

# The Expanding Universe

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IB PHYSICS | ASTROPHYSICS

# IB Physics Data Booklet

| Sub-topic D.1 – Stellar quantities  | Sub-topic D.2 – Stellar characteristics and stellar evolution           |
|---|---|
| $d \text{ (parsec)} = \frac{1}{p \text{ (arc-second)}}$ $L = \sigma AT^4$ $b = \frac{L}{4\pi d^2}$                      | $\lambda_{\max} T = 2.9 \times 10^{-3} \text{ m K}$ $L \propto M^{3.5}$ |
| Sub-topic D.3 – Cosmology   | Sub-topic D.5 – Further cosmology (HL only)                             |
| $z = \frac{\Delta\lambda}{\lambda_0} \approx \frac{v}{c}$ $z = \frac{R}{R_0} - 1$ $v = H_0 d$ $T \approx \frac{1}{H_0}$ | $v = \sqrt{\frac{4\pi G\rho}{3}} r$ $\rho_c = \frac{3H^2}{8\pi G}$      |

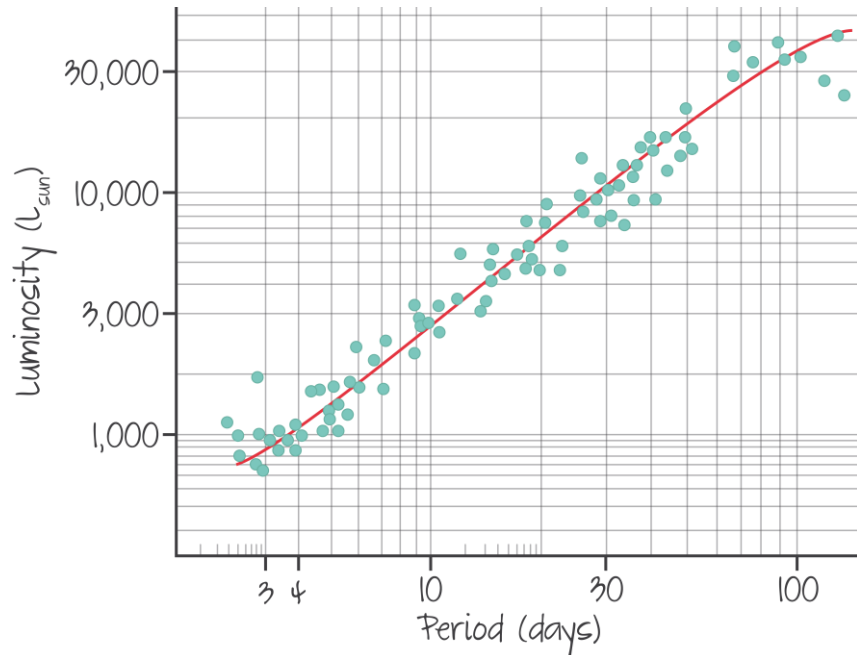
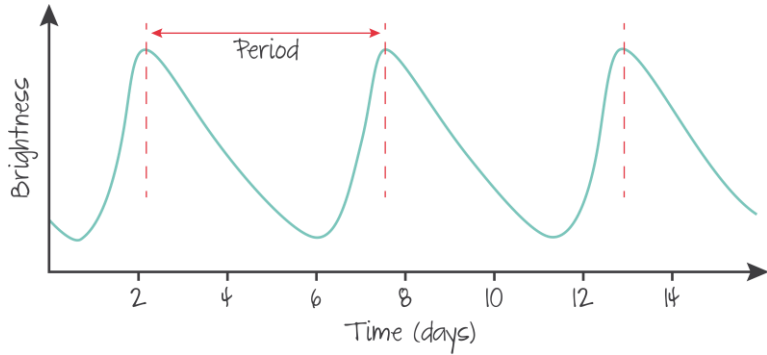
# Henrietta Swan Leavitt



# Cepheid Variables

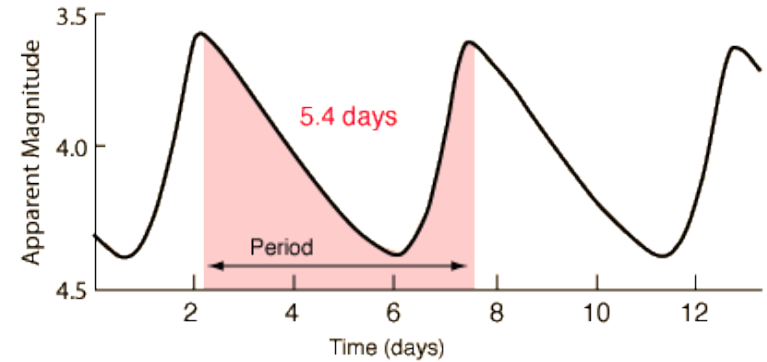
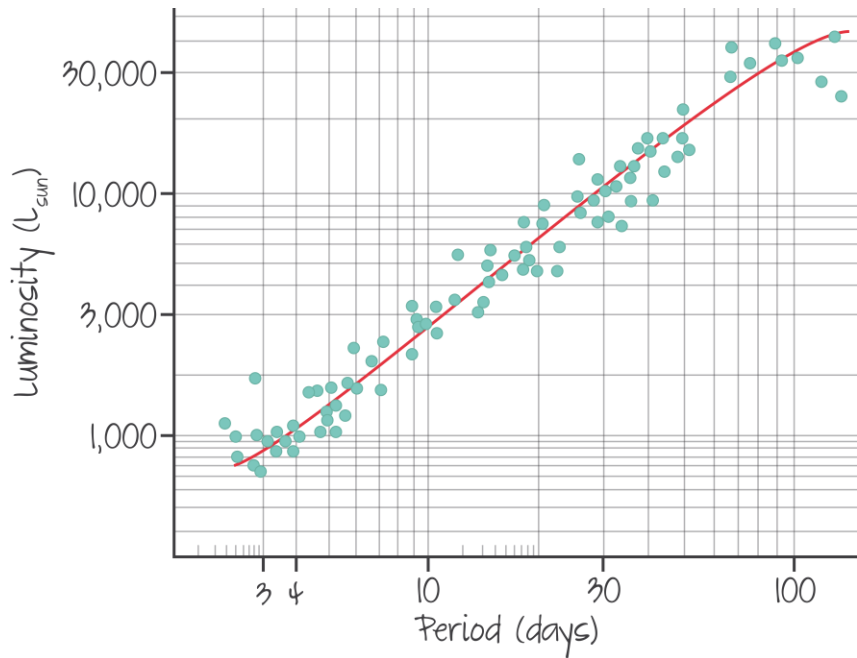


# “Standard Candle”



# Cepheid Variables

$$1 L_{\text{sun}} = 3.84 \times 10^{24} \text{ W}$$

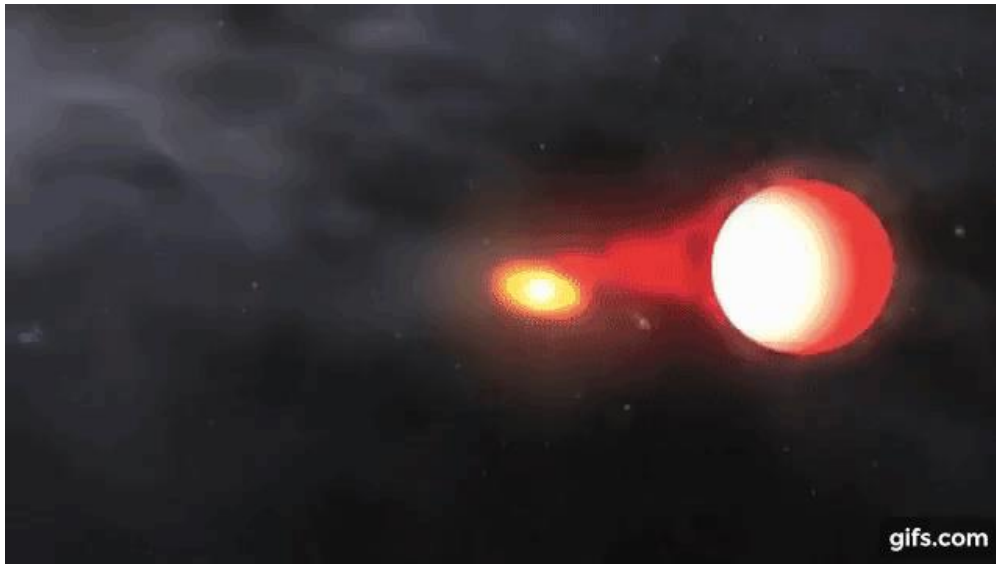


What is the distance of the Cepheid Variable with the period shown in the graph above? The brightness of this star is  $8 \times 10^{-10} \text{ W m}^{-2}$ .

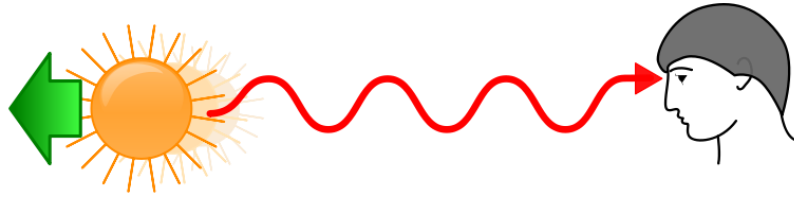
# Type Ia Supernova

A type Ia Supernova forms when a white dwarf accretes matter from a companion star until it exceeds the Chandrasekhar limit and explodes

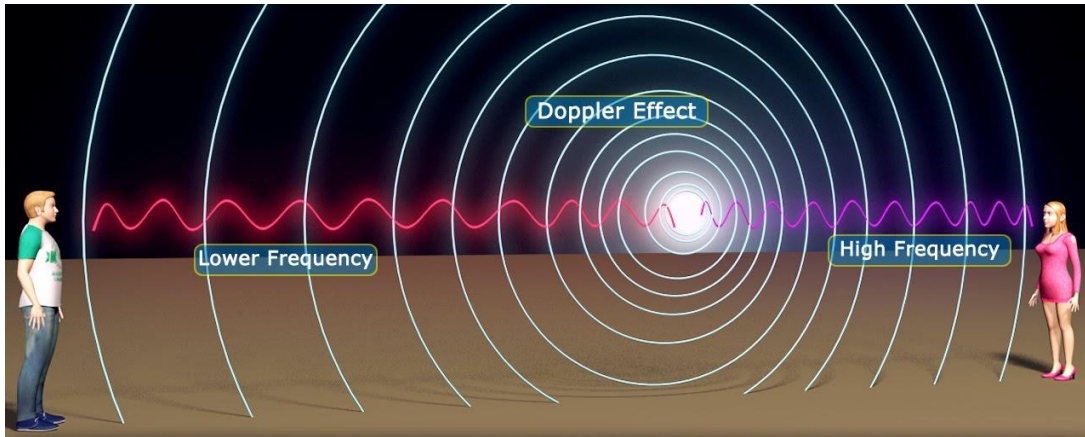
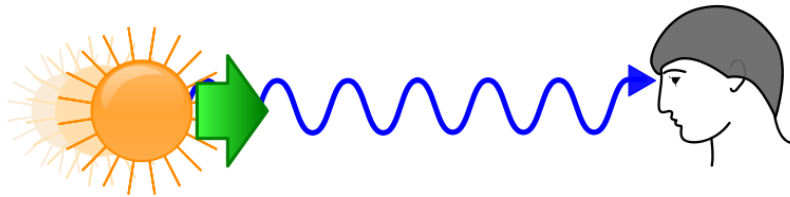
These supernovae have a constant luminosity so their brightness can be analyzed as a standard candle much like the Cepheid Variables



# Doppler Effect



Red-shifted

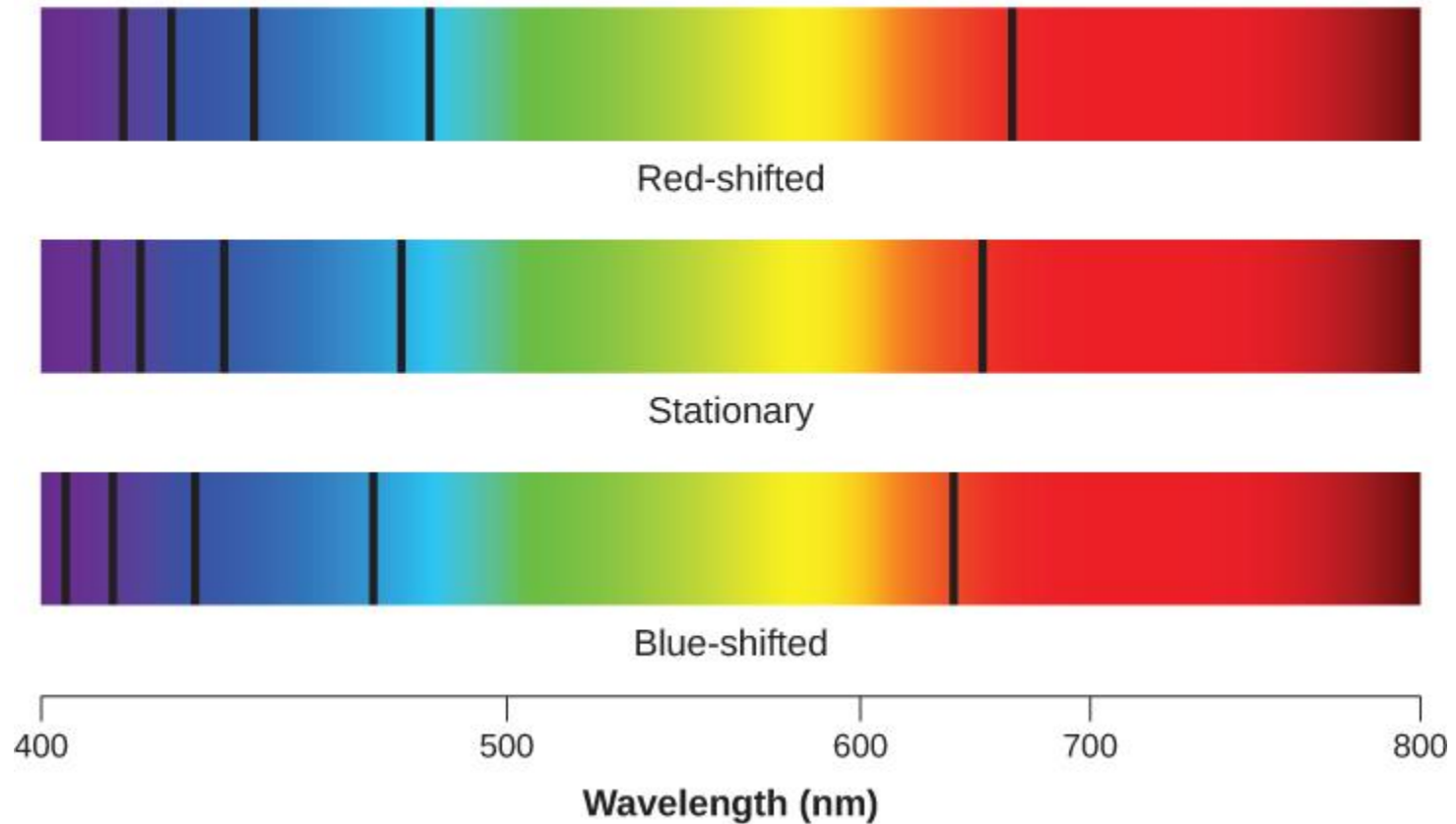


Fun Science: Light



# Red Shift, Blue Shift

## The Doppler Shift



# Calculating Redshift

$$z = \frac{\Delta\lambda}{\lambda_0} \approx \frac{v}{c}$$

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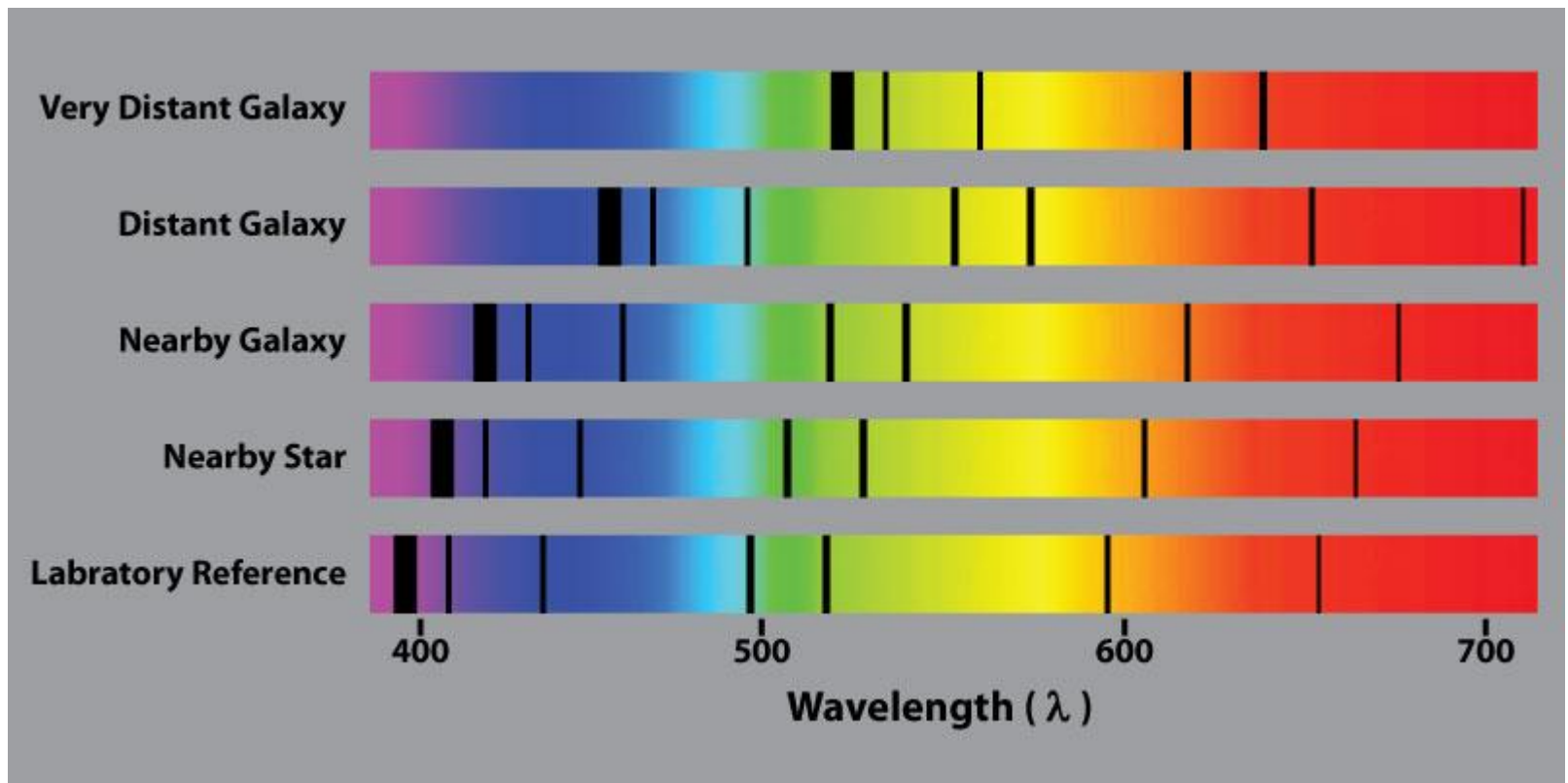
# Calculating Redshift

A characteristic absorption line often seen in stars is due to ionized helium. It occurs at 468.6 nm. If the spectrum of a star has this line at a measured wavelength of 499.3 nm what is the recession speed of the star?

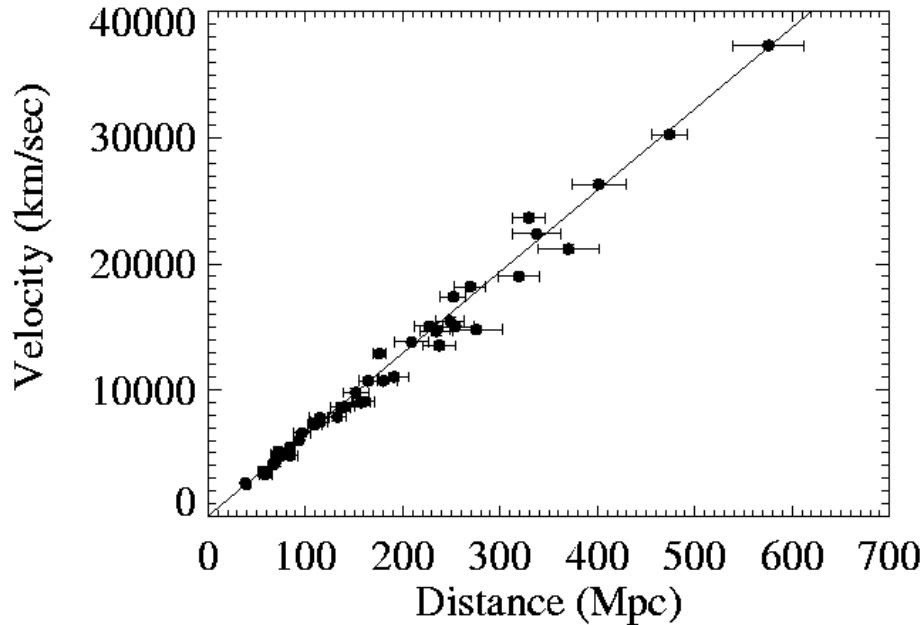
$$\frac{\Delta\lambda}{\lambda_0} \approx \frac{v}{c}$$

# Hubble's Big Discovery

Edwin Hubble discovered that the amount of redshift changed by the distance



# The Universe is Expanding



$$v = H_0 d$$

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# Using the Hubble “Constant”

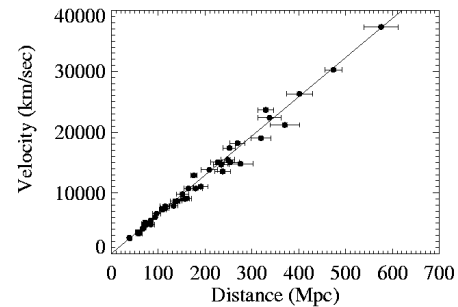
Estimate the distance from the Earth to a galaxy with a recessional velocity of  $150 \text{ km s}^{-1}$

If a galaxy is 20 Mpc from Earth, how fast will it be receding?



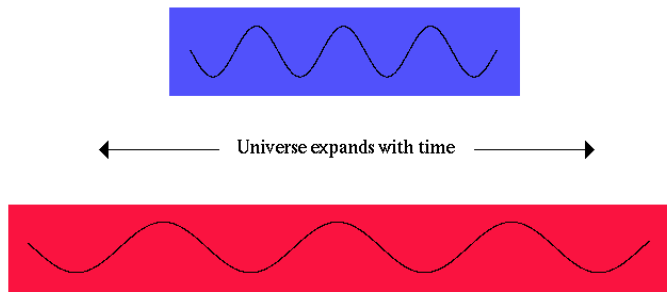
# Calculating Redshift

Nothing can go faster than the speed of light so the Doppler effect can't really hold up...



# Calculating Redshift

Think of the wavelength change due to the stretching of space-time



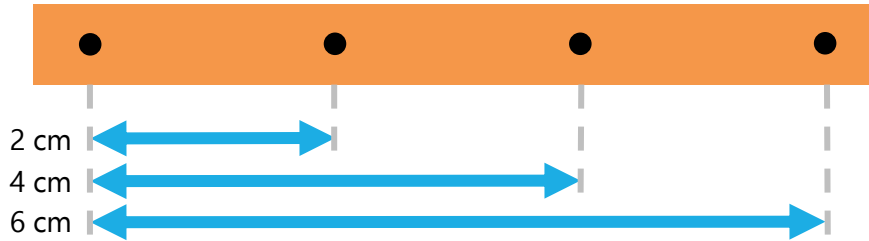
$$z = \frac{R}{R_0} - 1$$

# Calculating Redshift

If the redshift  $z = 3$ , what was the scale factor at the time that the light was emitted?

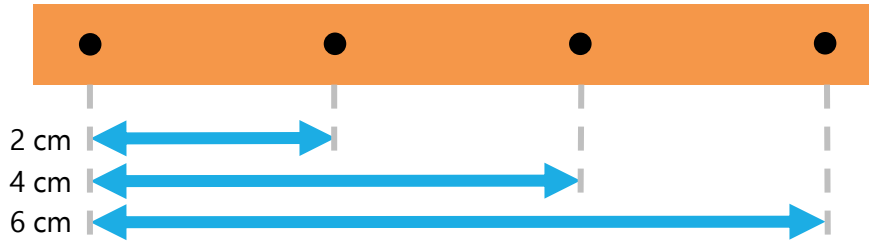
$$z = \frac{R}{R_0} - 1$$

# The Universe is Expanding



Think of a rubber band with marks when it is stretched out... Relative to the first dot, which dot moves the fastest?

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Think of a rubber band with marks when it is stretched out... Relative to the first dot, which dot moves the fastest?



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