS | ATOMIC PHYSICS

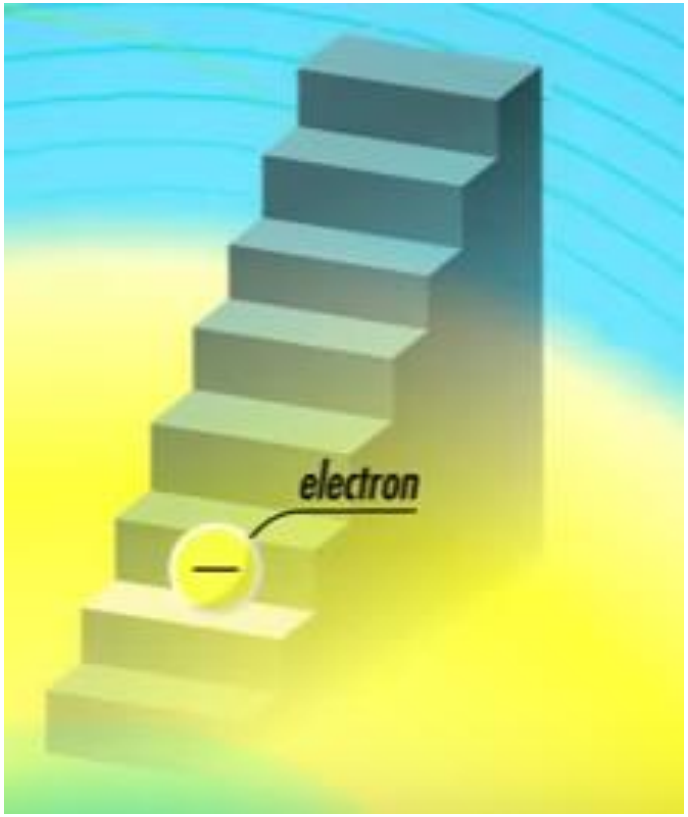
What is Light?

Light is Quantized

Photons of light can only have certain

_____

values of energy



Energy of a Photon

$$E = hf$$

Planck's Constant	h	6.63×10^{-34} J s

Energy of a Photon

$$E = hf \qquad c = f\lambda$$

$$c = 3.00 \times 10^{8}\ m\ s^{-1}$$

Quick Recap of eV

$eV \rightarrow$

IB Physics Data Booklet

Sub-topic 7.1 – Discrete energy and radioactivity	Sub-topic 7.2 – Nuclear reactions
$E = hf$ $\lambda = \frac{hc}{E}$	$\Delta E = \Delta m\, c^2$

Stefan–Boltzmann constant	σ	$5.67 \times 10^{-8}\,\mathrm{W\,m^{-2}\,K^{-4}}$
Coulomb constant	k	$8.99 \times 10^{9}\,\mathrm{N\,m^{2}\,C^{-2}}$
Permittivity of free space	ε_0	$8.85 \times 10^{-12}\,\mathrm{C^{2}\,N^{-1}\,m^{-2}}$
Permeability of free space	μ_0	$4\pi \times 10^{-7}\,\mathrm{T\,m\,A^{-1}}$
Speed of light in vacuum	c	$3.00 \times 10^{8}\,\mathrm{m\,s^{-1}}$
Planck's constant	h	$6.63 \times 10^{-34}\,\mathrm{J\,s}$
Elementary charge	e	$1.60 \times 10^{-19}\,\mathrm{C}$

Try This…

Calculate the energy carried by one photon of microwaves of wavelength 9 cm (as might be used in wifi signals) in J and eV

Shortcut time ☺

Unit conversions

1 radian (rad) $\equiv \frac{180^\circ}{\pi}$

Temperature (K) = temperature (°C) + 273

1 light year (ly) = 9.46×10^{15} m

1 parsec (pc) = 3.26 ly

1 astronomical unit (AU) = 1.50×10^{11} m

1 kilowatt-hour (kWh) = 3.60×10^{6} J

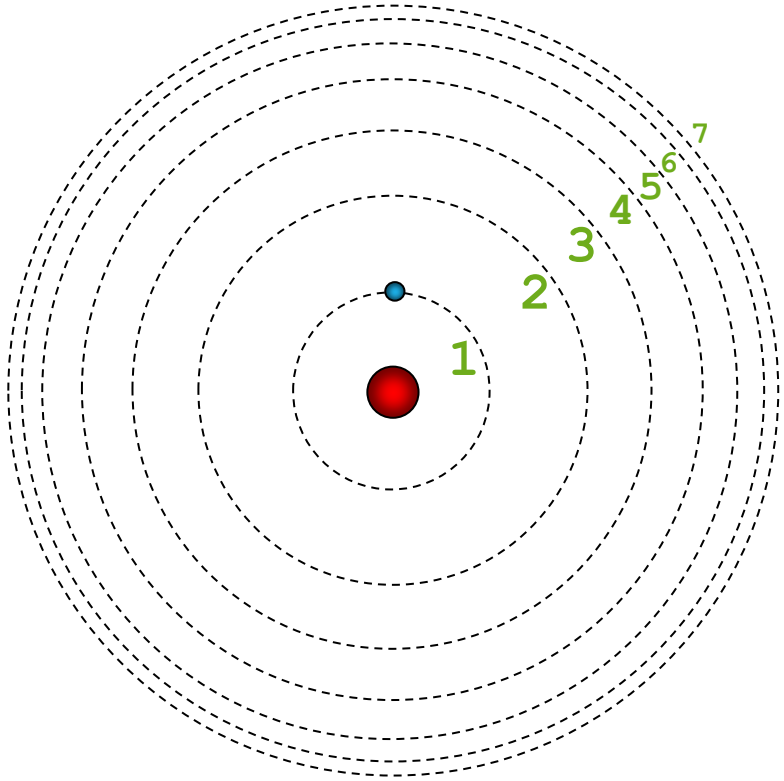
$hc = 1.99 \times 10^{-25}$ J m = 1.24×10^{-6} eV m

Since *h* and *c* are both constants, *hc* acts as a constant as well

$$E = \frac{hc}{\lambda}$$

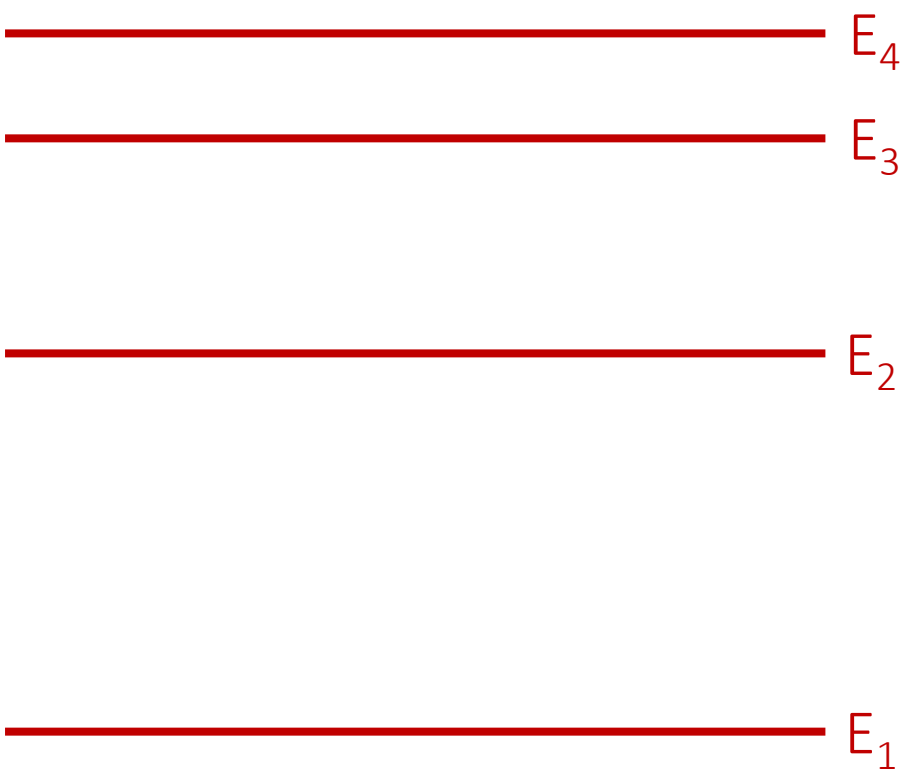
Energy Levels

Electrons in an atom exist at discrete energy levels



Energy Levels

A photon is emitted whenever an electron transitions from one energy level down to a lower energy level

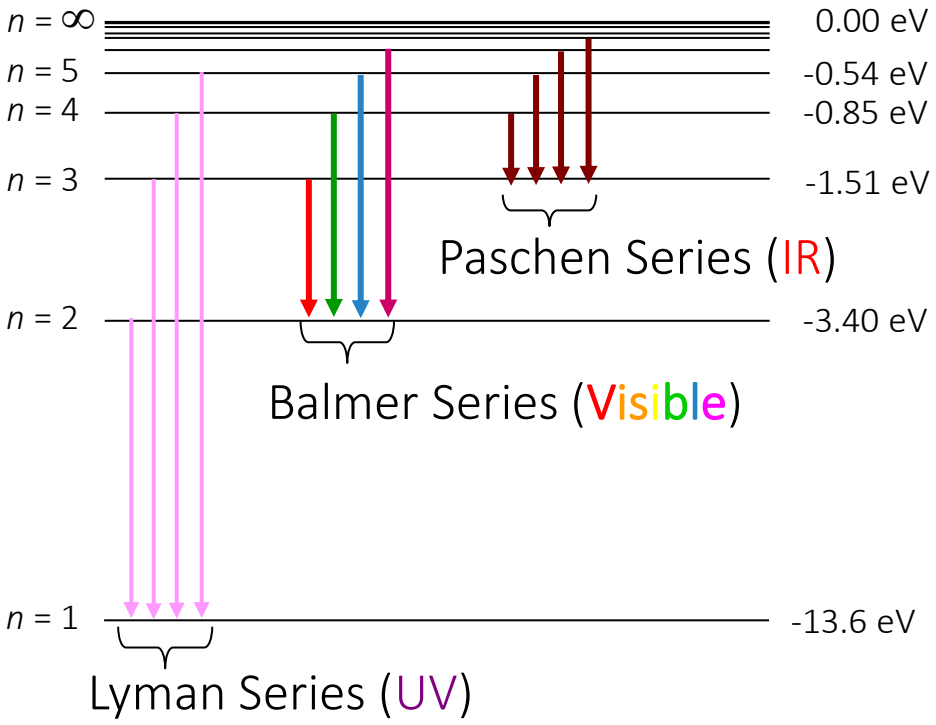


How many different transitions are possible between these four energy levels?

Energy Levels

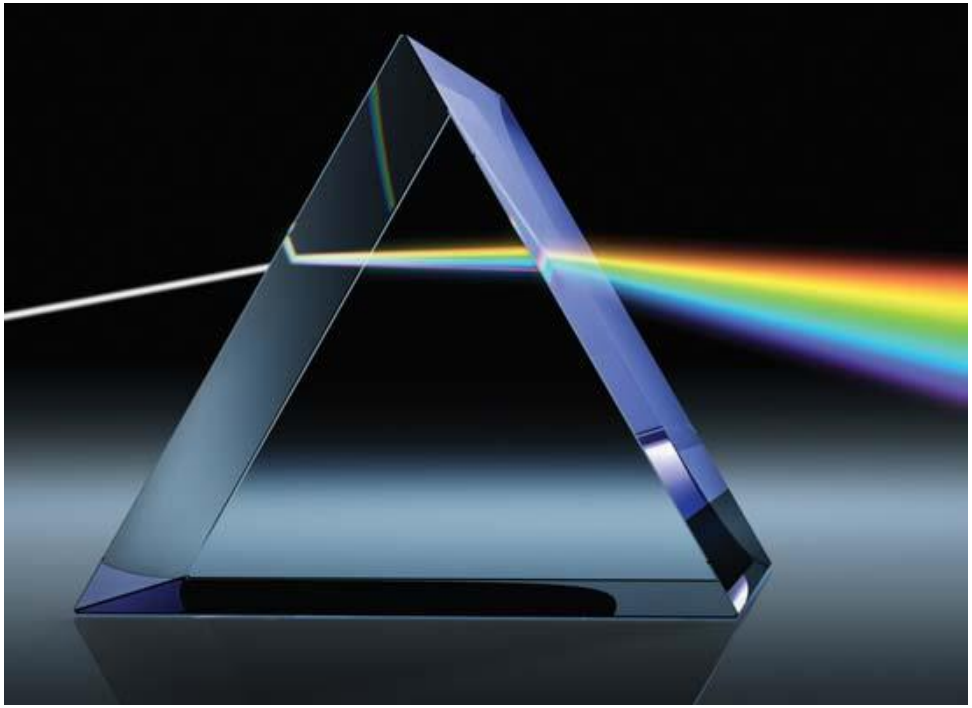
Energy Transitions

Different Energy transitions result in different energies (wavelengths) of light that are absorbed or emitted



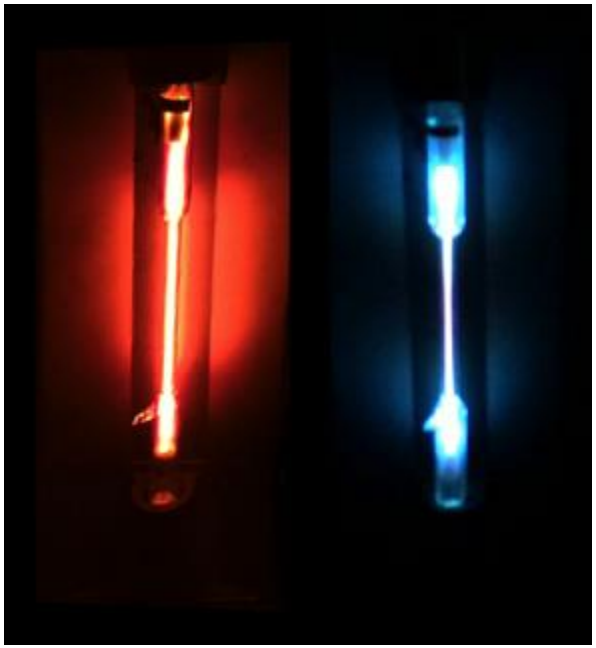
Continuous Spectrum

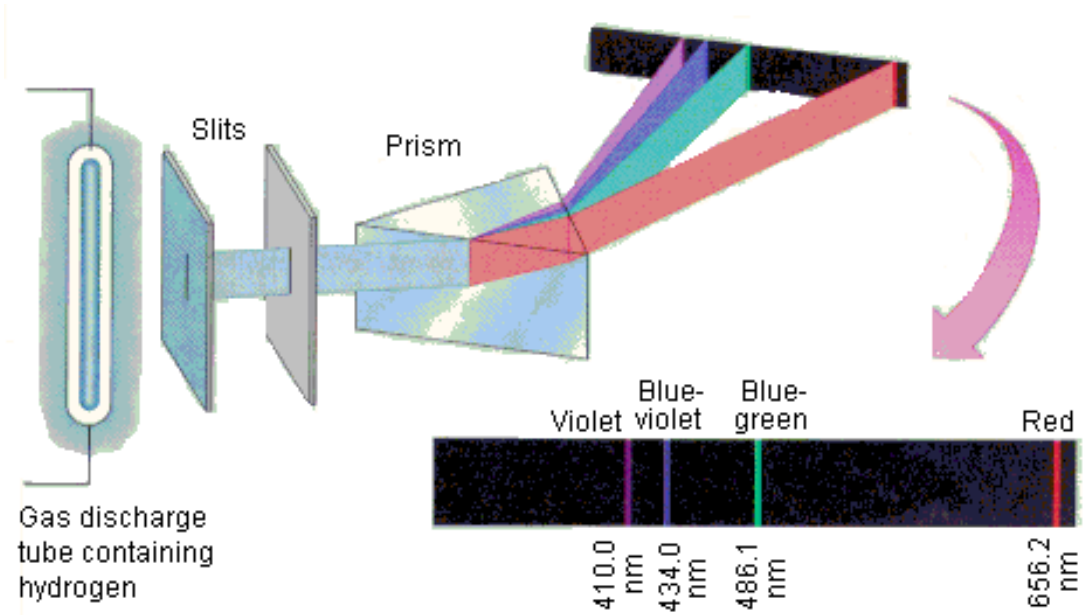
When white light from the sun passes through a prism, the light is dispersed into its component colors in a continuous spectrum



Emission Spectrum

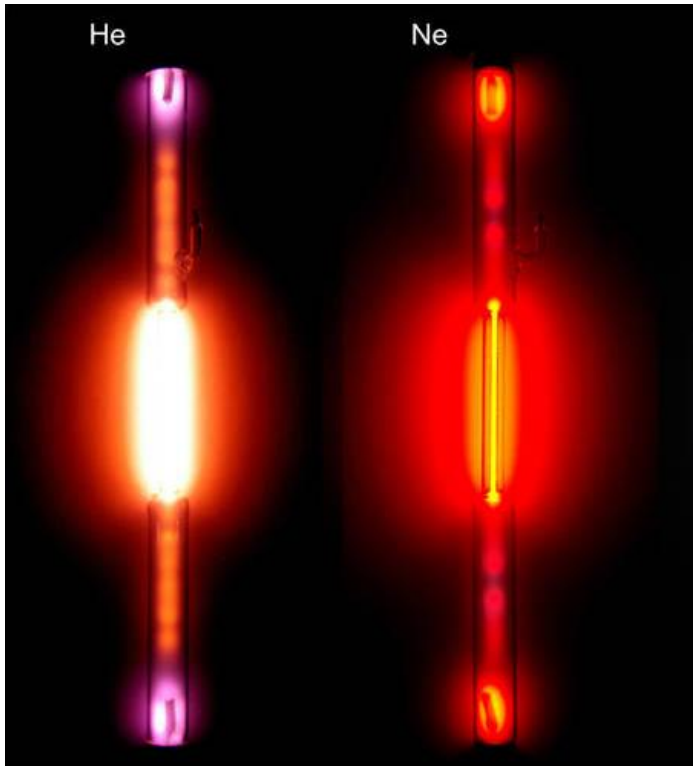
If an electric current is passed through an element in the form of a low-pressure gas, it will produce its own unique emission spectrum

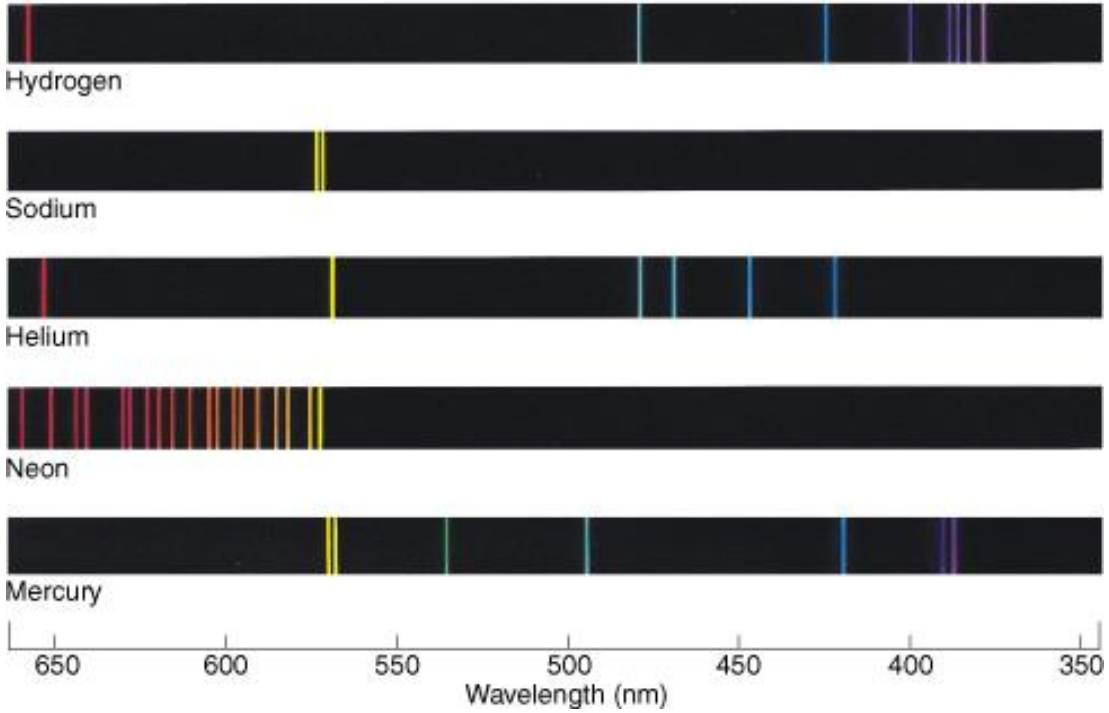




Emission Spectrum

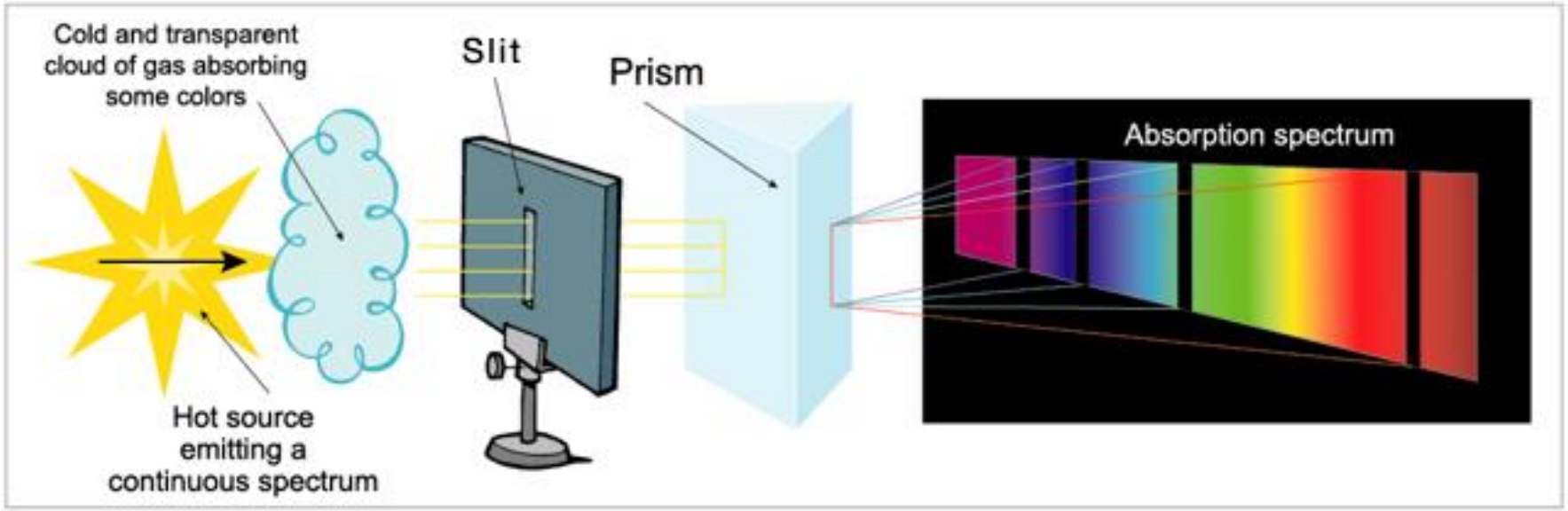
These spectra can be used to identify elements like a fingerprint



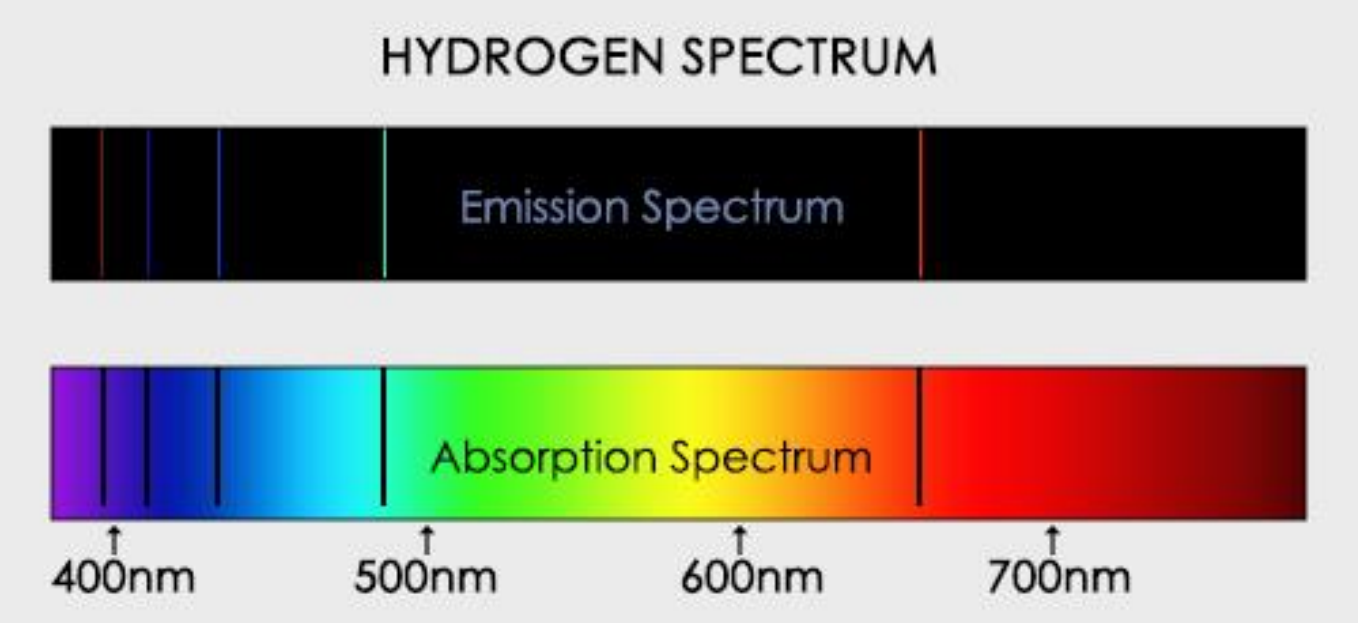


Absorption Spectrum

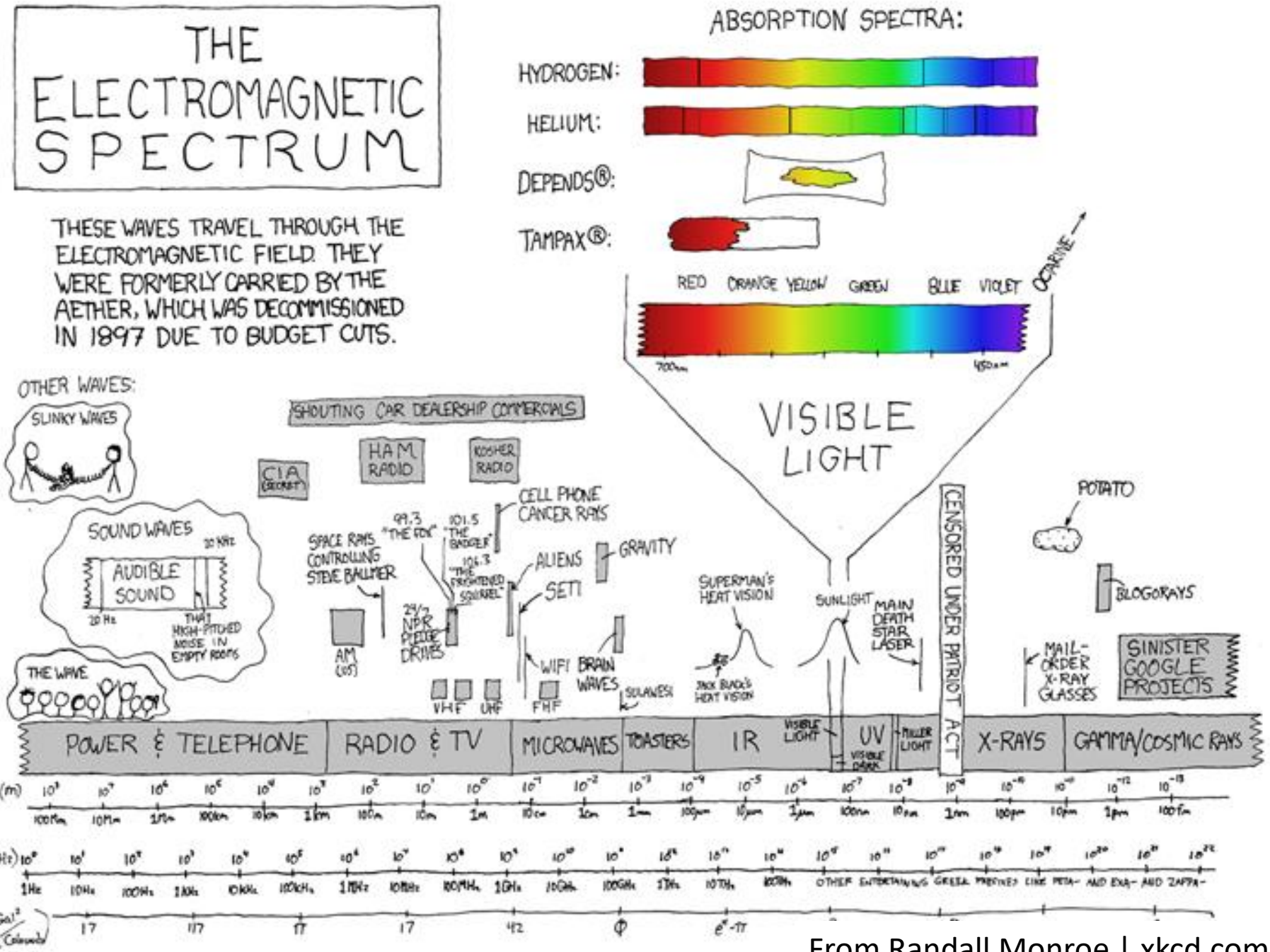
If white light is passed through a sample of gaseous atoms or molecules, it is found that the light of certain wavelengths is missing



Absorption Spectrum

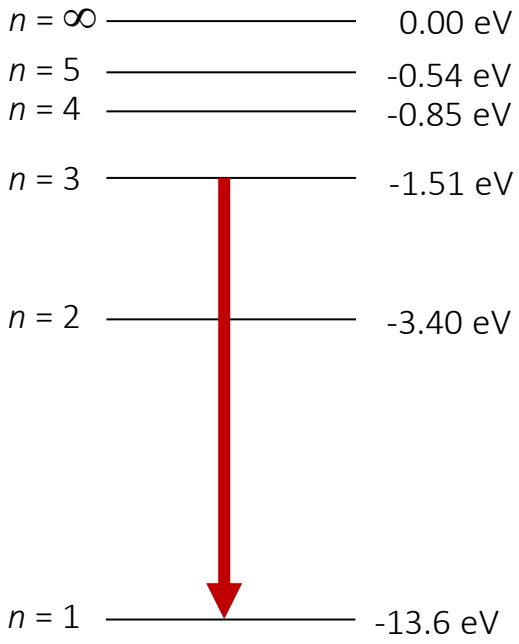


The emission and absorption spectra are negative images of each other



From Randall Monroe | xkcd.com

Calculating Wavelength Emitted

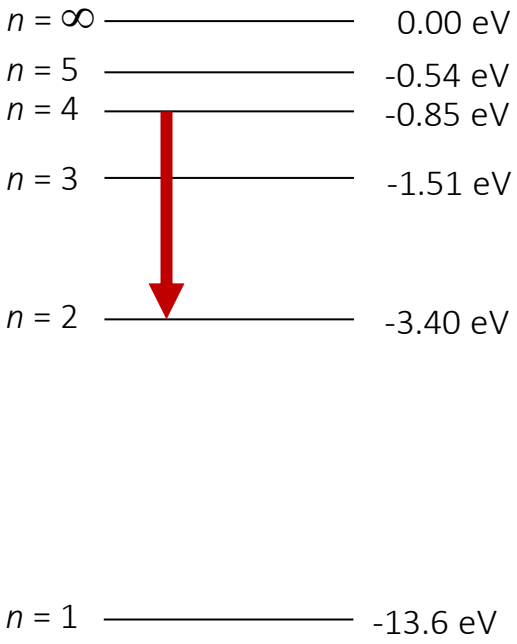


What is the wavelength emitted?

$$\lambda = \frac{hc}{E}$$

hc	1.99×10^{-25} J m	1.24×10^{-6} eV m

Try This…



What is the wavelength emitted?

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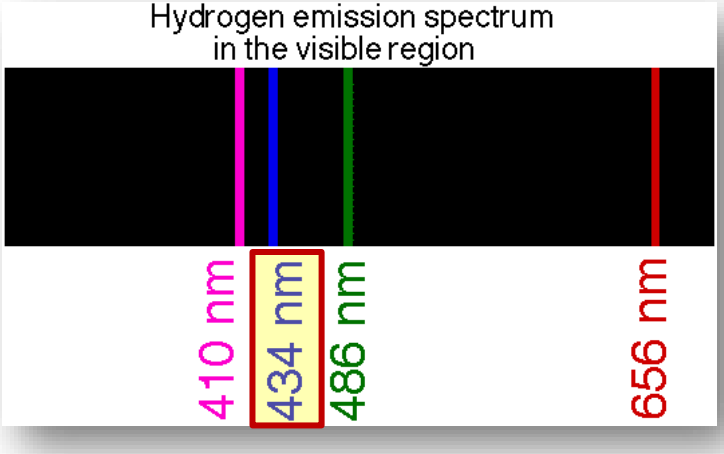
Working Backwards…

What is the energy in eV for a 434 nm blue emission line?



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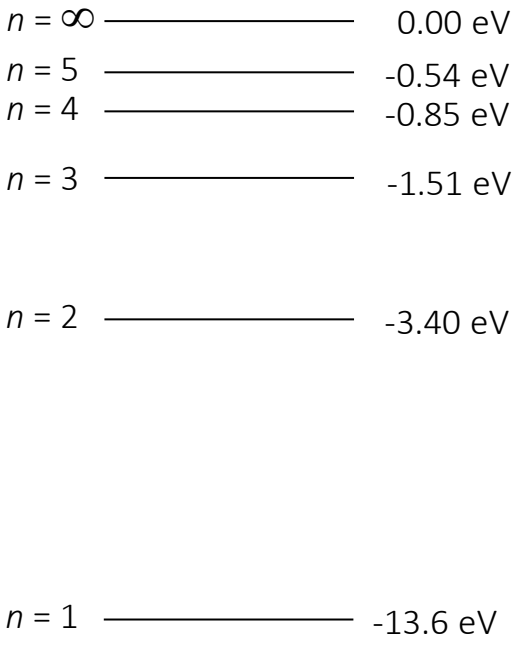
Working Backwards…



Draw in the Energy Transition for a 434 nm blue emission line?

What transition has an energy difference of 2.86 eV?