

Radiation from the Sun

IB PHYSICS | ENERGY PRODUCTION

Intensity

$$\textit{Intensity} = \frac{\textit{Power}}{A}$$

$$\textit{Units} = \frac{W}{m^2} = W m^{-2}$$

Intensity

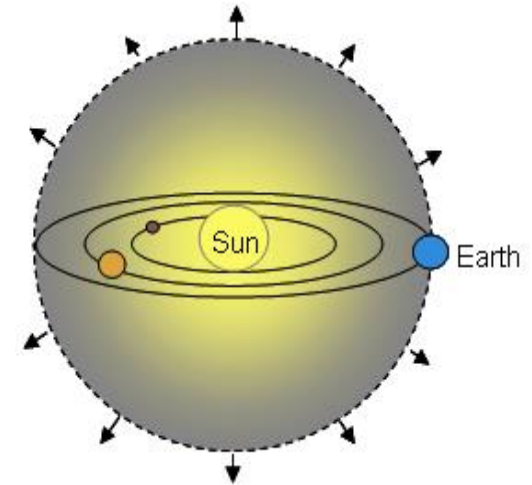
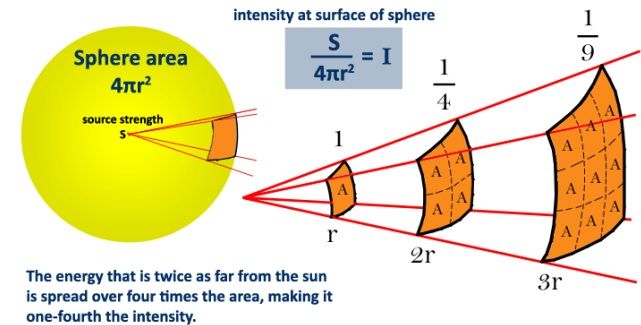
Calculate the intensity of the Sun's radiation arriving to Earth

Sun's Power = 3.84×10^{26} W

Earth's Distance from Sun = 150×10^6 km

$$I = \frac{P}{A} = \frac{3.84 \times 10^{26}}{4\pi(150 \times 10^9)^2}$$

$$I = 1358 \text{ Wm}^{-2}$$



Solar Constant

The average intensity falling on an area above the earth's atmosphere perpendicular to the direction traveled by the radiation

$$S = 1360 \text{ W m}^{-2} = 1.36 \times 10^3 \text{ W m}^{-2}$$

Quantity	Symbol	Approximate value
Elementary charge	e	$1.60 \times 10^{-19} \text{ C}$
Electron rest mass	m_e	$9.110 \times 10^{-31} \text{ kg} = 0.000549 \text{ u} = 0.511 \text{ MeV c}^{-2}$
Proton rest mass	m_p	$1.673 \times 10^{-27} \text{ kg} = 1.007276 \text{ u} = 938 \text{ MeV c}^{-2}$
Neutron rest mass	m_n	$1.675 \times 10^{-27} \text{ kg} = 1.008665 \text{ u} = 940 \text{ MeV c}^{-2}$
Unified atomic mass unit	u	$1.661 \times 10^{-27} \text{ kg} = 931.5 \text{ MeV c}^{-2}$
Solar constant	S	$1.36 \times 10^3 \text{ W m}^{-2}$
Fermi radius	R_0	$1.20 \times 10^{-15} \text{ m}$

Average Solar Intensity on Earth

Earth's Radius = 6.37×10^6 m

Area of sun power captured:

$$\pi r^2 = \pi(6.37 \times 10^6)^2 =$$

$$1.27 \times 10^{14} \text{ m}^2$$

1360 W m^{-2}



Total sun power captured:

$$1.27 \times 10^{14} \text{ m}^2 \times \frac{1360 \text{ W}}{1 \text{ m}^2} =$$

$$1.7 \times 10^{17} \text{ W}$$

Total Power Received by the Earth

$$1.7 \times 10^{17} \text{ W}$$

Average spread out across Earth's surface:

$$\frac{P}{A_{sphere}} = \frac{1.7 \times 10^{17} \text{ W}}{4\pi(6.37 \times 10^6)^2} =$$

$$340 \text{ W m}^2$$

Average Solar Intensity on Earth

$$340 \text{ W m}^2$$

Albedo vs. Emissivity

Albedo

$$\frac{\text{power scattered by a body}}{\text{incident power}}$$

% Reflected

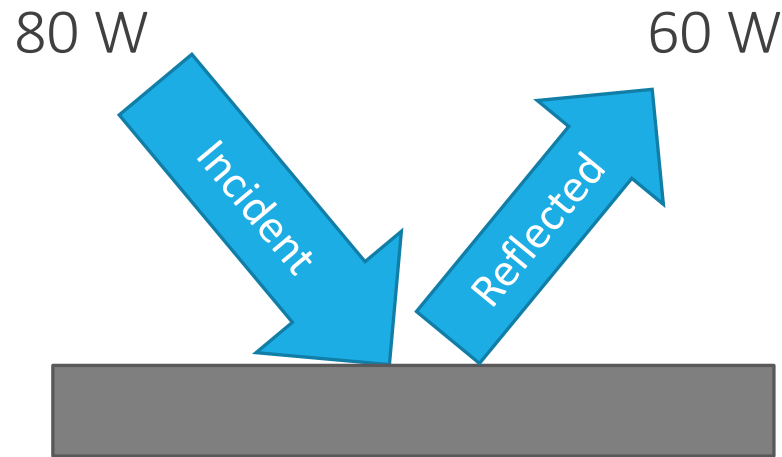
Emissivity

e

$$\frac{\text{power radiated by a surface}}{\text{power radiated from a black body}}$$

% Absorbed

Albedo vs. Emissivity



Albedo

$$\frac{60}{80} = 0.75$$

Emissivity

$$\frac{20}{80} = 0.25$$

Albedo of Earth



$$\textit{Albedo} = \frac{102}{340} = 0.3$$

Albedo of Earth

Highest Albedo?

0.66

Snow

Lowest Albedo?

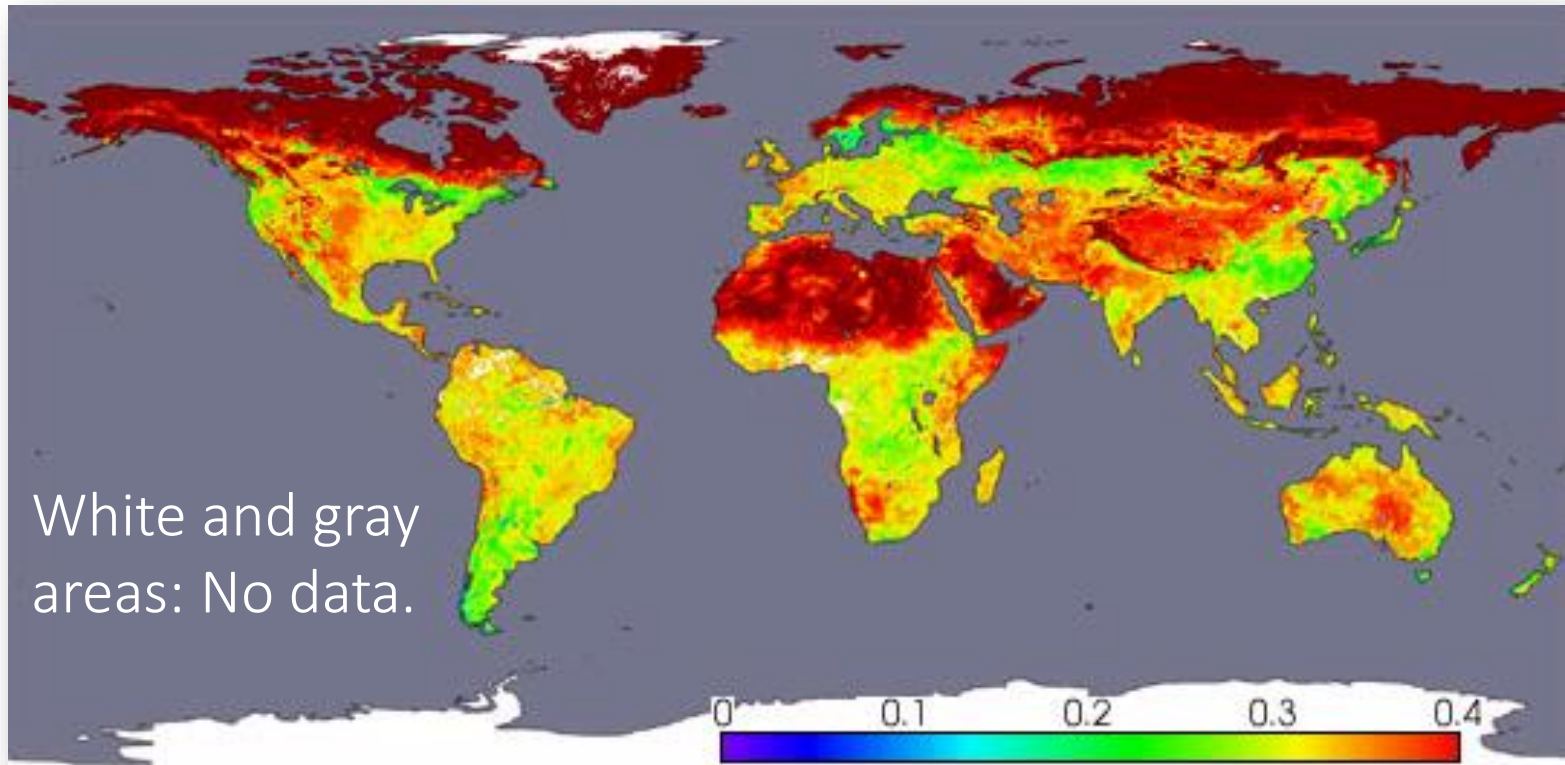
0.07

Ocean

Surface	Albedo ^a	
	Summer	Winter
Mixed farming, tall grass	0.16	0.18
Tall/medium grassland, evergreen shrubland	0.20	0.21
Short, grassland, meadow and shrubland	0.21	0.20
Evergreen forest (needle leaved)	0.12	0.13
Mixed deciduous, evergreen forest	0.16	0.16
Deciduous forest	0.17	0.18
Tropical evergreen broadleaved forest	0.12	0.15
Medium/tall grassland, woodland	0.15	0.18
Desert	0.36	0.36
Tundra	0.17	0.17
Snow	0.66	0.66
Sea ice	0.62	0.62
Ocean	0.07	0.07

Data taken from Briegleb *et al.* (1986).

Albedo of Earth



April, 2002, *Terra* satellite, NASA

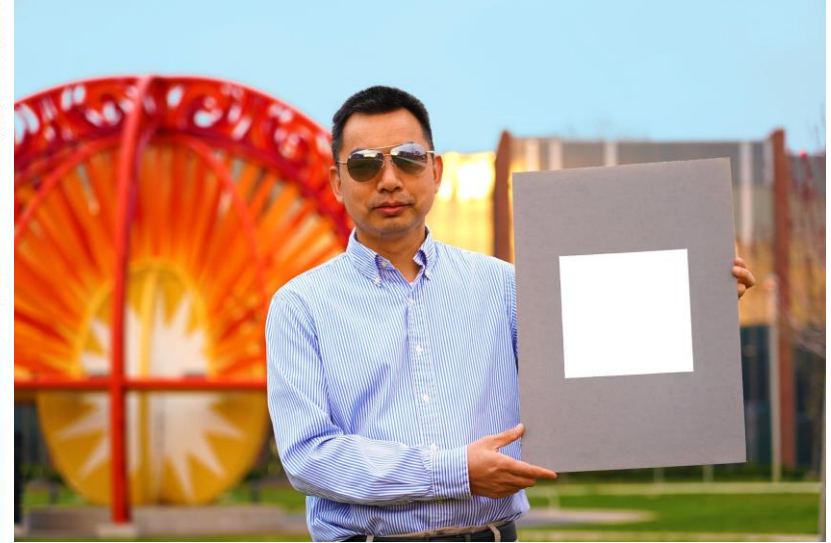
Adjusting our Albedo



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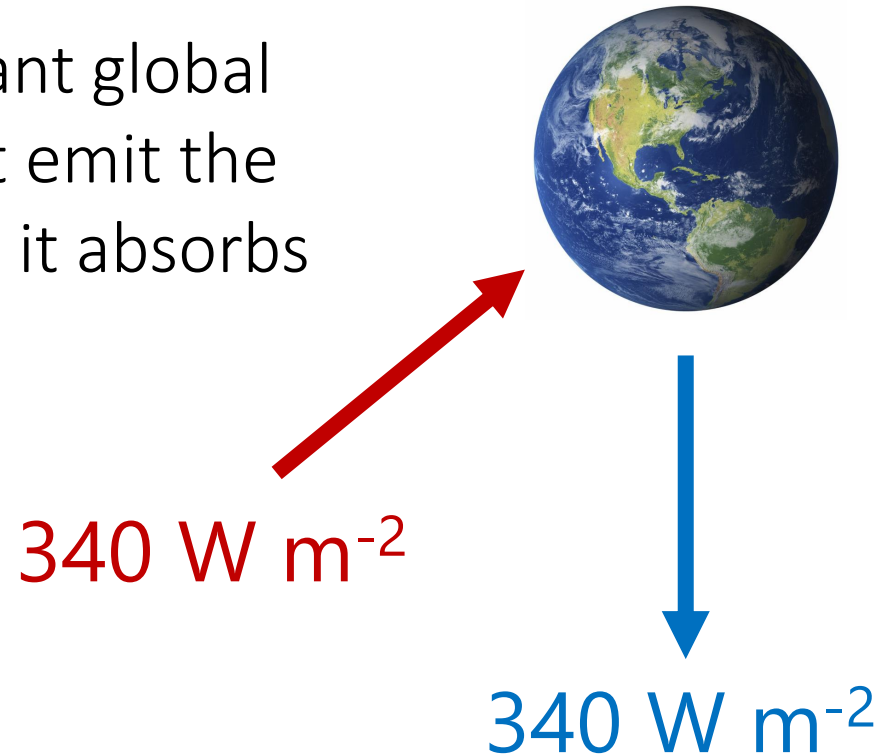
Los Angeles paints streets white to stay cool

By David Shultz | Sep. 7, 2017, 5:00 PM



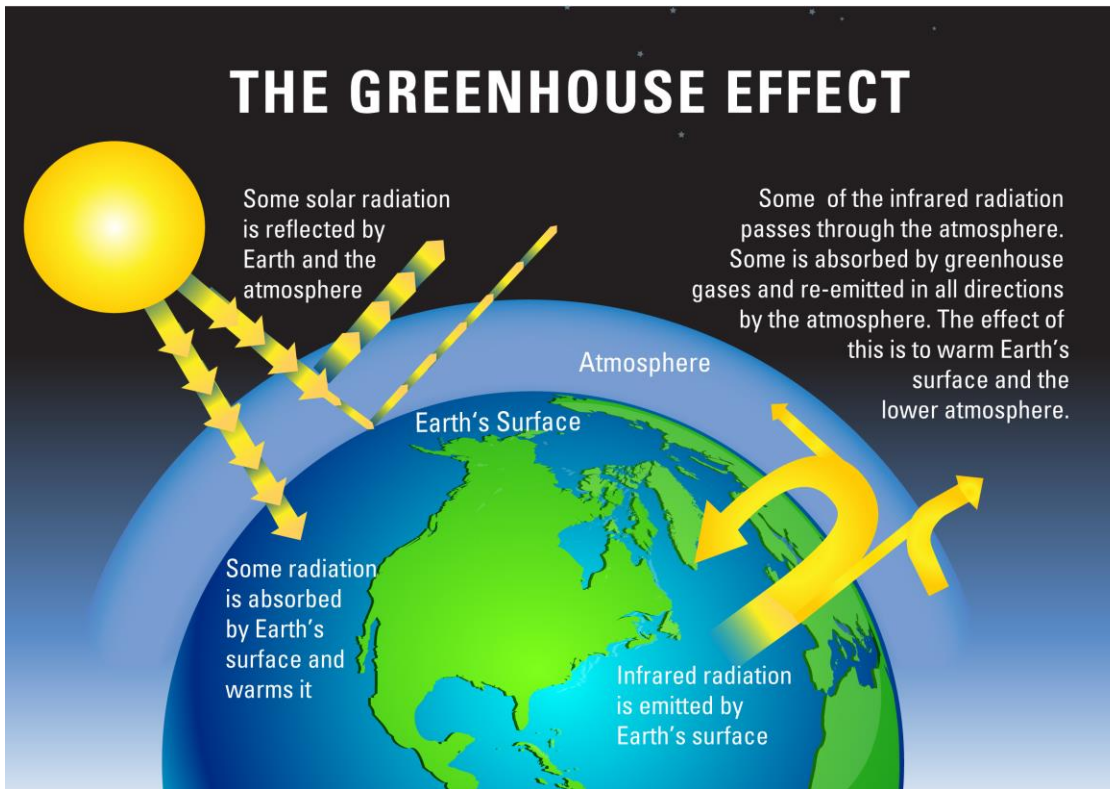
Thermal Equilibrium

In order to maintain a constant global temperature, the Earth must emit the same amount of energy that it absorbs

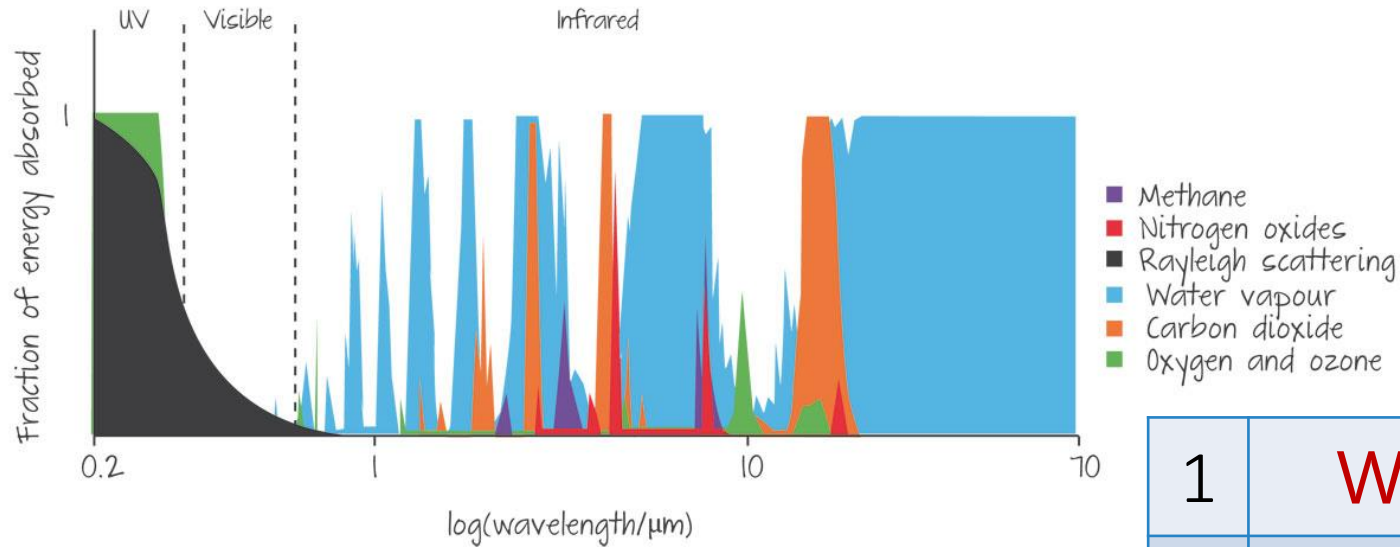


Greenhouse Effect

If there was no atmosphere, the earth would experience a net loss of energy and reach equilibrium at an average temperature about 30°C colder than it is currently.



Role of the Atmosphere



Rank the following Greenhouse Gases based on the amount of infrared energy they absorb



1	Water Vapor
2	Carbon Dioxide
3	Methane
4	Nitrogen Oxides
5	Oxygen/Ozone

More on this later...