

# ENERGY PRODUCTION

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IB PHYSICS | COMPLETED NOTES

# Energy Sources Overview

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IB PHYSICS | ENERGY PRODUCTION

# What is Energy Used For?

## Residential/Commercial



## Industrial



## Electric Power



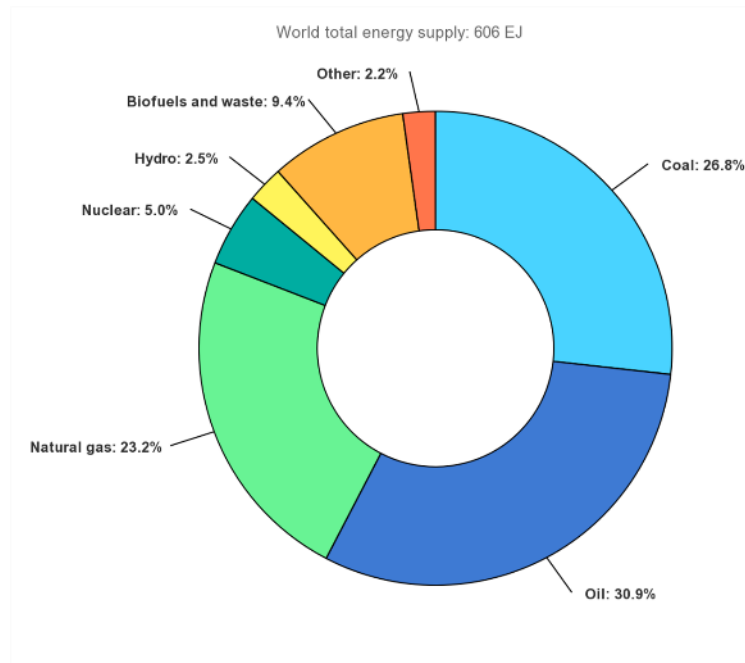
## Transportation



# Where does our Energy Come From?

## World Sources

### 2019 Global Energy Supply

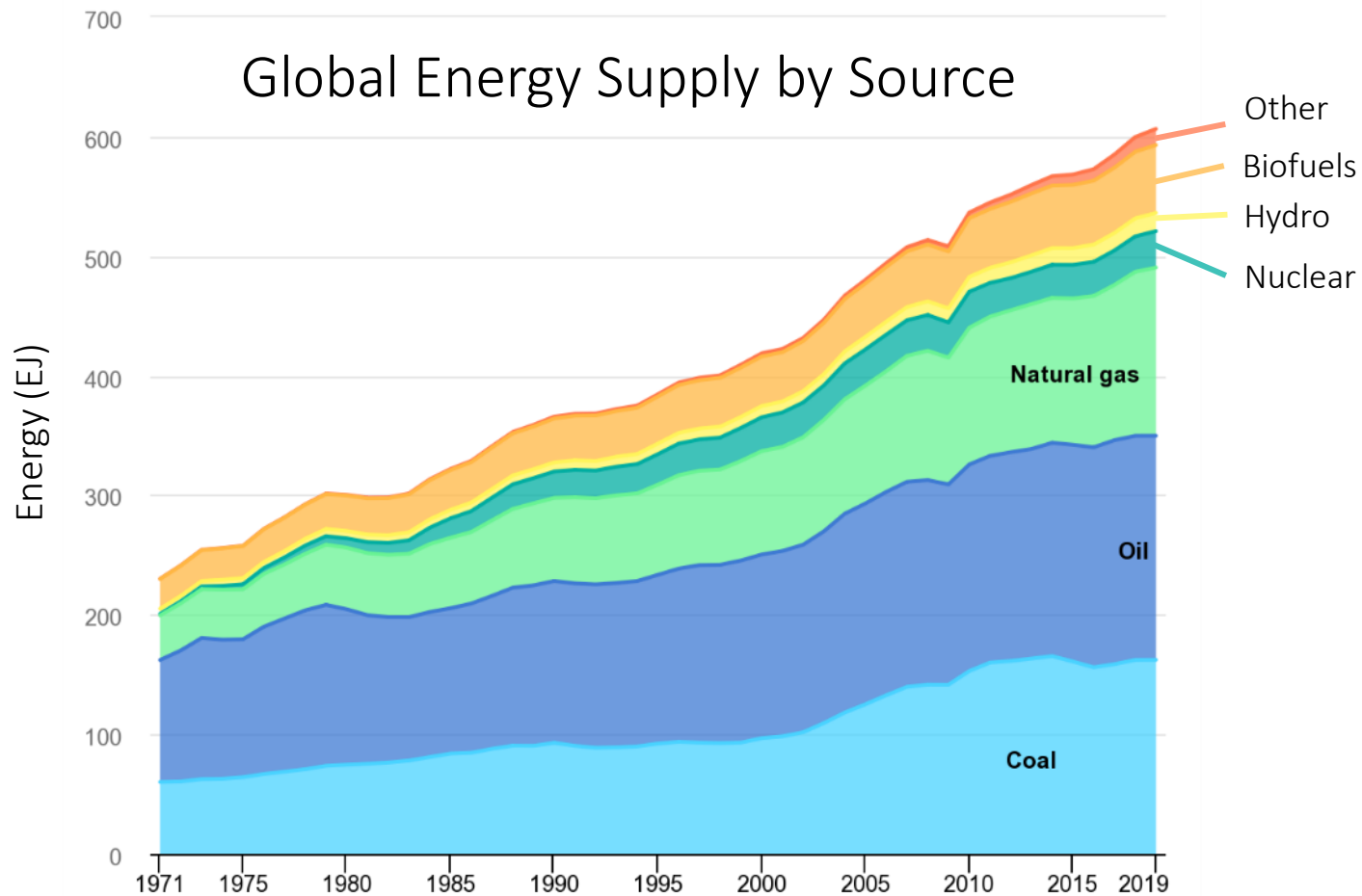


## List Sources in Order:

<b>Oil</b>	~31%
<b>Coal</b>	~27%
<b>Natural Gas</b>	~23%
<b>Biofuels</b>	~9%
<b>Nuclear</b>	~5%
<b>Hydropower</b>	~2.5%

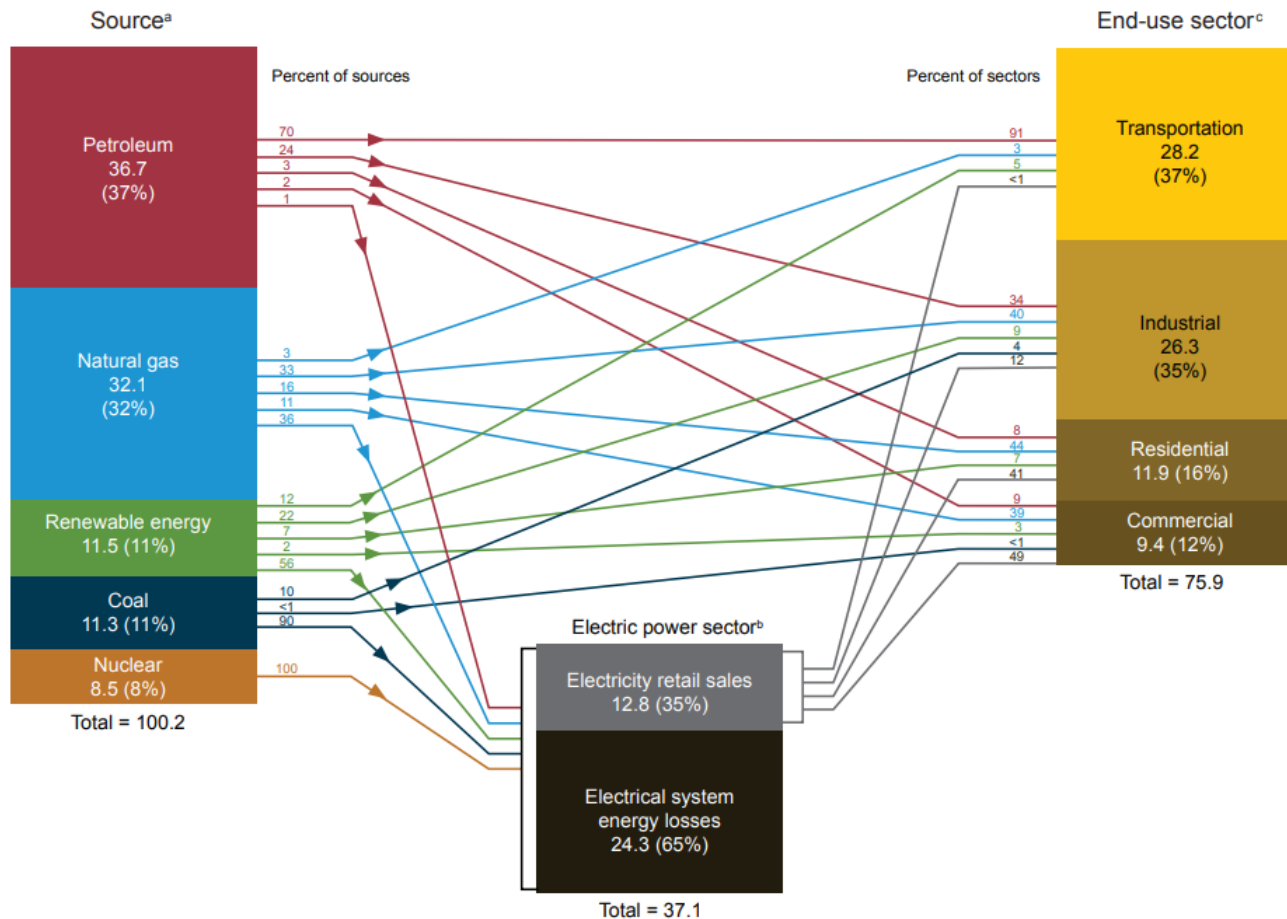


# This changes over time



# Used in Many Ways...

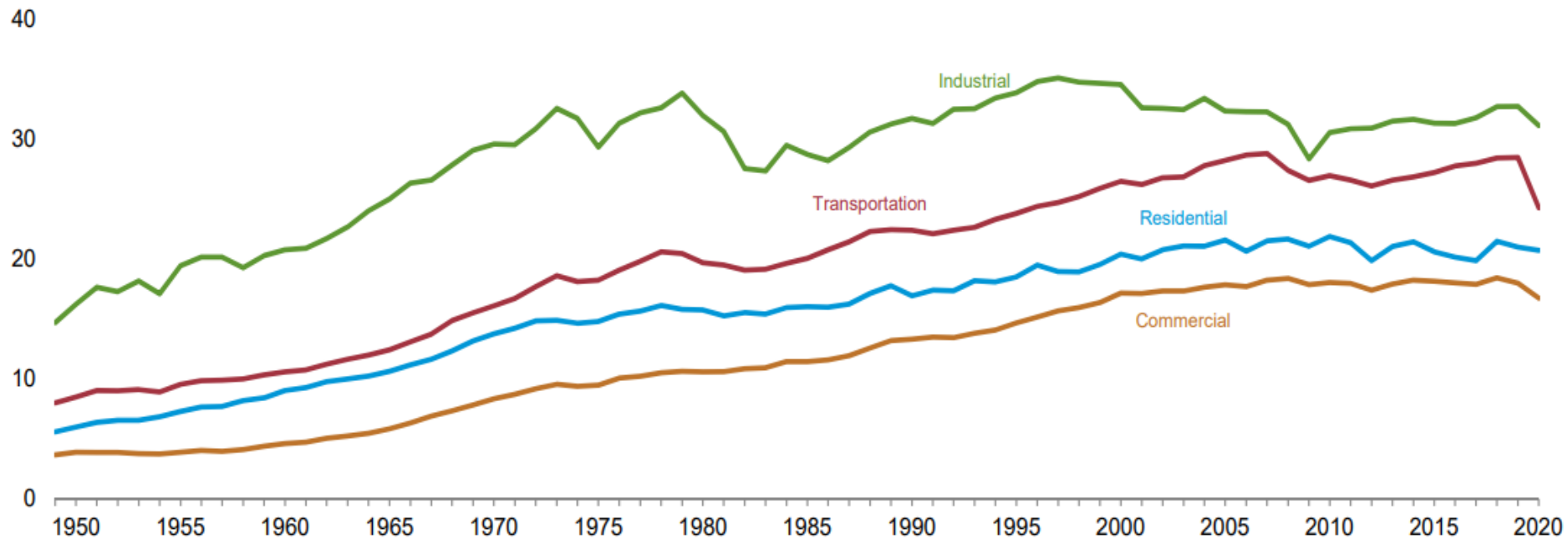
**U.S. energy consumption by source and sector, 2019**  
(Quadrillion Btu)



# Used in Many Ways...

**Figure 2.1 Energy Consumption by Sector**  
(Quadrillion Btu)

Total Consumption by End-Use Sector, 1949–2020



# Primary vs. Secondary Sources

**Primary energy sources** are sources found in the natural environment



**Petroleum**



**Natural Gas**



**Coal**



**Propane**



**Uranium**



**Solar**



**Geothermal**



**Wind**



**Hydropower**



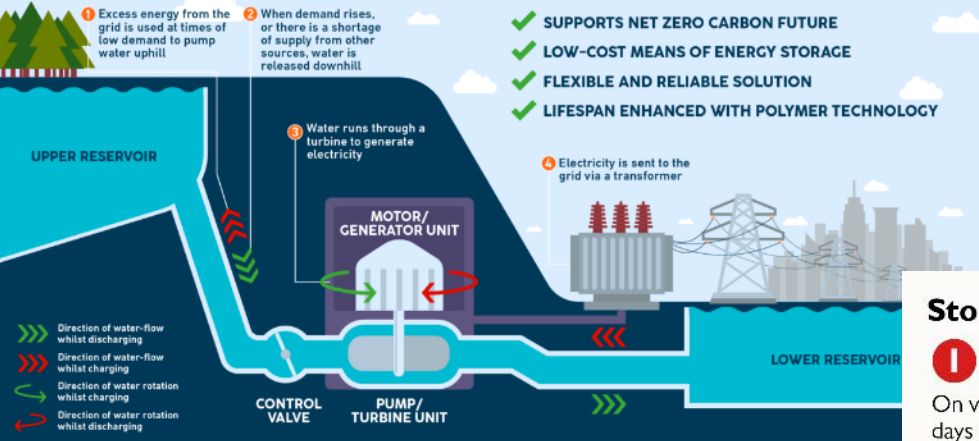
**Biomass**

**Secondary energy sources** are useful transformations of the primary sources. (typically used to **store** energy)

**Electricity – Batteries, Stored Hydropower**

# Other Secondary Sources...

## PUMPED HYDROPOWER STORAGE GUIDE



## Storing excess energy

**1**

On very windy days excess electricity produced by turbines would be used to pull the weight to the top of the shaft



**2**

On still days when electricity is required the weight is lowered generating energy which is then passed back to the grid

Each unit can be configured to produce between 1 and 20MW peak power, with output duration from 15 minutes to 8 hours

Electrical power is absorbed or generated by raising or lowering the weight. The winch system can be accurately controlled through the electrical drives to keep the weight stable in the hole

What is the Ultimate Primary Source?

**The Sun**

# Which one is the best?

That's a hard question and depends on many factors

Some of the big ones are:

- Energy Density
- Cost
- Availability and Location
- Politics
- Safety
- Environmental

# U.S. Electricity Generation

**Figure 2.6 Electric Power Sector Energy Consumption**  
(Quadrillion Btu)

By Major Source, 1949–2020

24

20

16

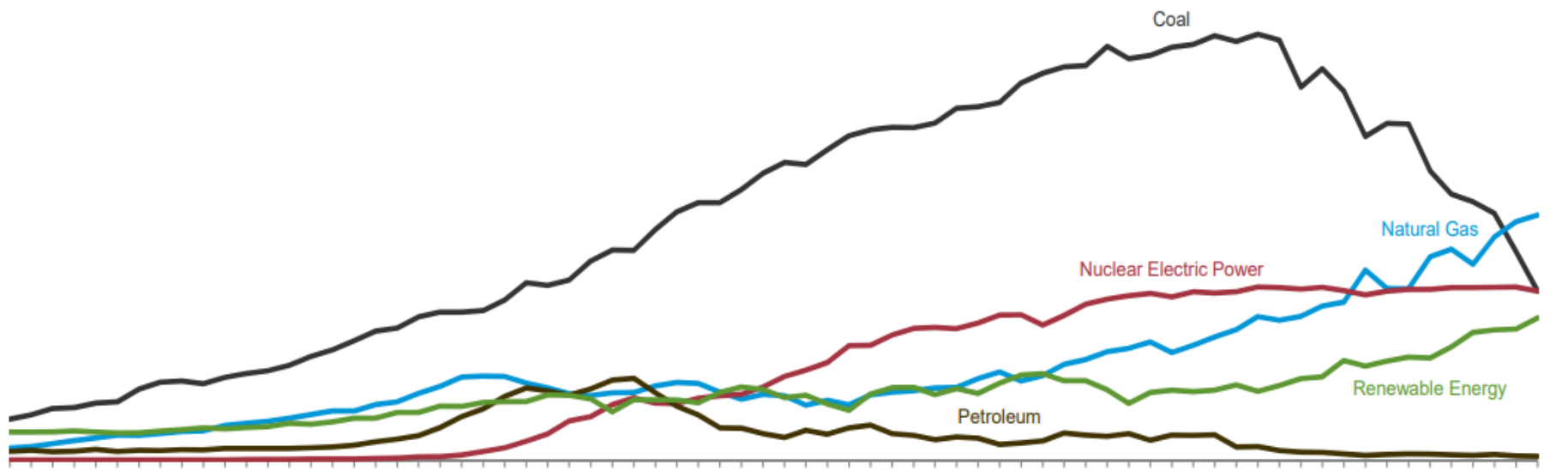
12

8

4

0

1950 1955 1960 1965 1970 1975 1980 1985 1990 1995 2000 2005 2010 2015 2020



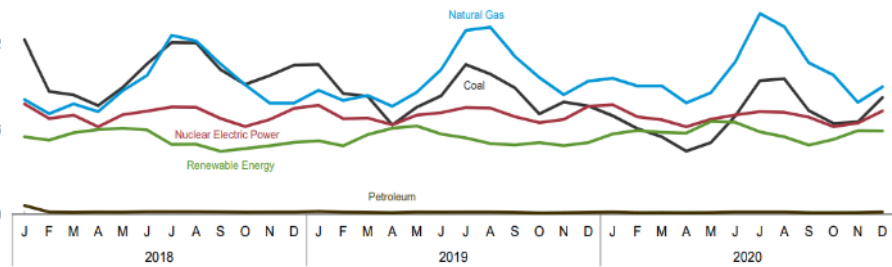
By Major Source, Monthly

1.8

1.2

0.6

0.0





# Renewable vs. Non Renewable

Highlight the primary energy sources that are considered **renewable**



**Petroleum**



**Geothermal**



**Coal**



**Hydropower**



**Uranium**



**Solar**



**Natural Gas**



**Wind**



**Propane**

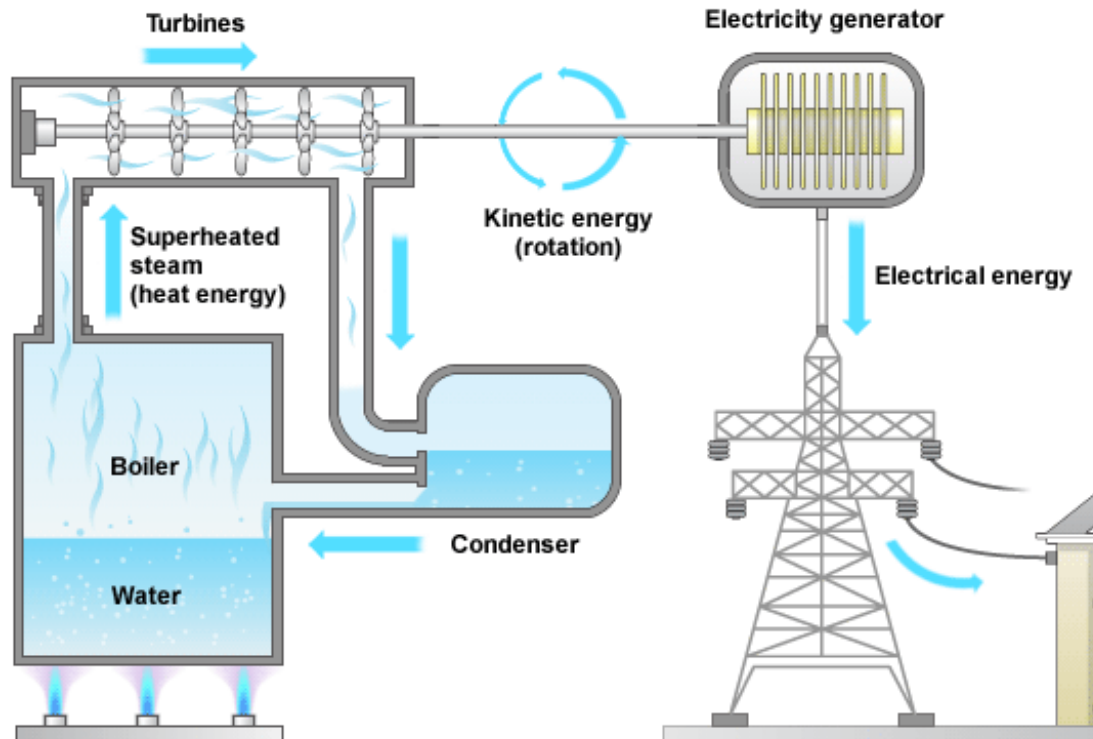


**Biomass**

\*Note: this doesn't mean that it cannot ever be replaced, just that it won't happen in any sort of useful time frame...

# Efficiency

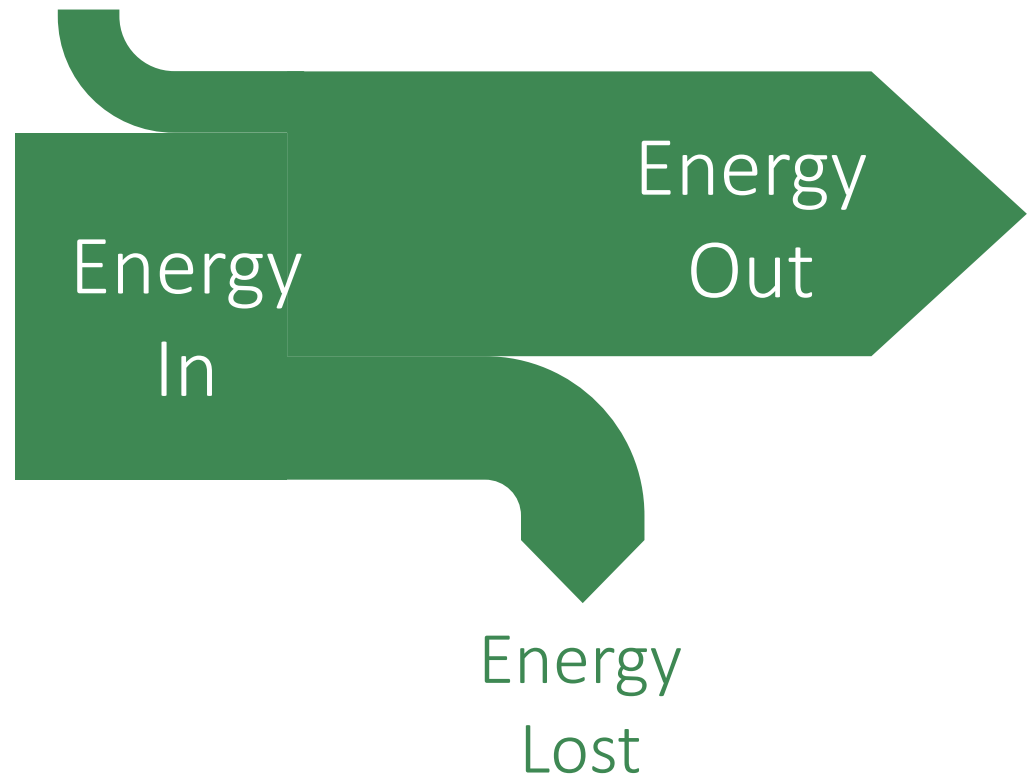
Almost every energy source has the same general path



# Sankey Diagram

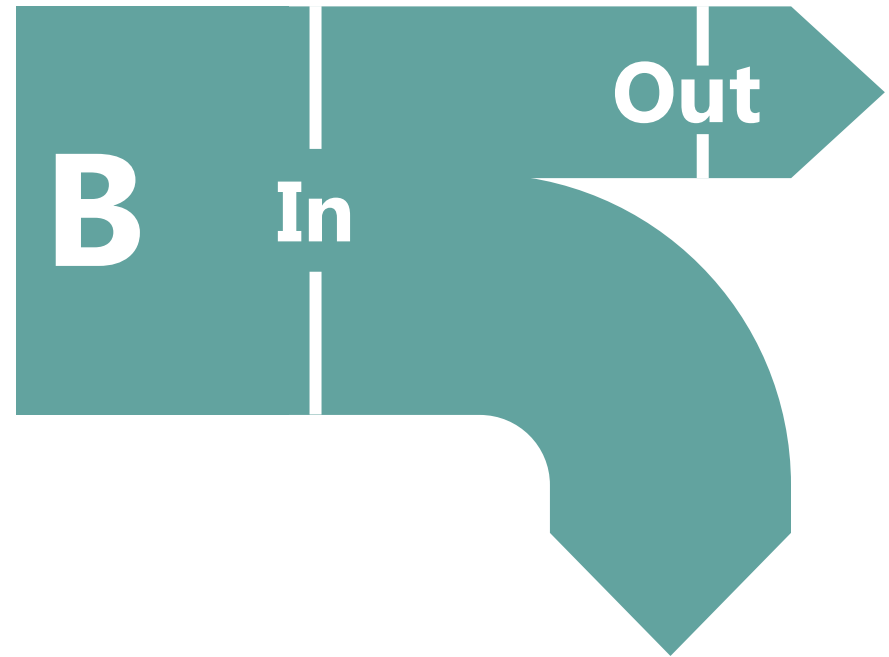
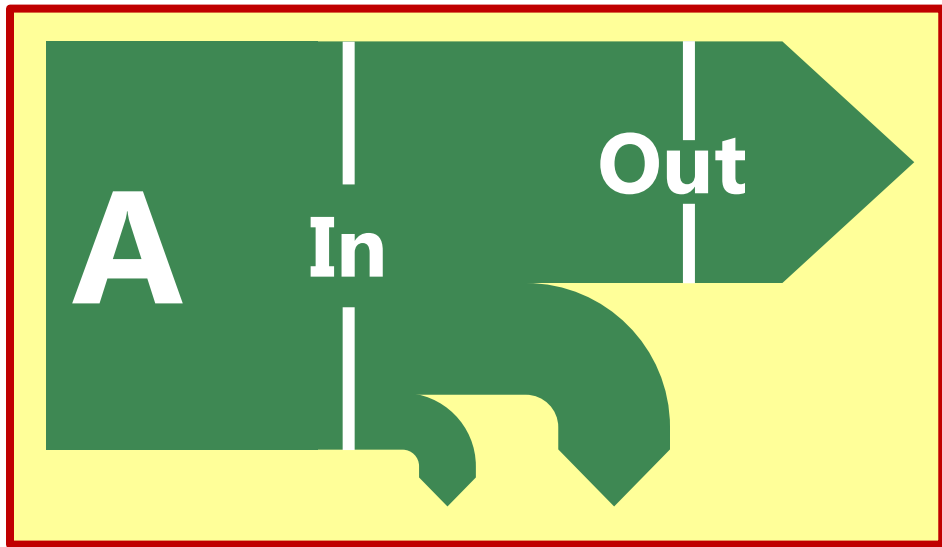
Sometimes it's easiest to represent energy flow in a picture.

In these diagrams the width of the arrow represents the amount of energy

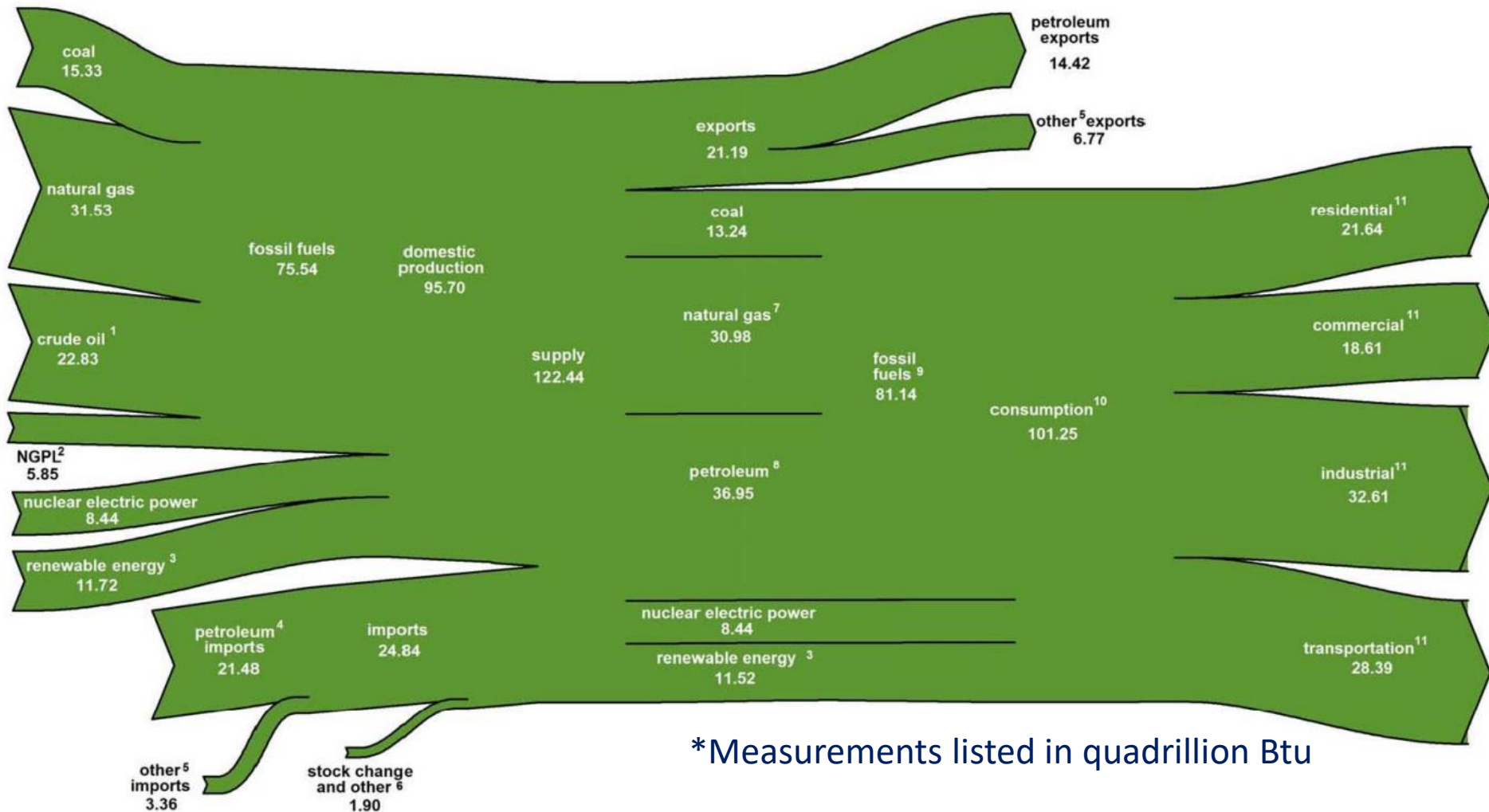


# Sankey Diagrams

Which process is more efficient?



# U.S. Energy Flow 2018



\*Measurements listed in quadrillion Btu

# Cost

## Levelized Cost of Energy Comparison—Unsubsidized Analysis

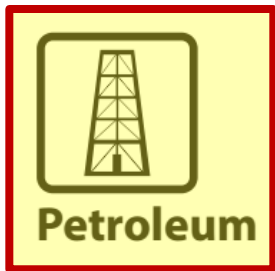
Selected renewable energy generation technologies are cost-competitive with conventional generation technologies under certain circumstances



Source: Lazard estimates.

# CO<sub>2</sub> Emissions

Highlight the primary energy sources that are produce Carbon Dioxide



**Geothermal**



**Hydropower**



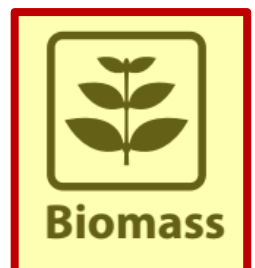
**Uranium**



**Solar**



**Wind**



\*Note: this is just one of several greenhouse gases. We'll discuss this.

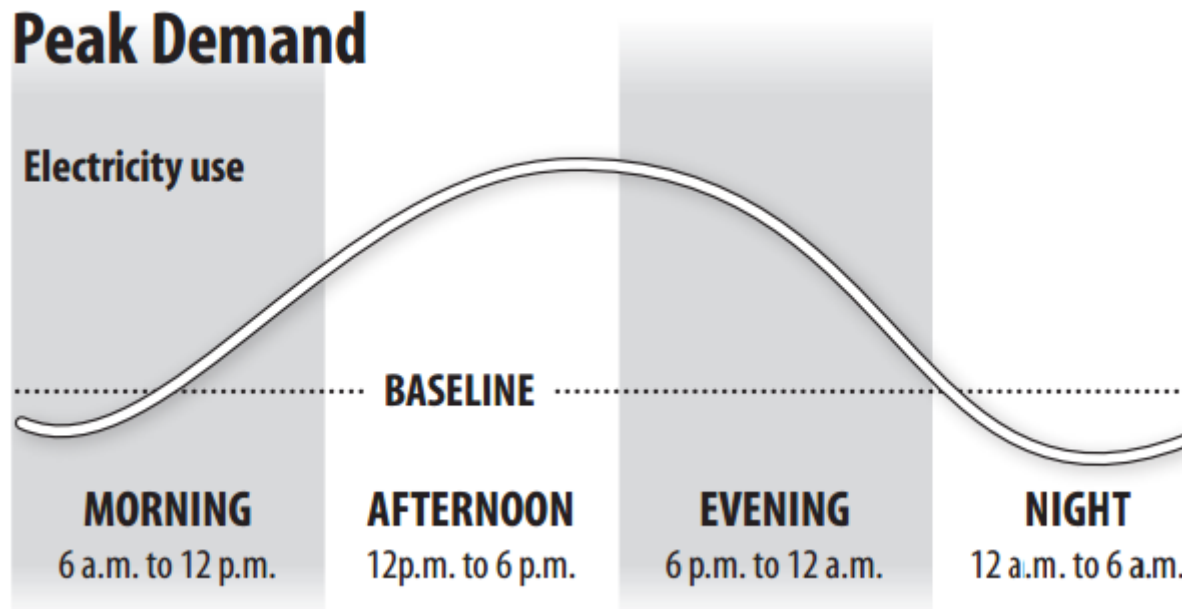
# Location Dependency and Politics





# Energy Load Requirements

Energy needs to be available when electricity is most needed but should also be available other times as well.



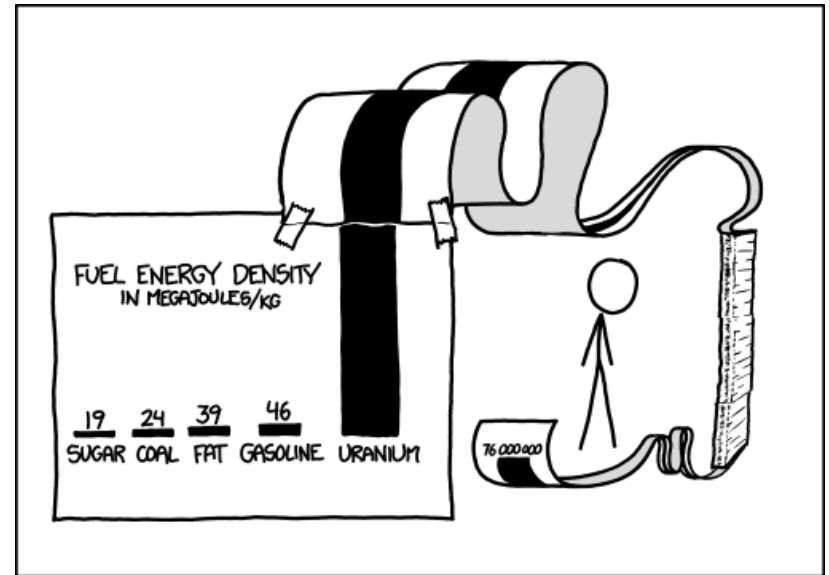
# Energy Density

Specific Energy = Energy per Unit Mass [ $\text{J kg}^{-1}$ ]

Energy Density = Energy per Unit Volume [ $\text{J m}^{-3}$ ]

$$E_S = \frac{E}{m}$$

$$E_D = \frac{E}{V}$$



SCIENCE TIP: LOG SCALES ARE FOR QUITTERS WHO CAN'T FIND ENOUGH PAPER TO MAKE THEIR POINT PROPERLY.

# Energy Density

Material	Specific Energy / MJ kg <sup>-1</sup>	Energy Density / MJ m <sup>-3</sup>
Uranium – (Nuclear Fission)	83,000,000	15,000,000,000,000
Natural Gas (Methane)	54	37
Gasoline/Petrol	46	34,000
Crude Oil	42	36,500
Coal	32	23,000
Ethanol	30	21,000
Wood	17	Varies
Average Food	17	Varies

Why are **Specific Energy** and **Energy Density** important?

**Higher energy density lowers transportation and storage costs**

# Power and Energy

**Energy**

Joules [J]

**Power**

Watts [W]

How are these quantities related?

$$1 \text{ W} = 1 \frac{\text{J}}{\text{s}}$$

# Energy Density

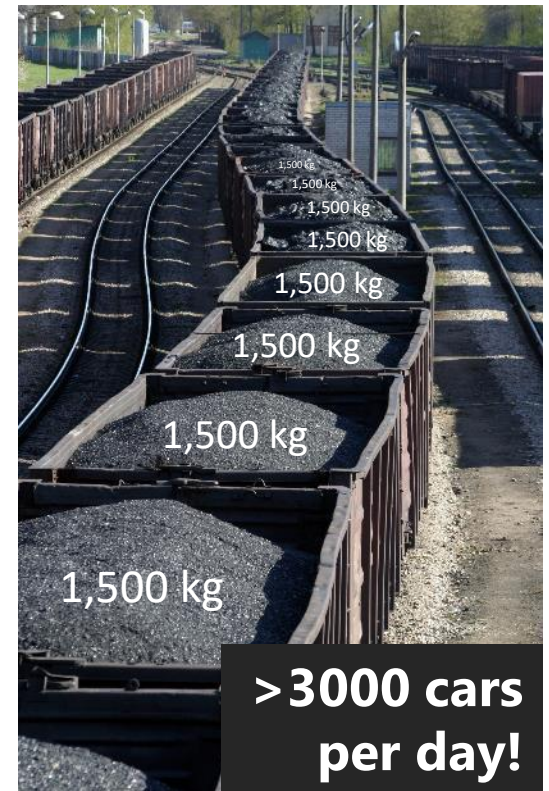
How much coal must be supplied per day to run a 500 MW power plant at 30% efficiency? (Specific Energy of coal is 32 MJ kg<sup>-1</sup>)

$$E_{in} \times 0.3 = 500 \frac{\text{MJ}}{\text{s}}$$

$$E_{in} = 1,700 \frac{\text{MJ}}{\text{s}}$$

$$\frac{1,700 \text{ MJ}}{1 \text{ s}} \times \frac{1 \text{ kg}}{32 \text{ MJ}} = 53.1 \frac{\text{kg}}{\text{s}}$$

$$\sim 4,600,000 \frac{\text{kg}}{\text{day}}$$



How many train cars per day?

# Energy Density

If a nuclear power plant powered by uranium-235 ( $83,000,000 \text{ MJ kg}^{-1}$ ) has the same output (500 MW) and the same efficiency (30%) as the coal-fired plant of the previous example, how many kg of nuclear fuel will it burn per day? Per year?

30% efficient

$$E_{in} \times 0.3 = 500 \frac{\text{MJ}}{\text{s}}$$

$$E_{in} = 1,700 \frac{\text{MJ}}{\text{s}}$$

$$\frac{1,700 \text{ MJ}}{1 \text{ s}} \times \frac{1 \text{ kg}}{83,000,000 \text{ MJ}} = 0.00002 \frac{\text{kg}}{\text{s}}$$

$$\sim 646 \frac{\text{kg}}{\text{year}}$$

$$\sim 1.77 \frac{\text{kg}}{\text{day}}$$



# Nuclear Power

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# Remember Binding Energy per Nucleon?

Nuclide	# of p	# of n	Nucleus Mass
Iodine- <b>127</b>	<b>53</b>	<b>74</b>	<b>126.87544u</b>

$$53 \times 1.007276 \text{ u}$$

$$74 \times 1.008665 \text{ u}$$

Mass Defect



$$128.026838 \text{ u} - 126.87544 \text{ u} = 1.15140 \text{ u}$$

$m_e$	0.000549u
$m_p$	1.007276u
$m_n$	1.008665u
1u	931.5 MeV $c^{-2}$

$$1.15140 \text{ u} \times \frac{931.5 \text{ MeV } c^{-2}}{1 \text{ u}} = 1072.53 \text{ MeV } c^{-2}$$

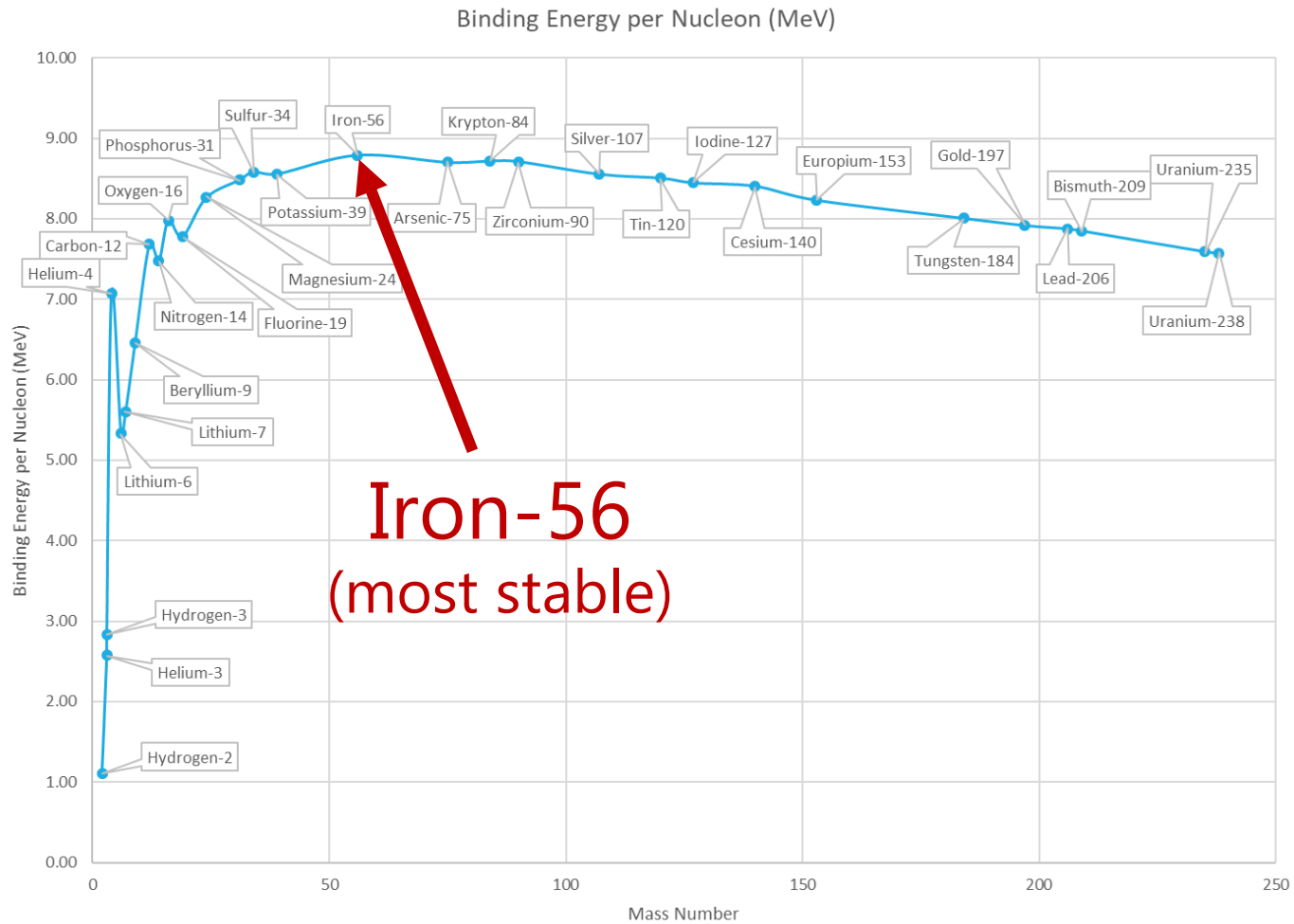
Convert mass  
to MeV  $c^{-2}$

$$E = mc^2 = (1072.53 \text{ MeV } \cancel{c^{-2}}) \cancel{c^2} = 1072.53 \text{ MeV}$$

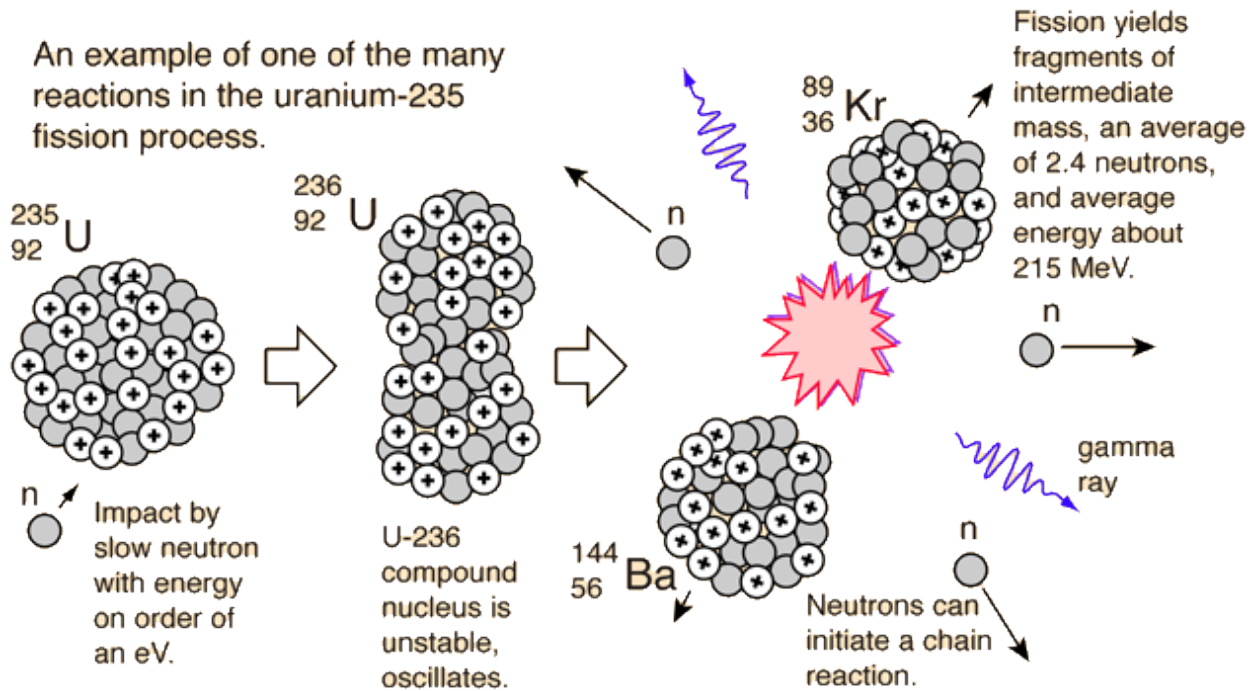
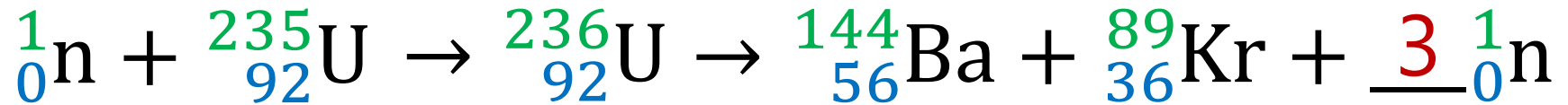
$$1072.53 \text{ MeV} / 127 = \boxed{8.45 \text{ MeV per Nucleon}}$$



# Binding Energy per Nucleon



# Fission

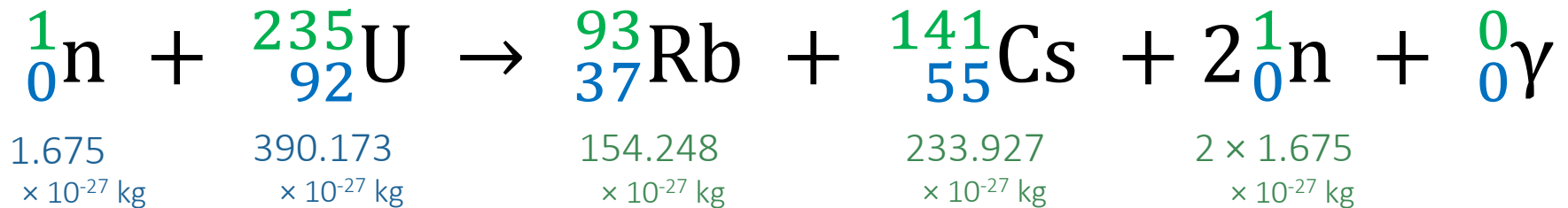


# IB Physics Data Booklet

## Fundamental constants

Quantity	Symbol	Approximate value
Speed of light in vacuum	$c$	$3.00 \times 10^8 \text{ m s}^{-1}$
Planck's constant	$h$	$6.63 \times 10^{-34} \text{ J s}$
Elementary charge	$e$	$1.60 \times 10^{-19} \text{ C}$
Electron rest mass	$m_e$	$9.110 \times 10^{-31} \text{ kg} = 0.000549 \text{ u} = 0.511 \text{ MeV c}^{-2}$
Proton rest mass	$m_p$	$1.673 \times 10^{-27} \text{ kg} = 1.007276 \text{ u} = 938 \text{ MeV c}^{-2}$
Neutron rest mass	$m_n$	$1.675 \times 10^{-27} \text{ kg} = 1.008665 \text{ u} = 940 \text{ MeV c}^{-2}$
Unified atomic mass unit	$u$	$1.661 \times 10^{-27} \text{ kg} = 931.5 \text{ MeV c}^{-2}$
Solar constant	$S$	$1.36 \times 10^3 \text{ W m}^{-2}$
Fermi radius	$R_0$	$1.20 \times 10^{-15} \text{ m}$

# Fission



$$391.848 \times 10^{-27} \text{ kg} > 391.525 \times 10^{-27} \text{ kg}$$

$$391.848 \times 10^{-27} \text{ kg} - 391.525 \times 10^{-27} \text{ kg} = 0.323 \times 10^{-27} \text{ kg}$$

$$0.323 \times 10^{-27} \text{ kg} \times \frac{1 \text{ u}}{1.661 \times 10^{-27} \text{ kg}} = \mathbf{0.19446 \text{ u}}$$

$$\mathbf{0.19446 \text{ u} \times 931.5 = 181.14 \text{ MeV}}$$

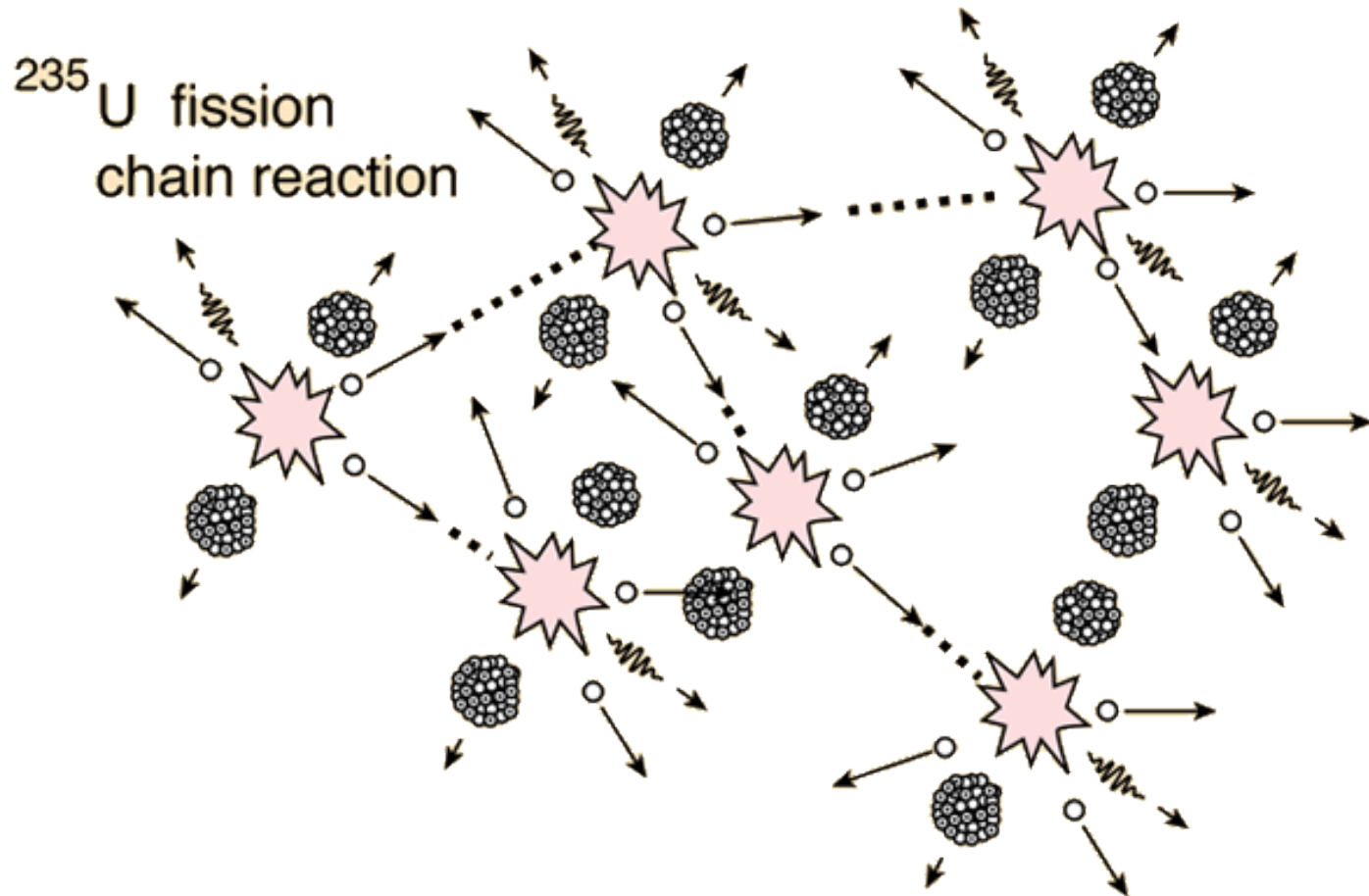
Mass Defect

**0.19446 u**

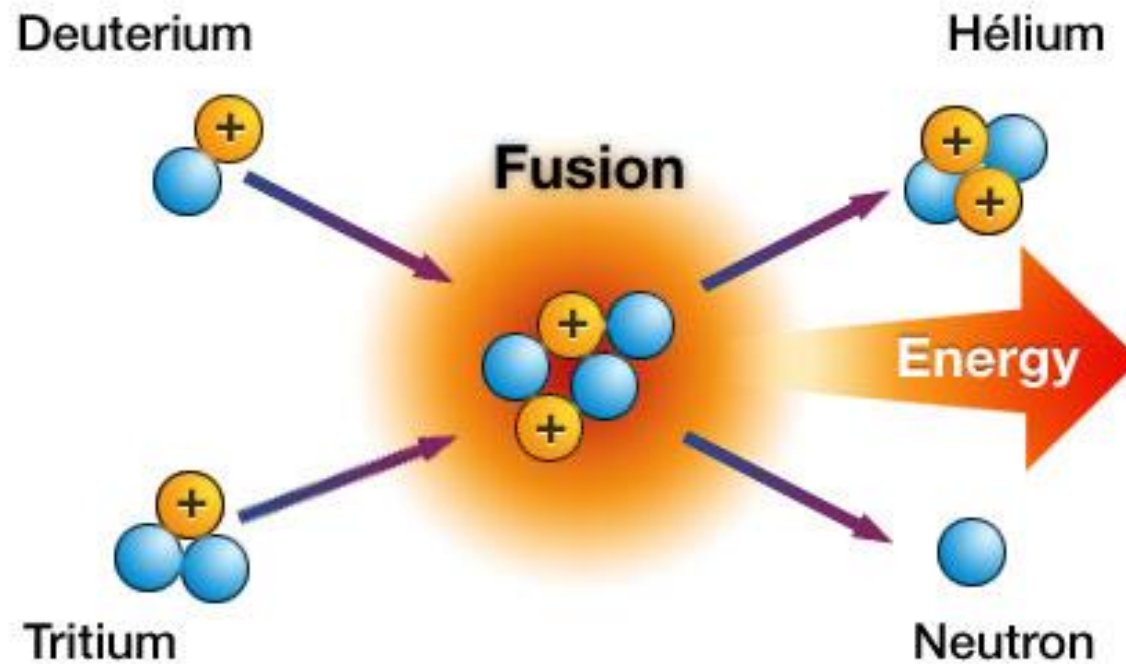
Energy Released

**181.14 MeV**

# Chain Reaction

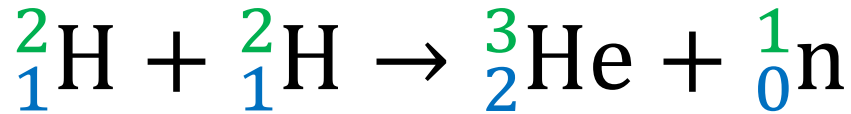


# Fusion



# Fusion

Hydrogen-2	2.0141 u
Helium-3	3.0161 u
Neutron	1.0087 u



$$(2.0141 \text{ u} + 2.0141 \text{ u}) - (3.0161 \text{ u} + 1.0087 \text{ u}) = 0.0034 \text{ u}$$

4.0282 u                      4.0248 u

$$0.0034 \text{ u} \times 931.5 = 3.1671 \text{ MeV}$$

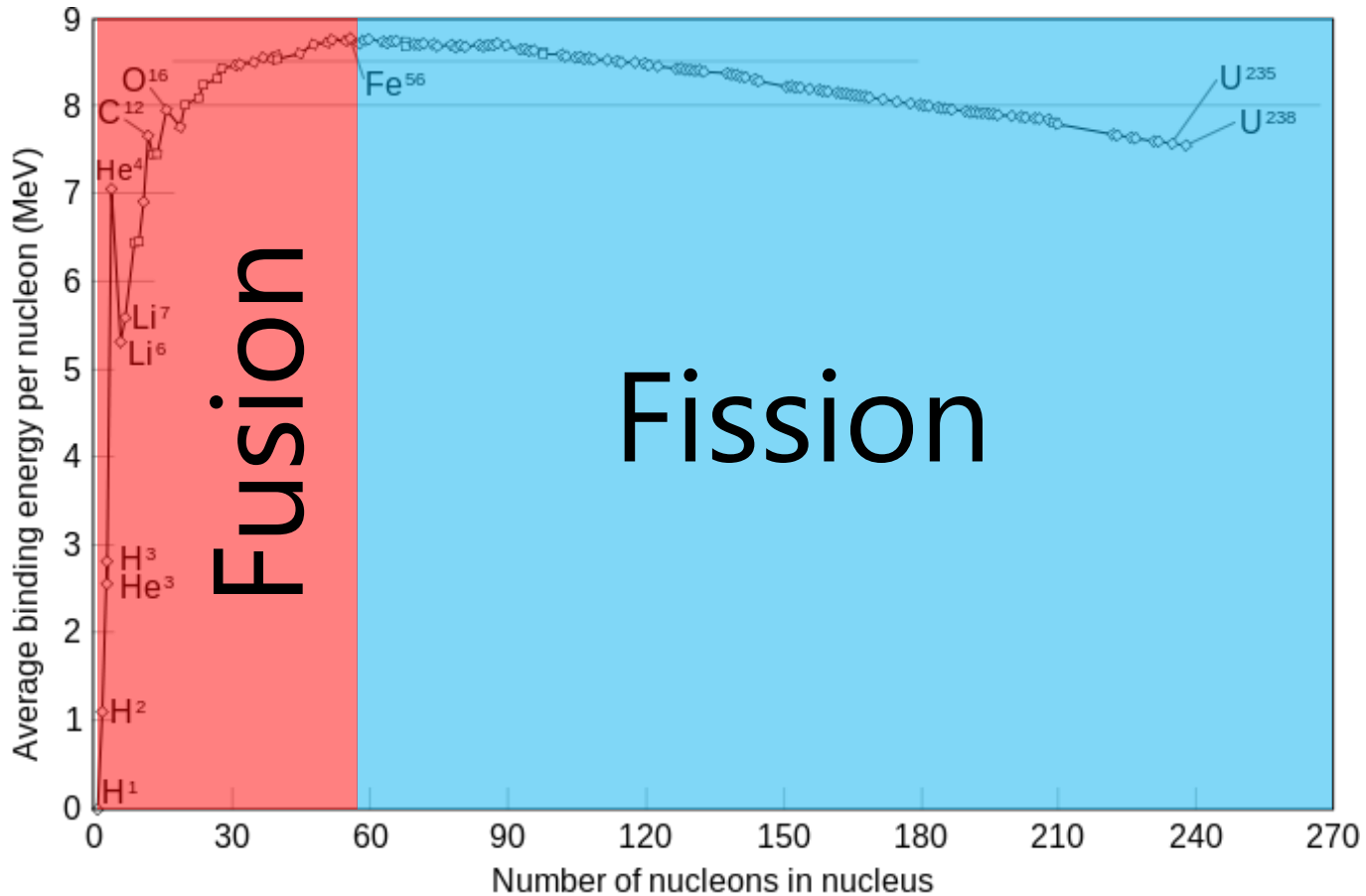
Mass Defect

0.0034 u

Energy Released

3.1671 MeV

# Fusion vs. Fission





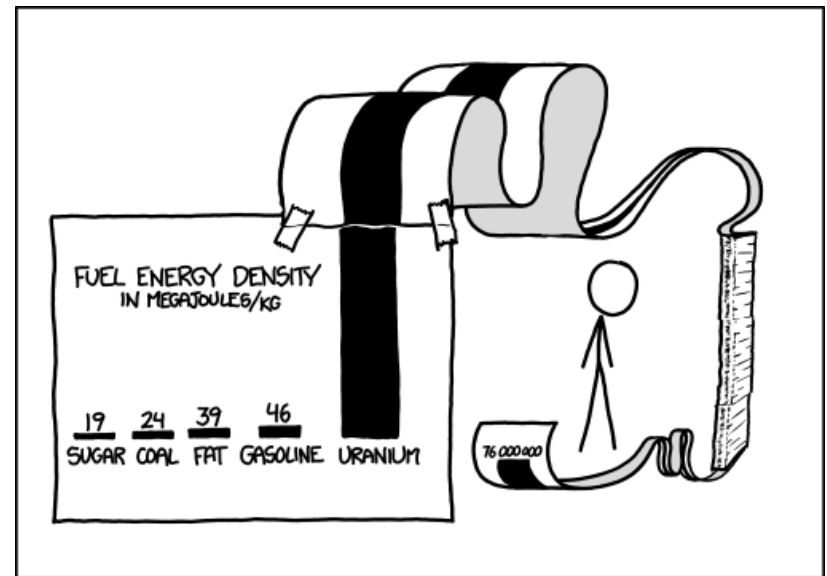
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SCIENCE TIP: LOG SCALES ARE FOR QUITTERS WHO CAN'T FIND ENOUGH PAPER TO MAKE THEIR POINT PROPERLY.

# Uranium



Uranium found in the earth's crust is primarily comprised of two different isotopes of Uranium



Uranium-238

99.3%



Uranium-235

0.7%

Where does the uranium used by the US come from?

37%	Kazakhstan, Russia, and Uzbekistan
30%	Canada
17%	Australia
10%	Malawi, Namibia, Niger, and South Africa
6%	United States

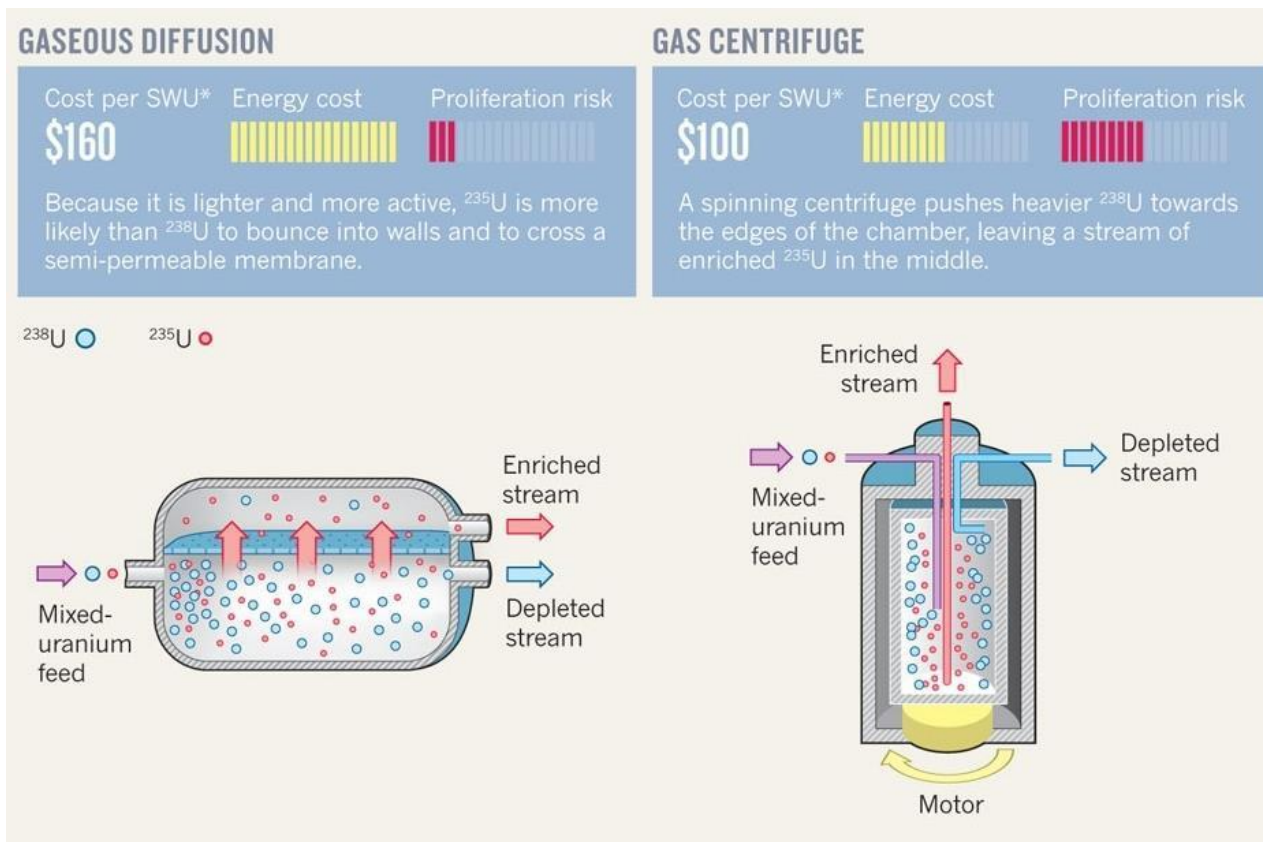
# Yellowcake Uranium



Uranium ore is milled into a  $U_3O_8$  powder known as **yellowcake**

# Uranium Enrichment

The yellowcake is converted into Uranium Hexafluoride gas and enriched to create a mixture with a higher percentage of U-235 nuclides

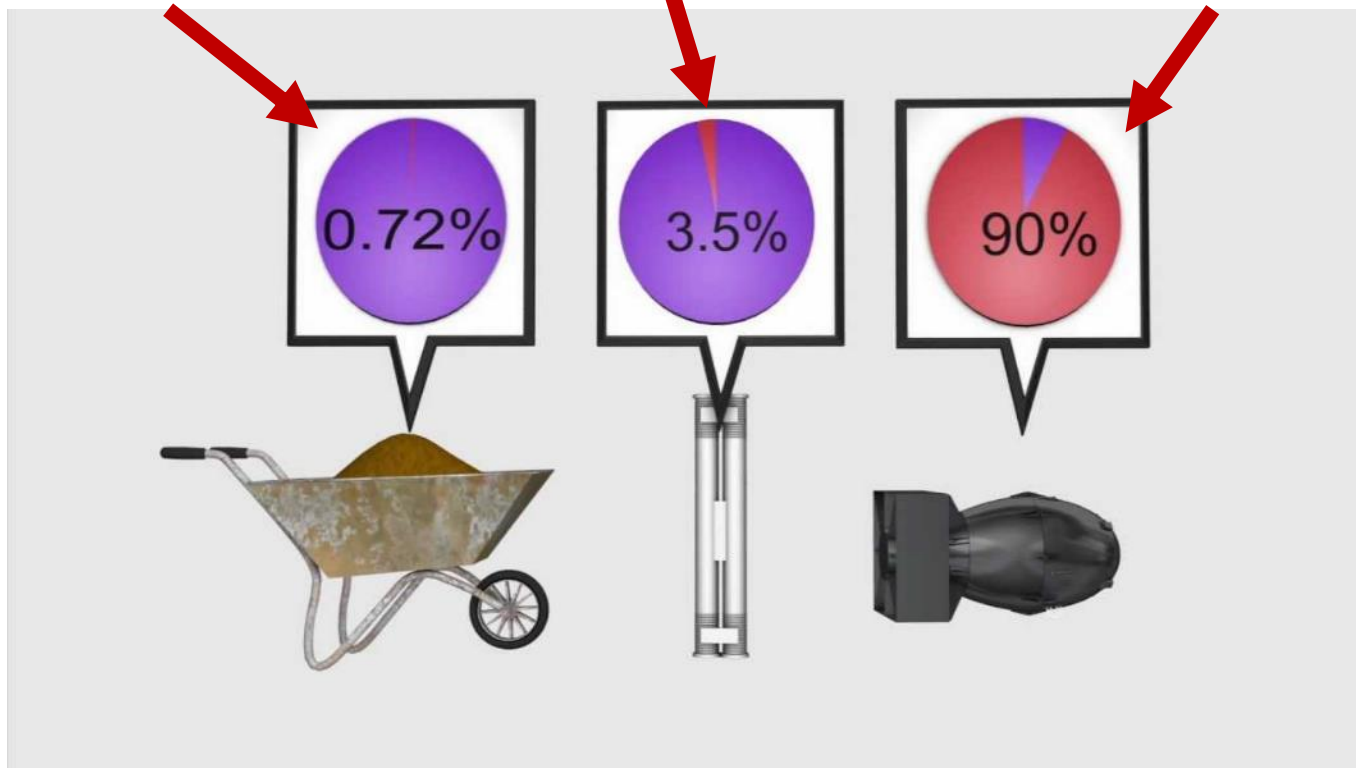


# Uranium Enrichment

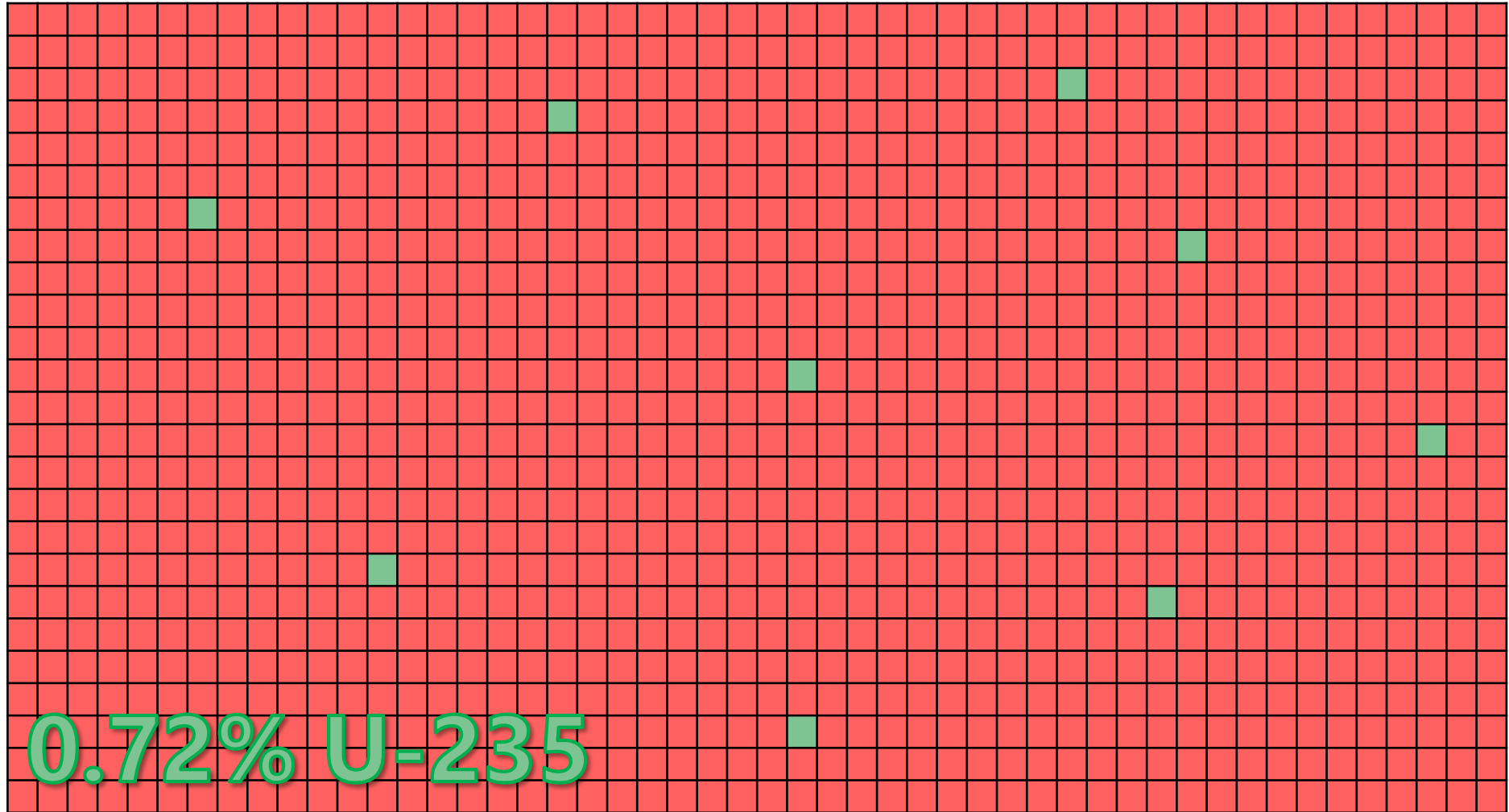
Fuel-Grade

Uranium Ore

Weapons-Grade

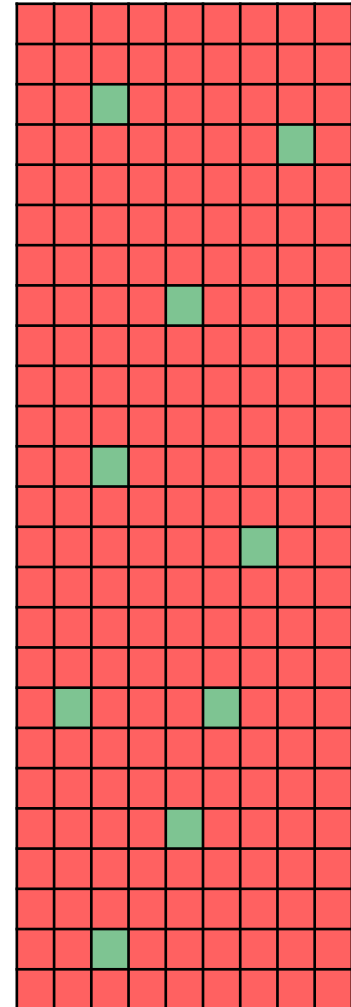


# Uranium Enrichment



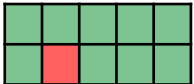
# Uranium Enrichment

4% U-235



# Uranium Enrichment

90% U-235



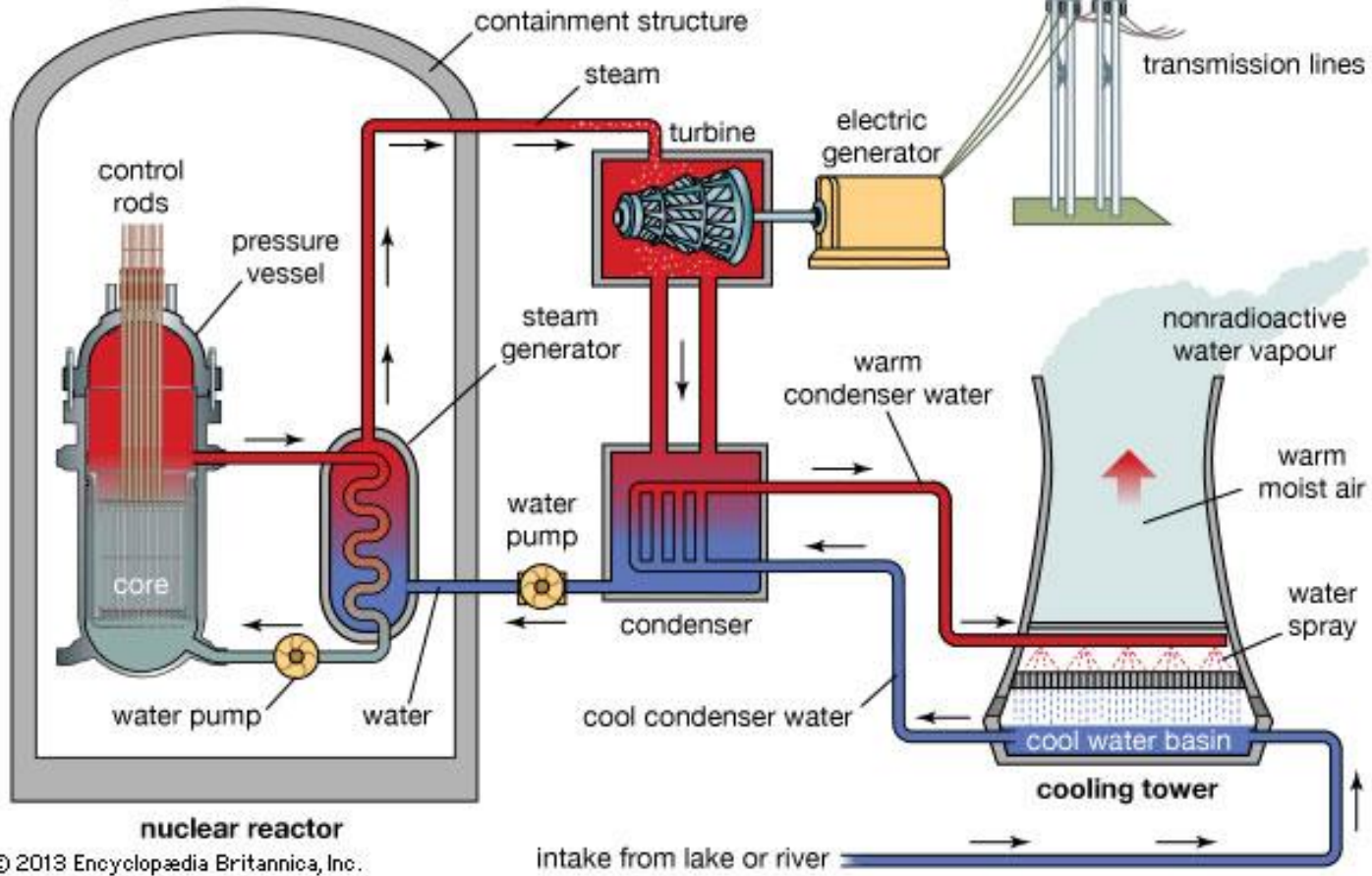


# Uranium Fuel Rods



# Nuclear Power Plant

Nuclear power plant

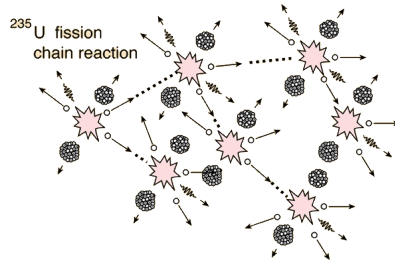


# Chain Reaction

For a controlled chain reaction, each reaction should trigger one other reaction

## Important Factors

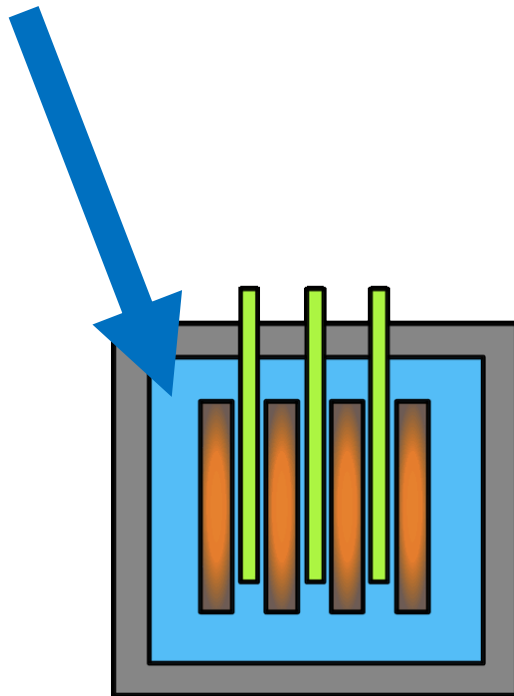
- Only about 4% of the fuel is actually comprised of fissionable U-235 atoms
- Neutrons have to be traveling relatively slowly to be captured by a U-235 atom



# Chain Reaction

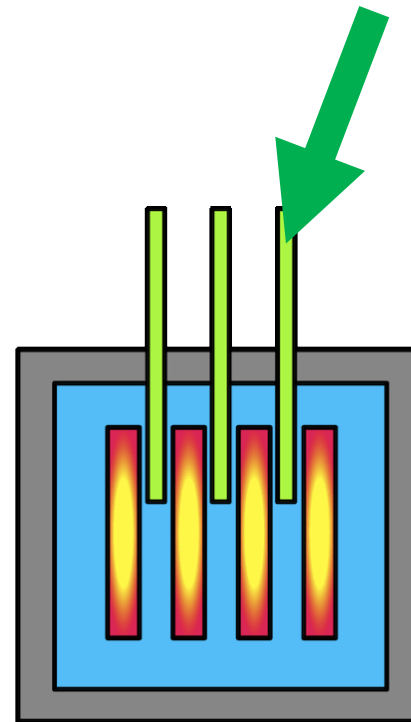
Moderator - water or graphite

Slows down neutrons

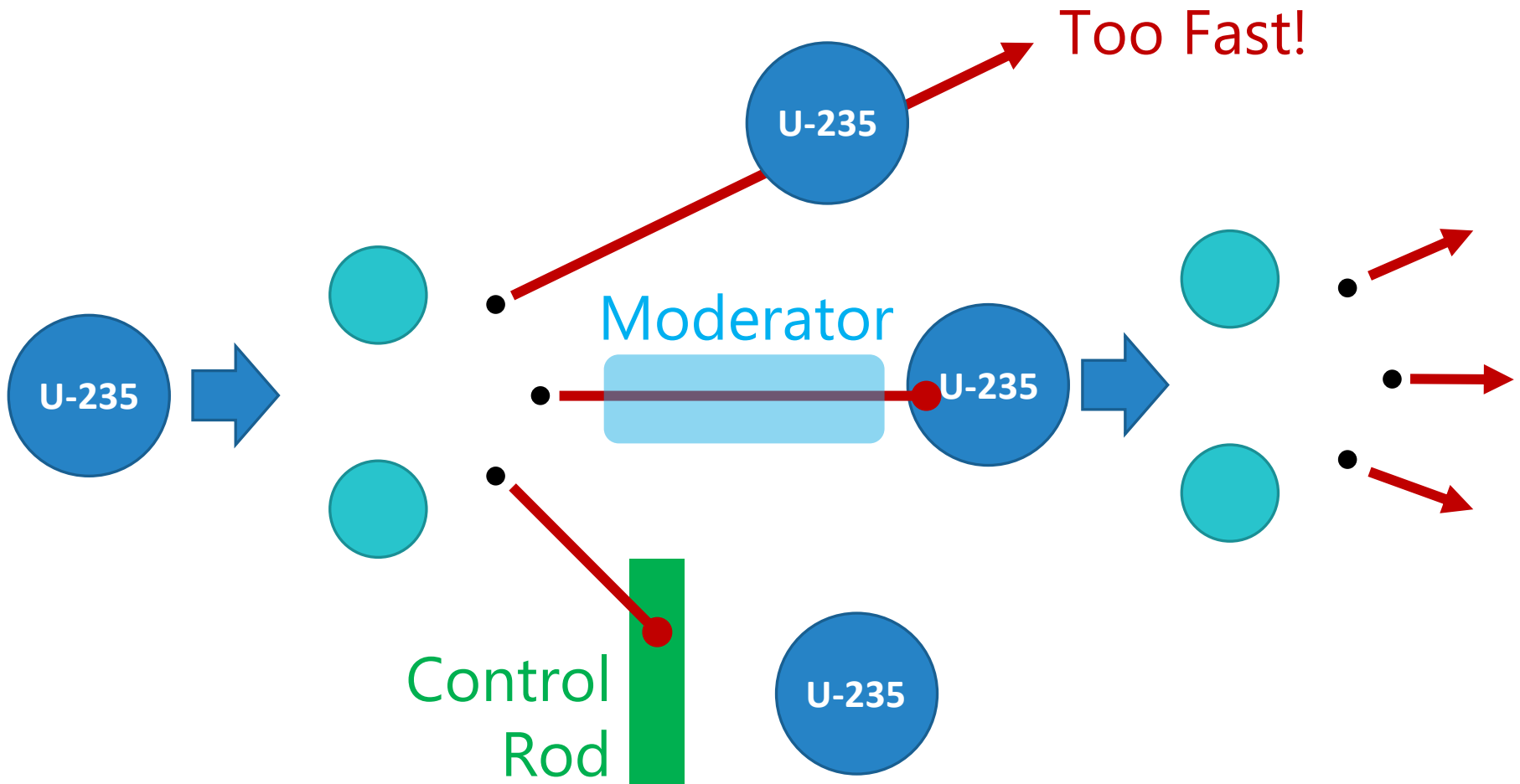


Control Rods - boron

Absorbs neutrons



# Chain Reaction



# Nuclear Waste

After the fuel can no longer be used to efficiently create electricity, the spent nuclear waste needs to be disposed of.

Isotope	Half Life
Strontium-90	28 years
Caesium-137	30 years
Plutonium-239	24,000 years
Caesium-135	2.3 million years
Iodine-129	15.7 million years

How long until  
it's safe?

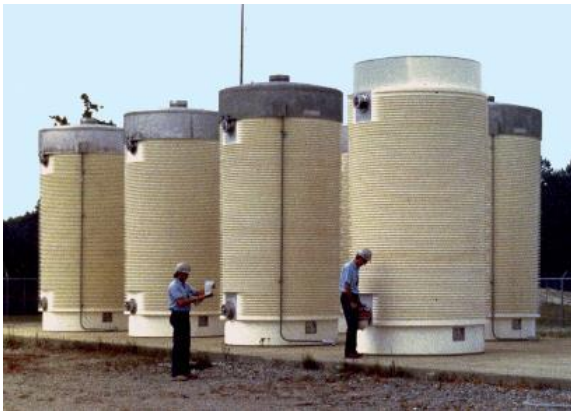
Radioactive Isotopes found  
in spent nuclear fuel



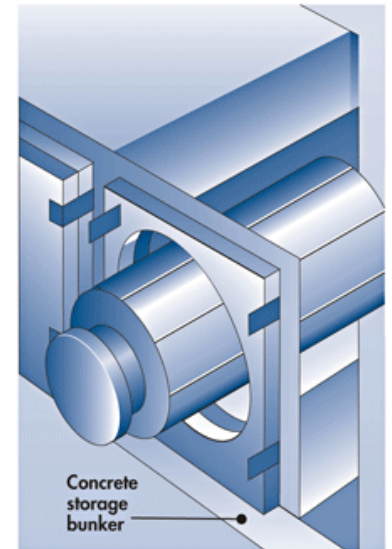
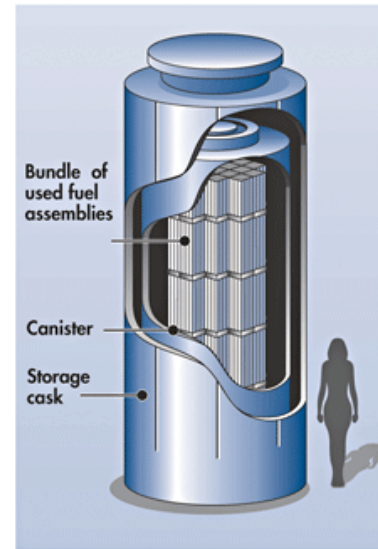
# Nuclear Waste - Disposal



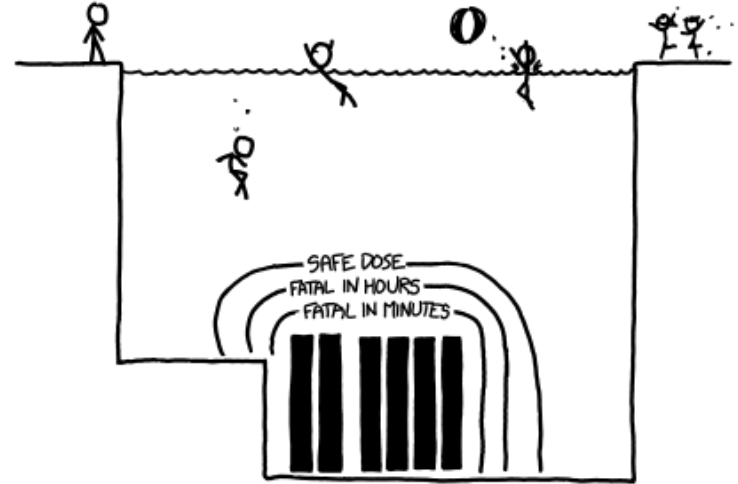
Most of the waste is stored onsite



Dry Storage of Spent Fuel



# Nuclear Waste - Disposal



## Spent Fuel Pool

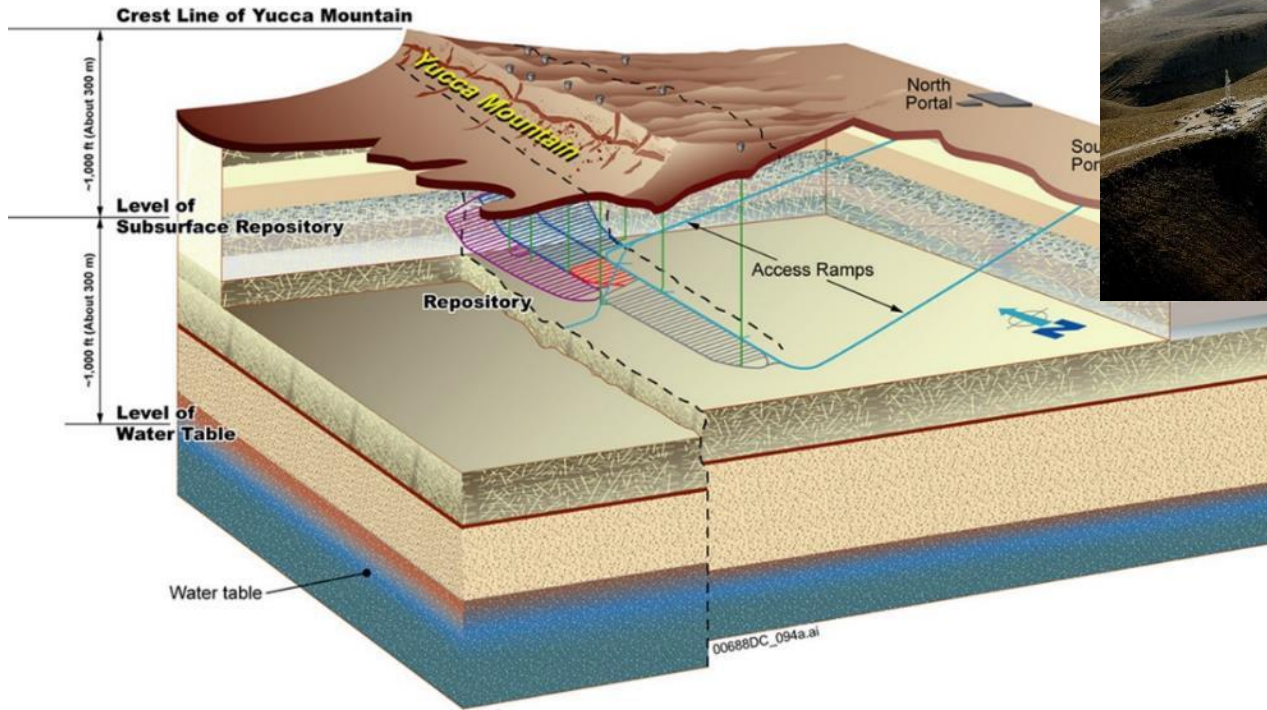
*What if I took a swim in a typical spent nuclear fuel pool? Would I need to dive to actually experience a fatal amount of radiation? How long could I stay safely at the surface?*

*—Jonathan Bastien-Filiatrault*



# Nuclear Waste - Disposal

## Yucca Mountain



# Nuclear Waste - Disposal

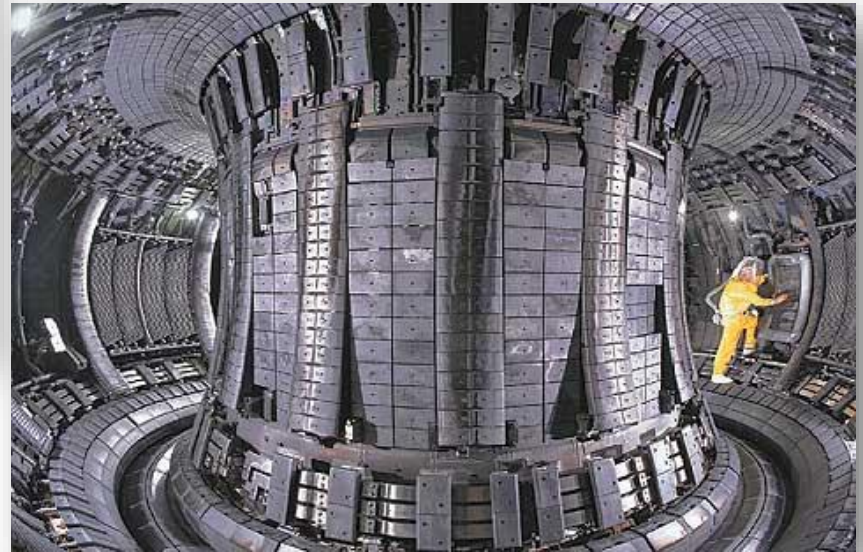
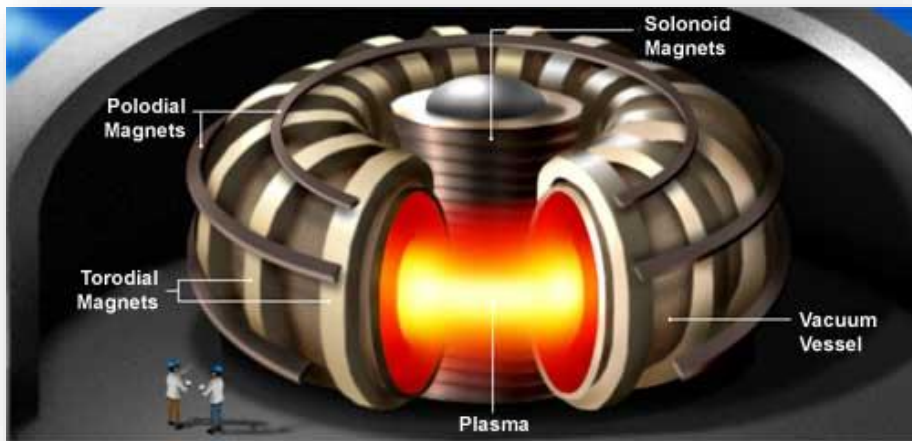
Any better options out there???



Ideas	Examples
<u>Long-term above ground storage</u>	<ul style="list-style-type: none"> <li>• Investigated in France, Netherlands, Switzerland, UK, and USA.</li> <li>• Not currently planned to be implemented anywhere.</li> </ul>
<u>Disposal in outer space</u> (proposed for wastes that are highly concentrated)	<ul style="list-style-type: none"> <li>• Investigated by USA.</li> <li>• Investigations now abandoned due to cost and potential risks of launch failure.</li> </ul>
<u>Rock-melting</u> (proposed for wastes that are heat-generating)	<ul style="list-style-type: none"> <li>• Investigated by Russia, UK, and USA.</li> <li>• Not implemented anywhere.</li> <li>• Laboratory studies performed in the UK.</li> </ul>
<u>Disposal at subduction zones</u>	<ul style="list-style-type: none"> <li>• Investigated by USA.</li> <li>• Not implemented anywhere.</li> <li>• Not permitted by international agreements.</li> </ul>
<u>Sea disposal</u>	<ul style="list-style-type: none"> <li>• Implemented by Belgium, France, Germany, Italy, Japan, Netherlands, Russia, South Korea, Switzerland, UK, and USA.</li> <li>• Not permitted by international agreements.</li> </ul>
<u>Sub seabed disposal</u>	<ul style="list-style-type: none"> <li>• Investigated by Sweden and UK (and organisations such as the OECD Nuclear Energy Agency).</li> <li>• Not implemented anywhere.</li> <li>• Not permitted by international agreements.</li> </ul>
<u>Disposal in ice sheets</u> (proposed for wastes that are heat-generating)	<ul style="list-style-type: none"> <li>• Investigated by USA.</li> <li>• Rejected by countries that have signed the Antarctic Treaty or committed to providing solutions within national boundaries.</li> </ul>
<u>Deep well injection</u> (for liquid wastes)	<ul style="list-style-type: none"> <li>• Implemented in Russia for many years for LLW and ILW.</li> <li>• Investigations abandoned in the USA in favour of deep geological disposal of wastes in solid form.</li> </ul>

# Fusion as a Power Source

Fusion reactions have been successfully controlled using strong magnetic fields but the energy used to run the magnets exceeds the energy released in the reaction...

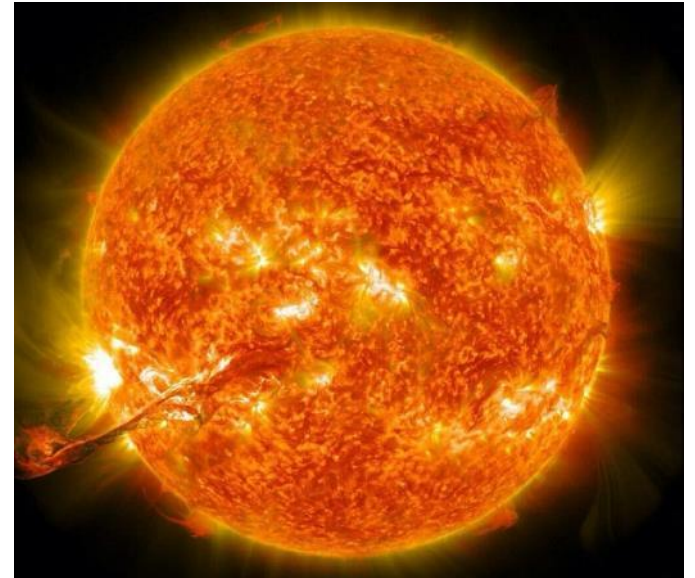




# Conditions for Fusion

It's significantly more difficult to create fusion reactions here on earth

- High Pressure
- High Temperature



# The Renewables

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# Renewable vs. Non Renewable

Highlight the primary energy sources that are considered **renewable**



**Petroleum**



**Geothermal**



**Coal**



**Hydropower**



**Uranium**



**Solar**



**Natural Gas**



**Wind**



**Propane**

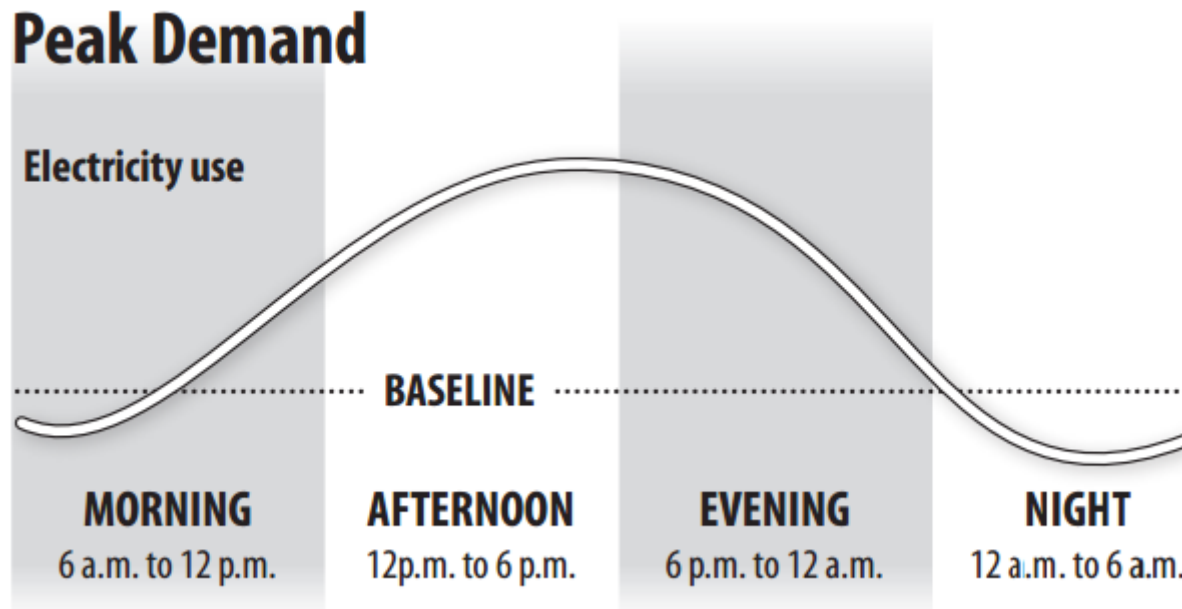


**Biomass**

\*Note: this doesn't mean that it cannot ever be replaced, just that it won't happen in any sort of useful time frame...

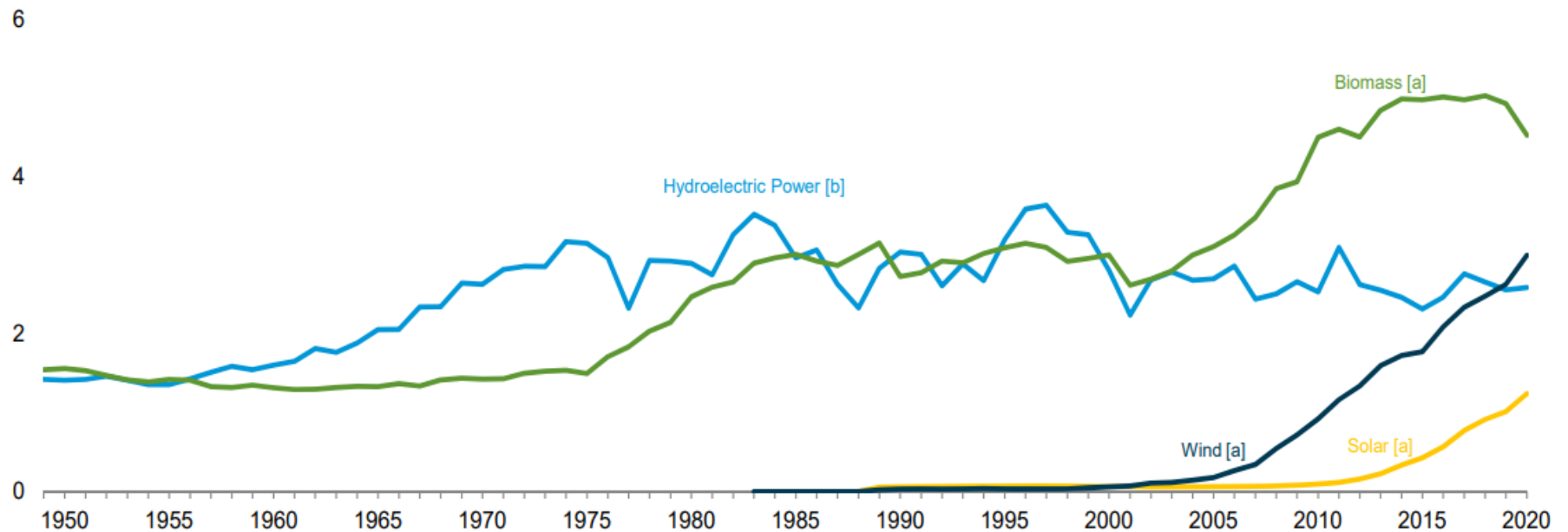
# Energy Load Requirements

Energy needs to be available when electricity is most needed but should also be available other times as well.



# Renewable Energy in the US

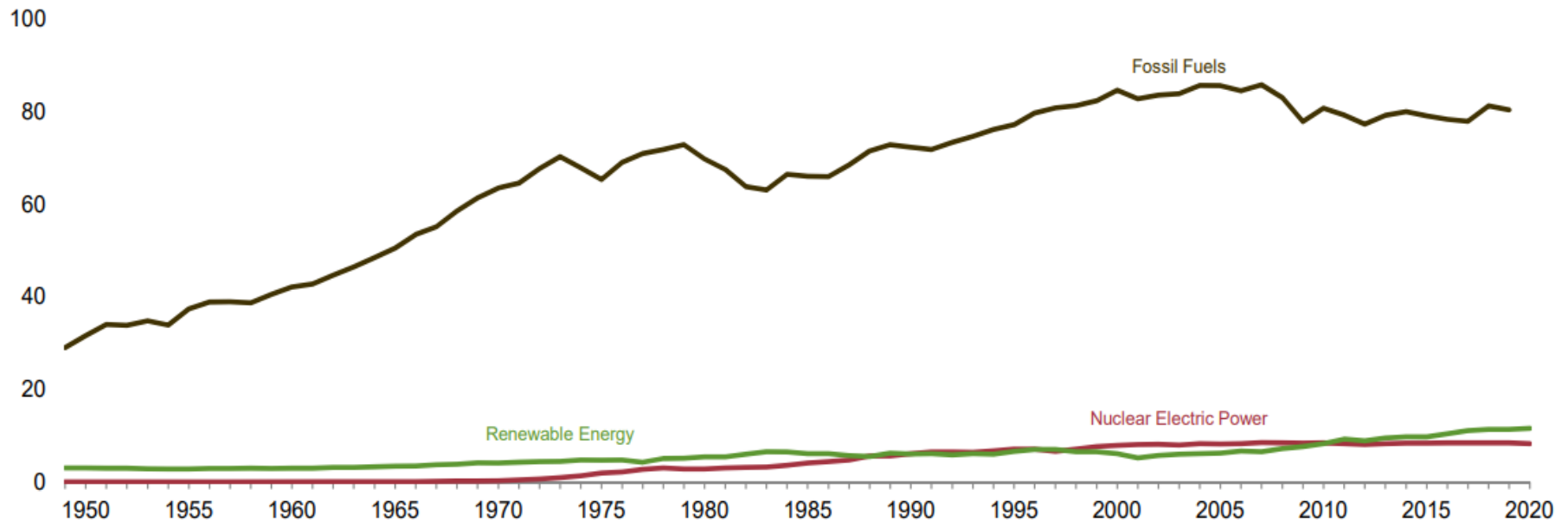
Major Sources, 1949–2020



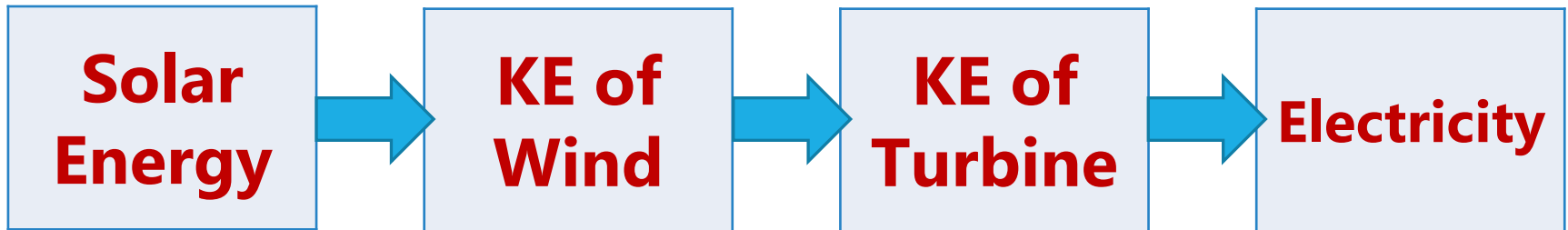
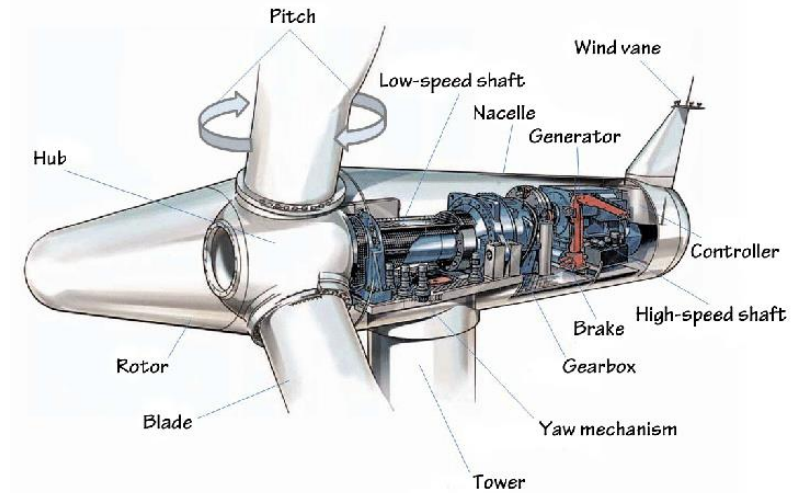
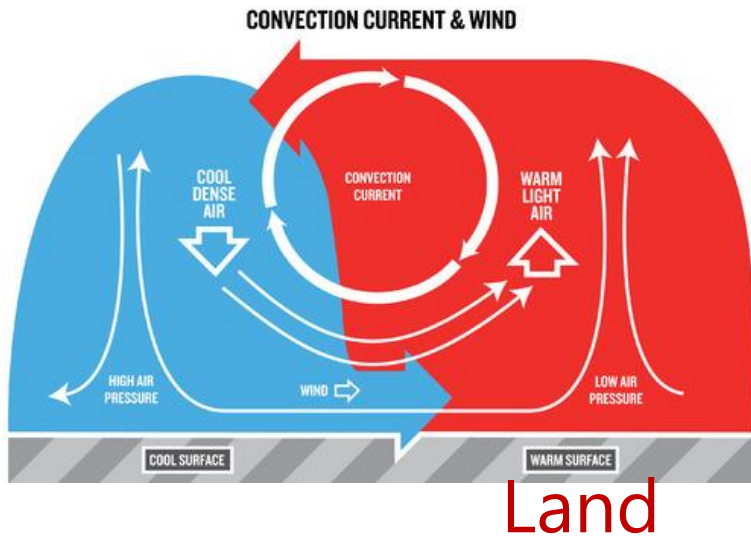


# Renewable Energy in the US

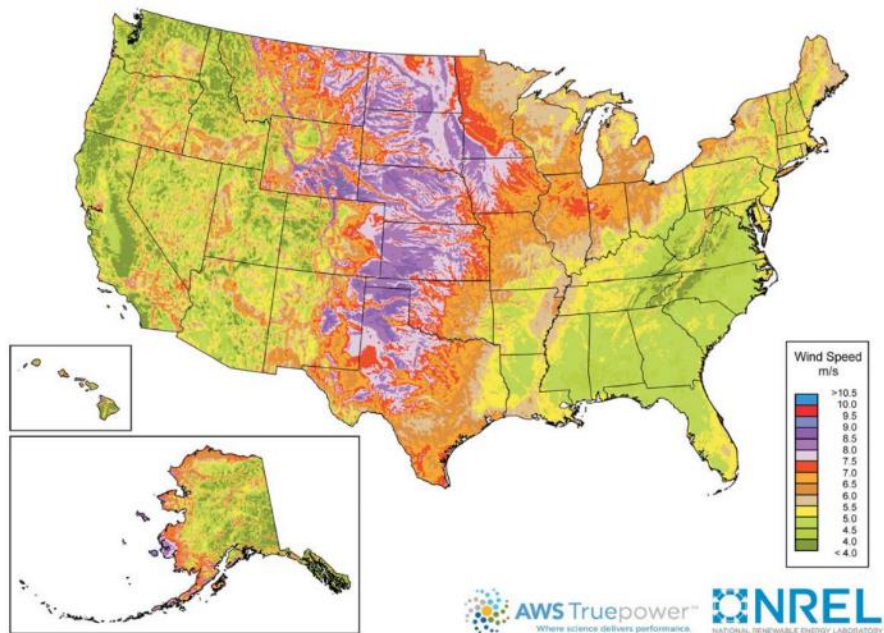
Compared With Other Resources, 1949–2020



# Wind Power

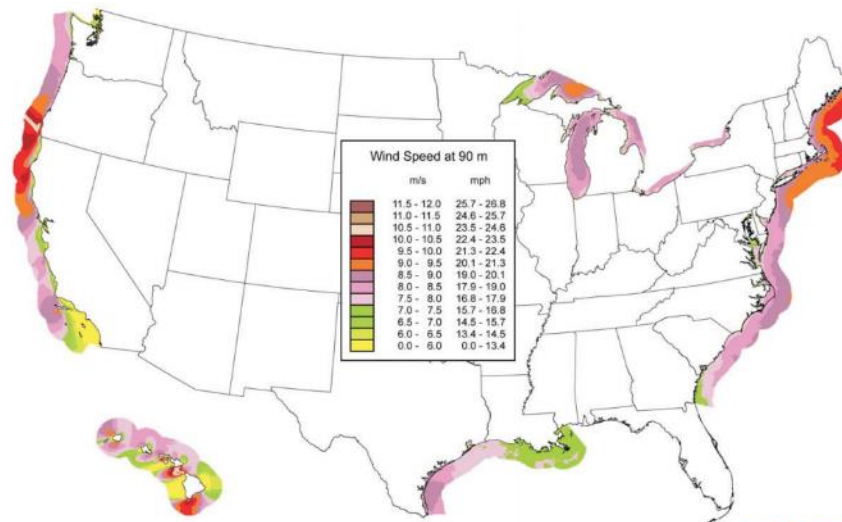


# Wind Speeds



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NATIONAL RENEWABLE ENERGY LABORATORY



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# NIMBY

## ARGUMENTS AGAINST-



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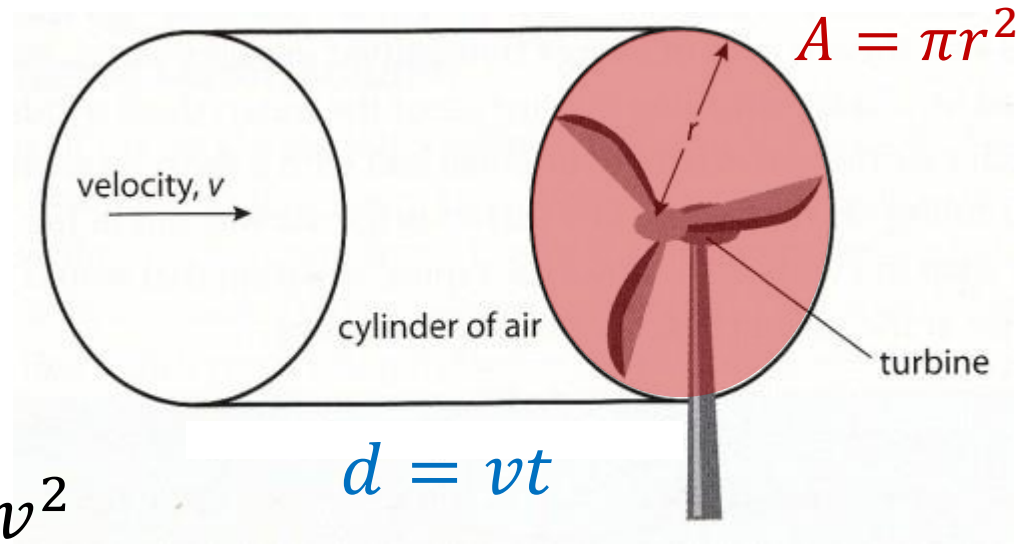


# Calculating the Wind's Energy

$$E_K = \frac{1}{2}mv^2$$

$$\text{Power} = \frac{E_K}{t} = \frac{\frac{1}{2}[Avt\rho]v^2}{t}$$

$$\text{Power} = \frac{1}{2}A\rho v^3$$



$\rho \rightarrow$  air density

$$V = A \times vt$$

$$m = (A \times vt) \times \rho$$

# IB Physics Data Booklet

$$Power = \frac{1}{2} A \rho v^3$$

← Velocity [m s<sup>-1</sup>]  
← Air Density [kg m<sup>-3</sup>]  
↑ Area [m<sup>2</sup>]  
 (A = πr<sup>2</sup>)

Sub-topic 8.1 – Energy sources	Sub-topic 8.2 – Thermal energy transfer
$Power = \frac{energy}{time}$ <div style="border: 2px solid red; background-color: yellow; padding: 5px; width: fit-content;"> <math>Power = \frac{1}{2} A \rho v^3</math> </div>	$P = e\sigma AT^4$ $\lambda_{max}(metres) = \frac{2.90 \times 10^{-3}}{T(kelvin)}$ $I = \frac{power}{A}$ $albedo = \frac{total\ scattered\ power}{total\ incident\ power}$



# Conceptual Meaning of Equation

$$Power = \frac{1}{2} A \rho v^3$$

If the wind speed is doubled, then the power is multiplied by a factor of **8**  $2^3$

If the wind speed is tripled, then the power is multiplied by a factor of **27**  $3^3$

# Try This...

$$Power = \frac{1}{2} A \rho v^3$$

Given a turbine having a blade length of 12 m, and a wind speed of 15 ms<sup>-1</sup> find the power output if the density of air is  $\rho = 1.2 \text{ kg m}^{-3}$ .

$$P = \frac{1}{2} (\pi \times 12^2) (1.2) (15)^3 = 916,000 \text{ W}$$

$$= 0.916 \text{ MW}$$

What is the actual power output if the efficiency is 45%?

$$0.916 \text{ MW} \times 0.45 = 0.412 \text{ MW}$$



# Try This...

$$Power = \frac{1}{2} A \rho v^3$$

Air of constant density  $1.2 \text{ kg m}^{-3}$  is incident at a speed of  $9.0 \text{ m s}^{-1}$  on the blades of a wind turbine. The turbine blades are each of length  $7.5 \text{ m}$ . The air passes through the turbine without any change of direction. Immediately after passing through the blades, the speed of the air is  $5.0 \text{ m s}^{-1}$ . The density of air immediately after passing through the blades is  $2.2 \text{ kg m}^{-3}$ .

$$P = \frac{1}{2}(\pi \times 7.5^2)(1.2)(9)^3 = 77,300 \text{ W}$$

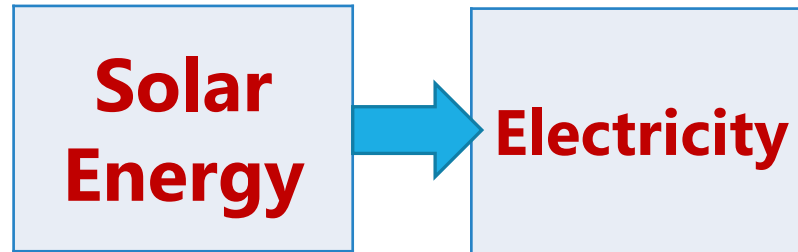
$$P = \frac{1}{2}(\pi \times 7.5^2)(2.2)(5)^3 = 24,300 \text{ W}$$

---

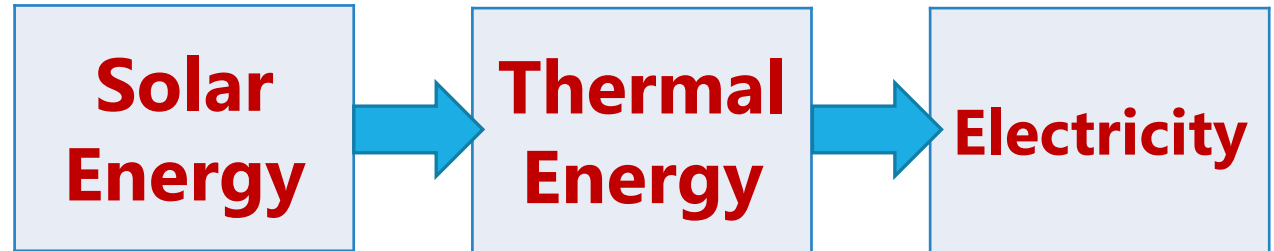
$$53,000 \text{ W}$$

# Solar Power

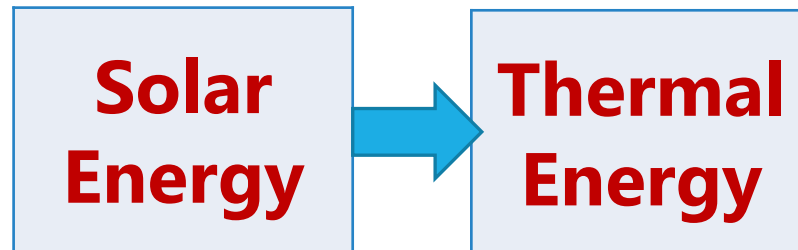
Photovoltaic Cells



Solar Concentrator



Solar Heating Panel



# Efficiency of PV Cells

Photovoltaic Cells



10%-20%  
Efficient

# Solar Power

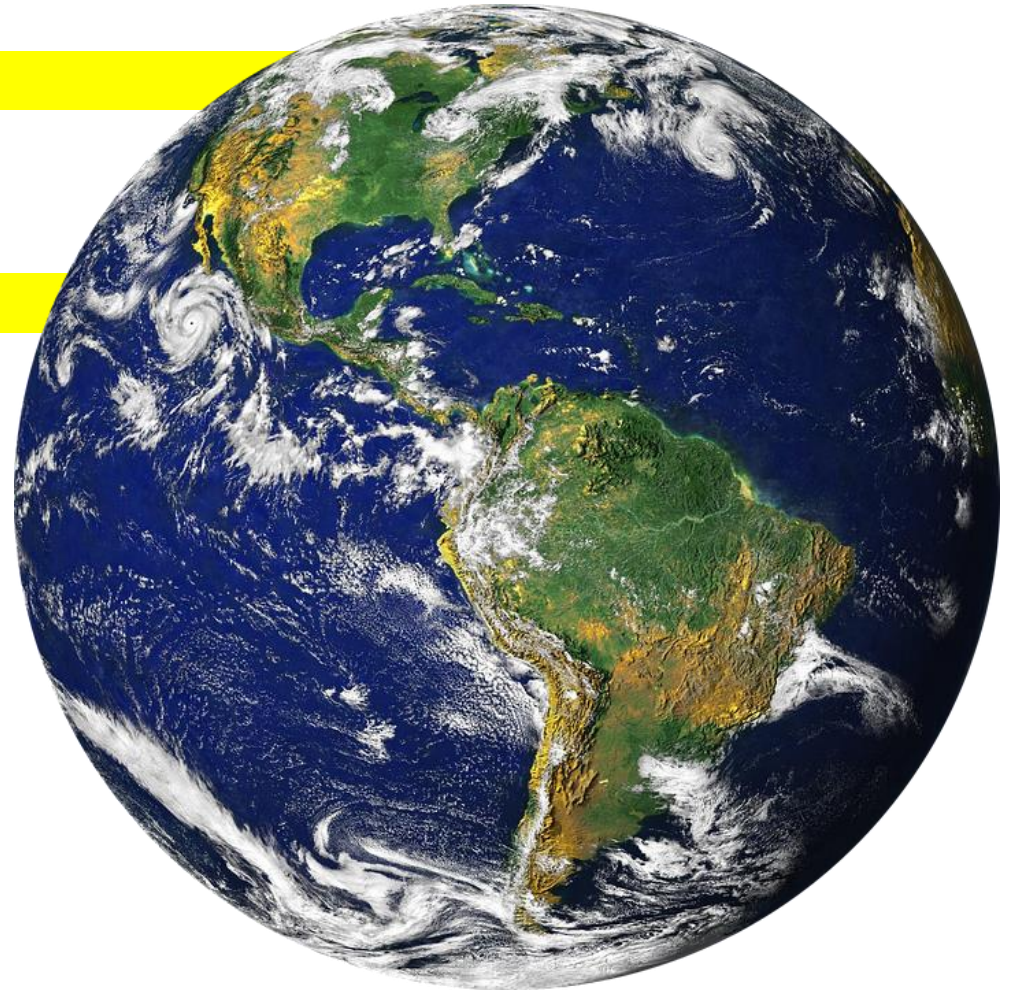


What Would It Take To Power The United States With Solar Energy?

## Issues

- Infrastructure to transport electricity
- Storage for non-sun times
- High up front cost

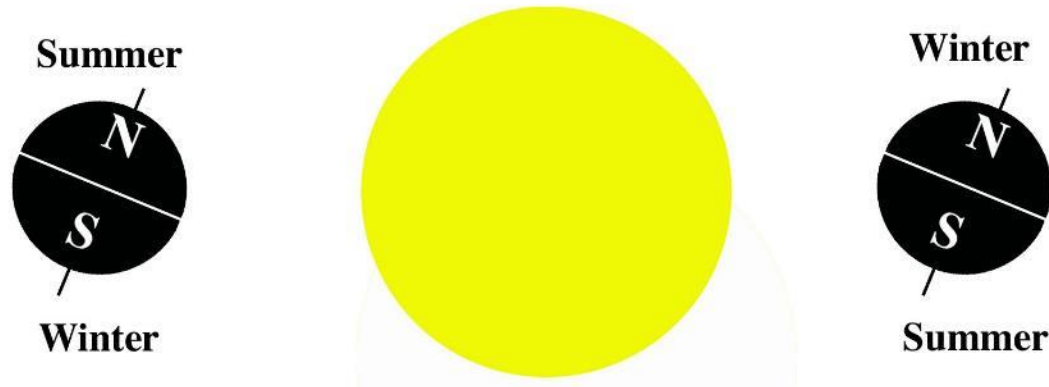
# Solar Power Intensity



Same Power  
**Different Area**

Solar intensity depends  
on the latitude

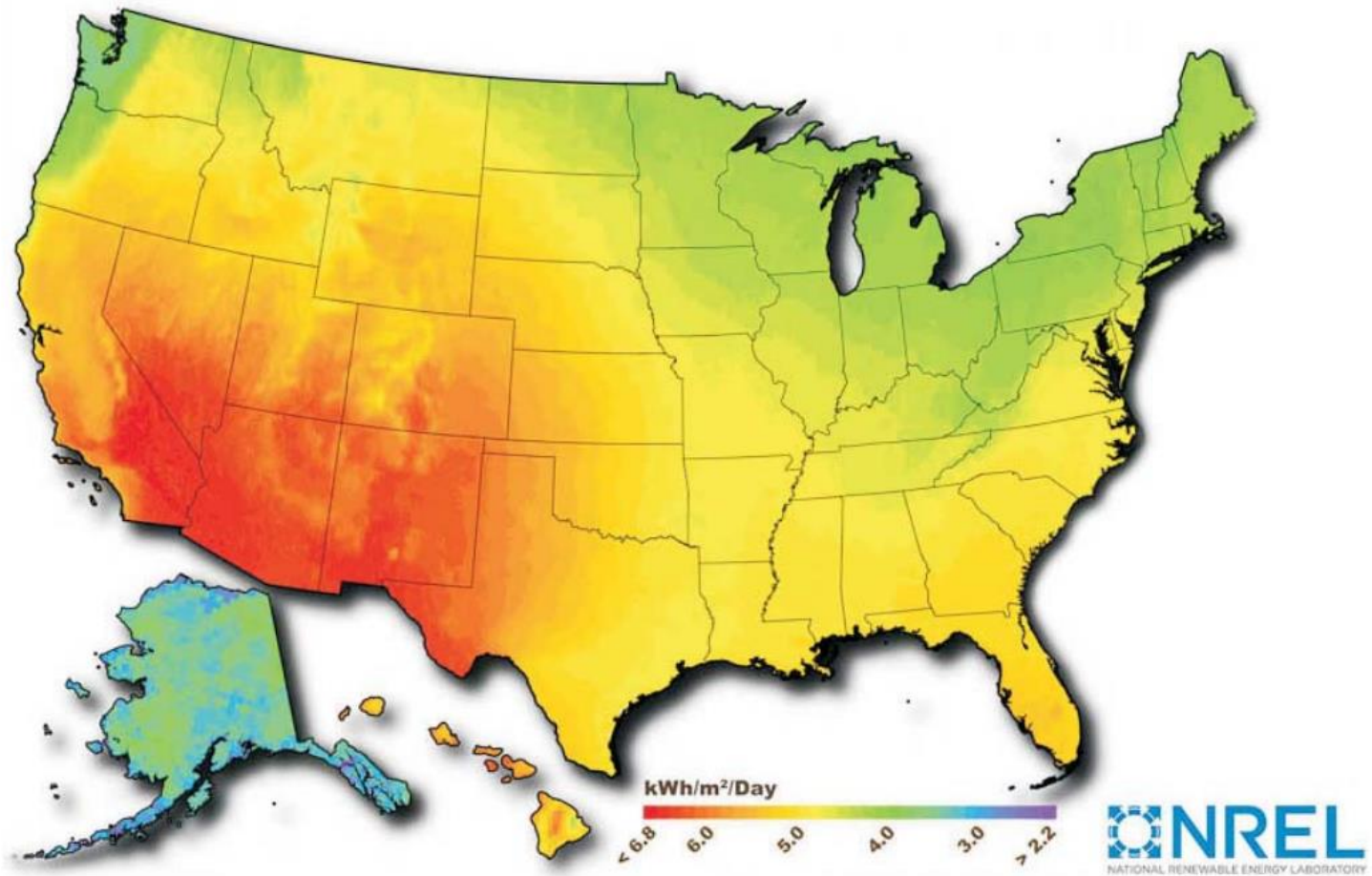
# This also affects the seasons



27. The annual variations of solar power incident per unit area at a particular point on the Earth's surface is mainly due to the change in the
- A. distance between the Earth and the Sun.
  - B. angle at which the solar rays hit the surface of the Earth.
  - C. average albedo of the Earth.
  - D. average cloud cover of the Earth.



# Solar Map



# Calculating Solar Power

A photovoltaic cell has an area of  $1.00 \text{ cm}^2$  and an efficiency of  $10.5\%$ . If the cell is placed in a position where the sun's intensity is  $1250 \text{ W m}^{-2}$ , what is the power output of the cell?

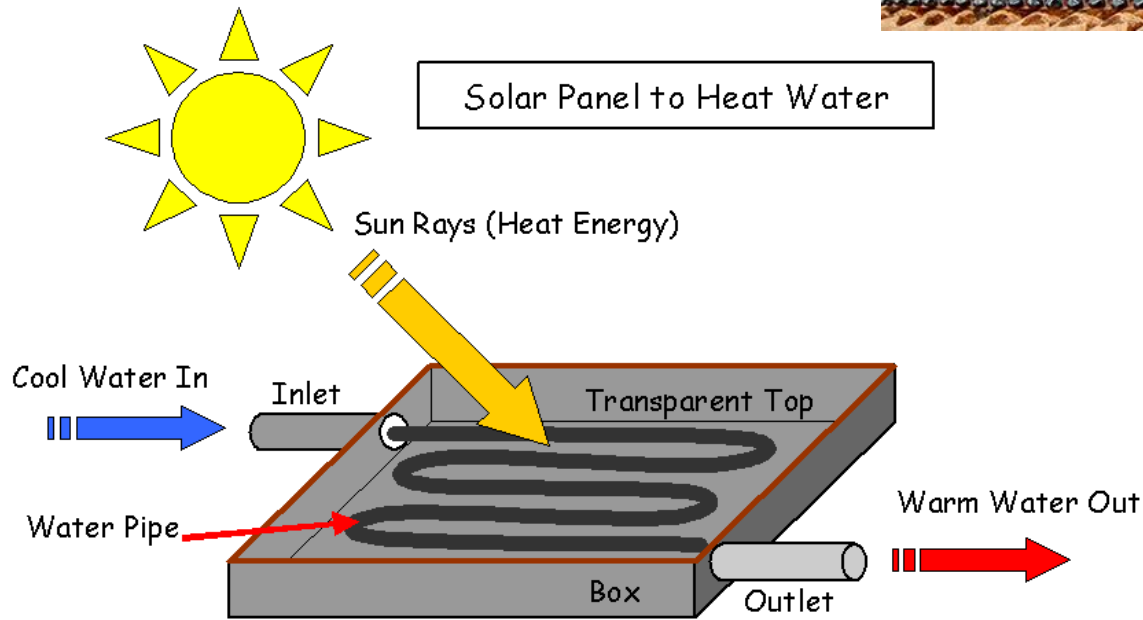


$$1 \text{ cm}^2 \times \frac{1 \text{ m}}{100 \text{ cm}} \times \frac{1 \text{ m}}{100 \text{ cm}} = 0.0001 \text{ m}^2$$

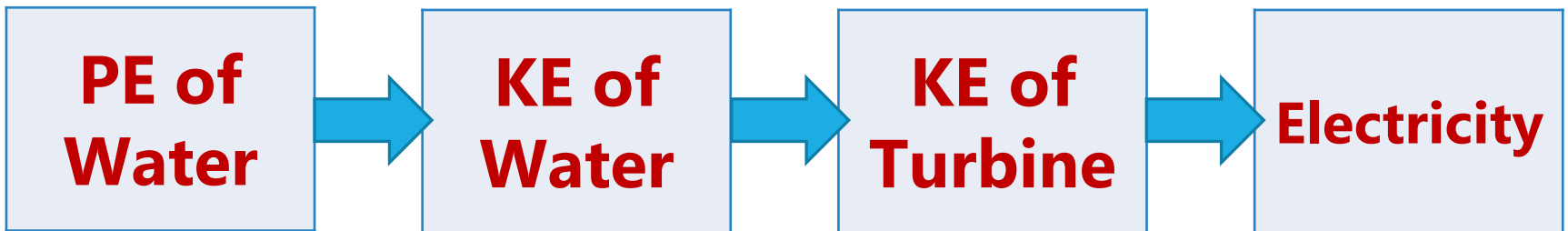
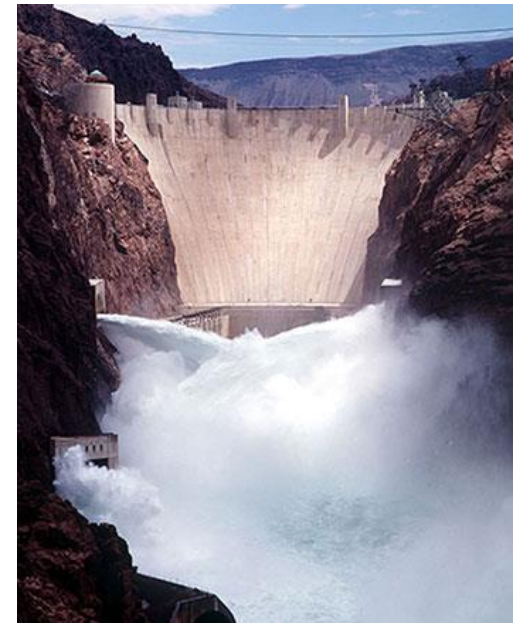
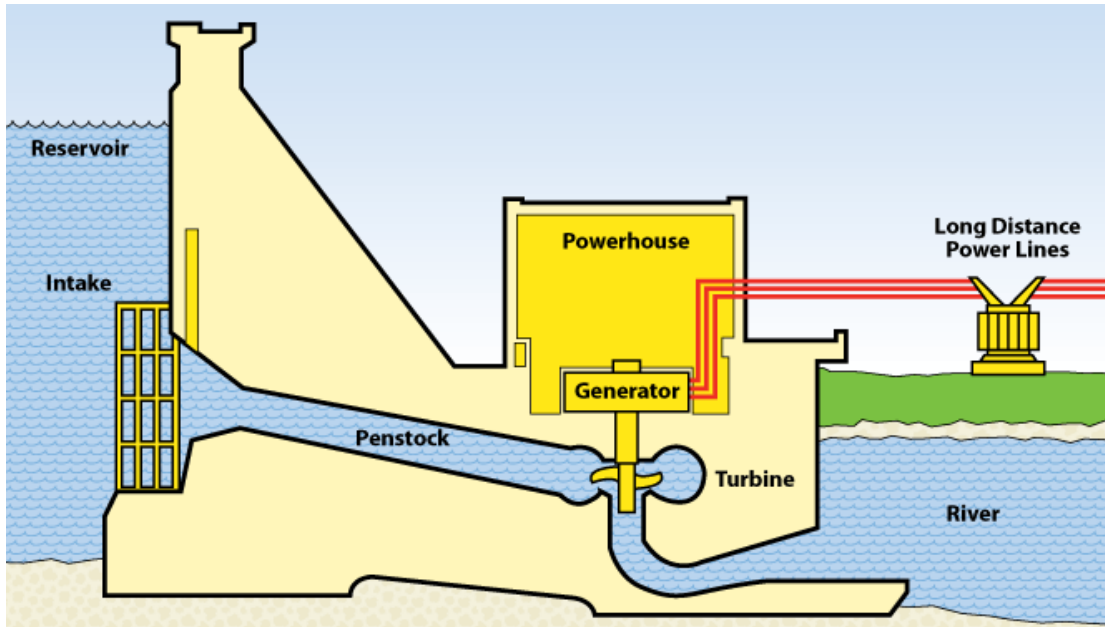
$$\frac{1250 \text{ W}}{1 \text{ m}^2} \times \underset{\substack{\uparrow \\ \text{Area}}}{0.0001 \text{ m}^2} \times \underset{\substack{\uparrow \\ \text{Efficiency}}}{0.105} = \boxed{0.0131 \text{ W}}$$



# Solar Heating Panel

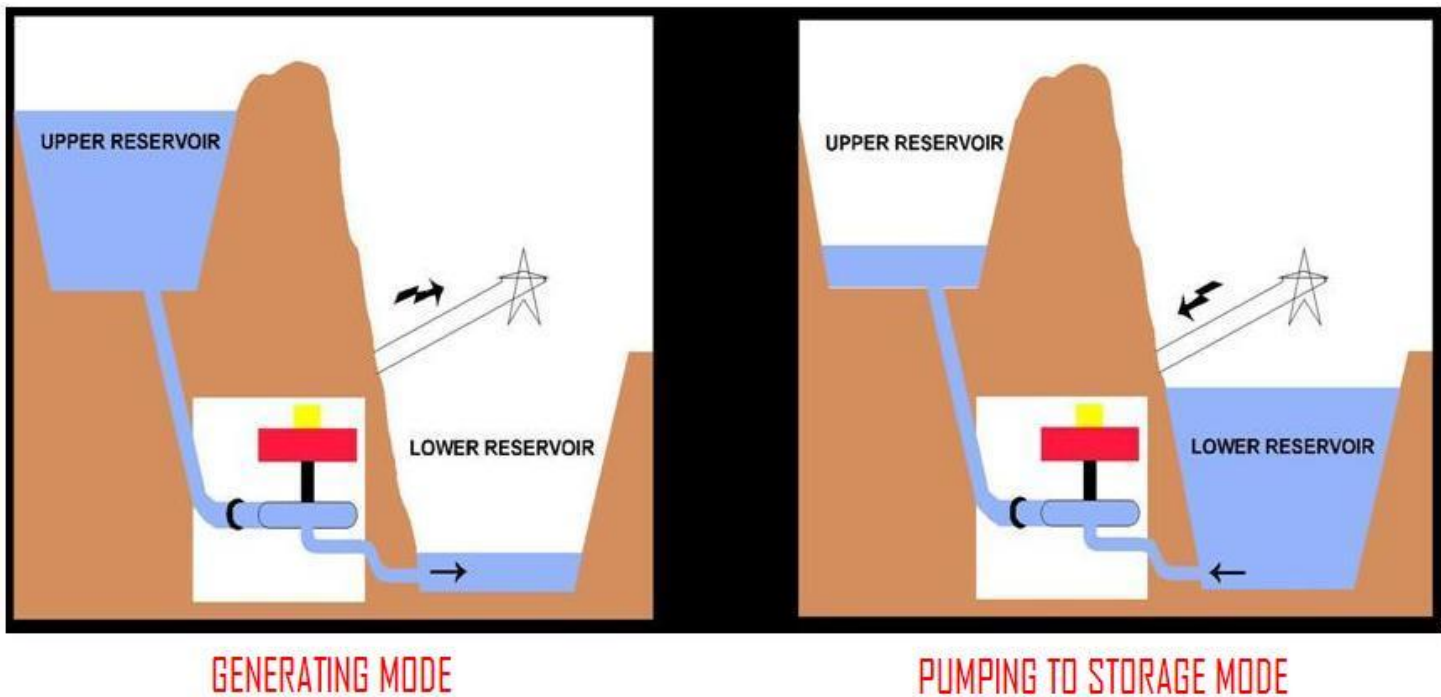


# Hydropower



# Storing Energy in Hydropower

If there is excess electricity, this energy can be stored by pumping water back up to the reservoir



pumped hydro operating principals

# Issues of the Renewables

- Storage
- Upfront cost
- Control over timing

# Thermal Energy Transfer & Black Body Radiation

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IB PHYSICS | ENERGY PRODUCTION

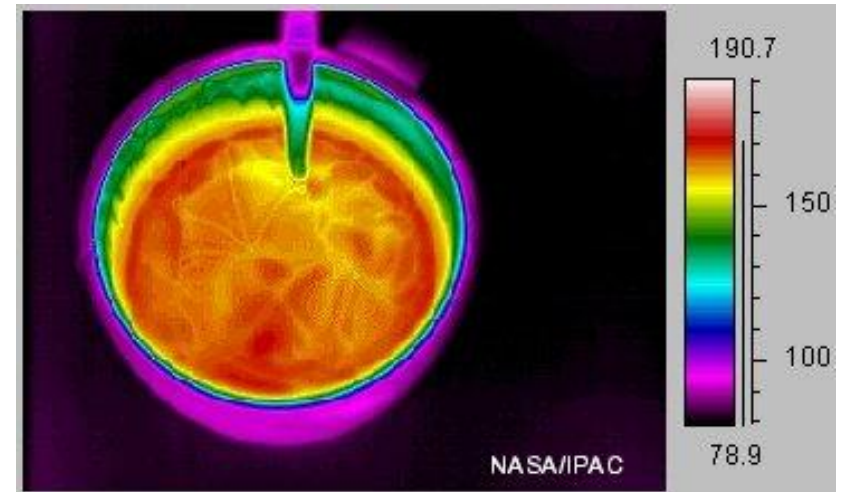
# Heat Transfer

There are 3 primary ways that heat is transferred:

- Conduction
- Convection
- Radiation

# Conduction

Conduction occurs between objects in direct contact



# Conduction

Why does this frying pan have a plastic handle?

Plastic has a high specific heat and doesn't conduct heat very quickly



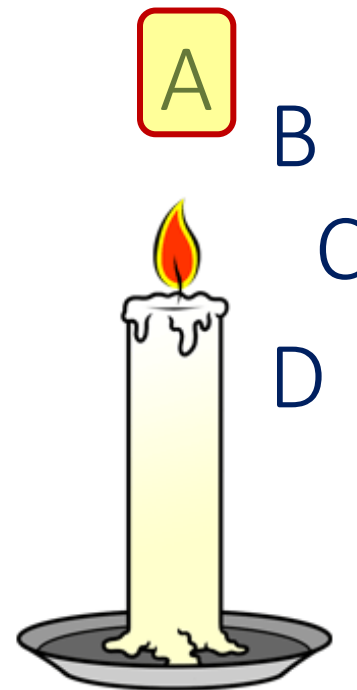


# Convection

**Convection** occurs when fluids (liquids or gases) move around due to temperature differences

Hot Air rises

Cold Air sinks



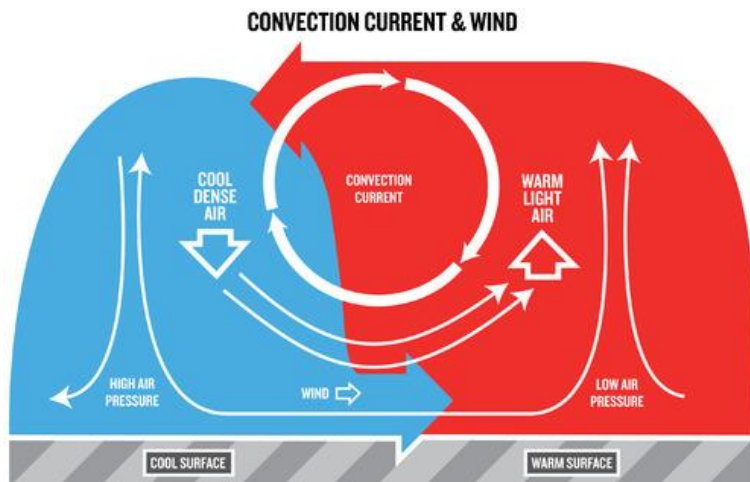
Where should I  
roast my  
marshmallow?

# Convection



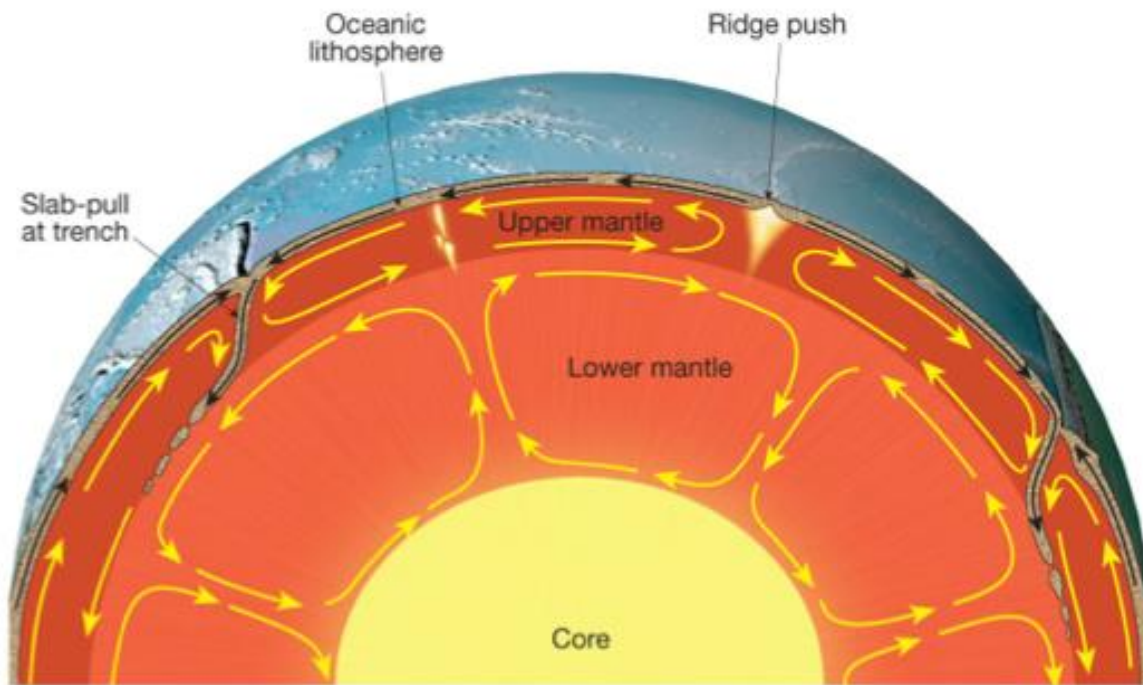
Why does hot air rise?

High Temperature  
High Volume



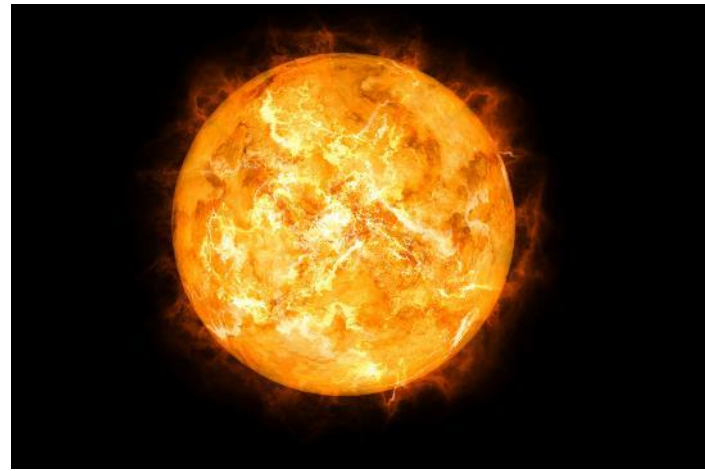
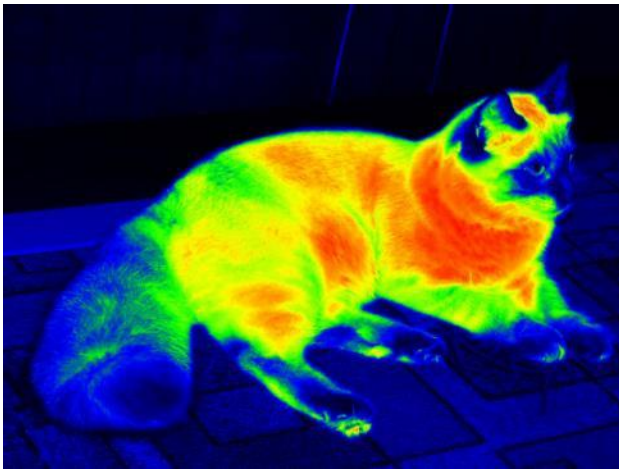
High Volume  
Same Mass  
Lower Density

# Convection



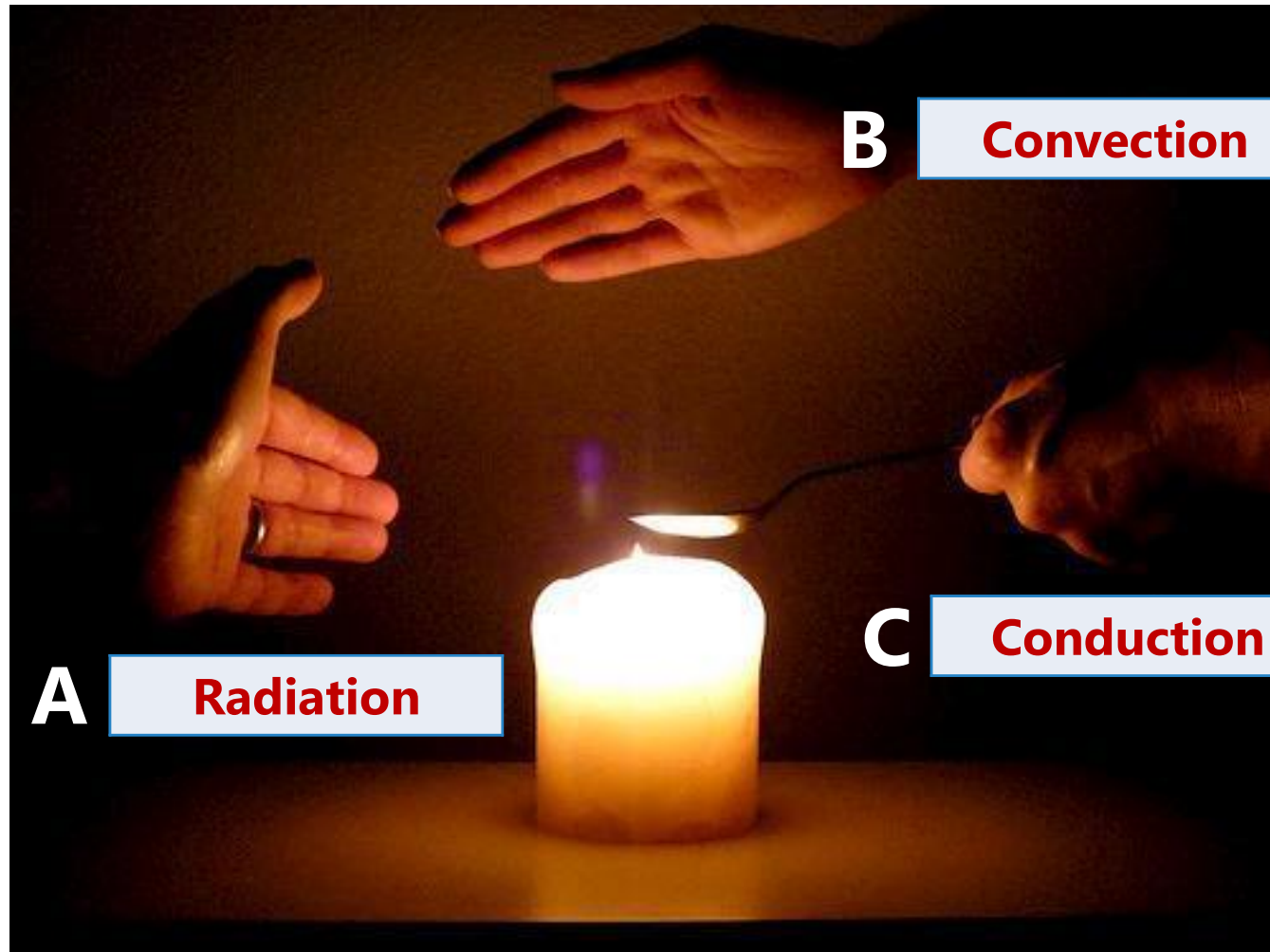
# Radiation

**Radiation** is energy that is transferred as waves such as visible light and infrared



**Radiation** can travel through           **a vacuum**

# Label Me



# Emissivity

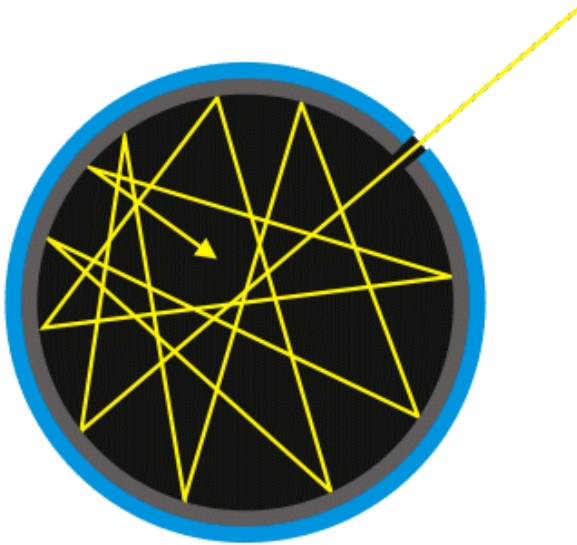
What color car heats up the most in the sun?

**Black** – Absorbs more light



# Black Body Radiator

A black body radiator is an object that is perfectly opaque and absorbs all energy



Conceptual Black Body



# Emissivity

Emissivity

$e$

$$\frac{\text{power radiated by a surface}}{\text{power radiated from a black body of the same temperature and area}}$$

The emissivity is used to adjust for an object that isn't a perfect black-body radiator

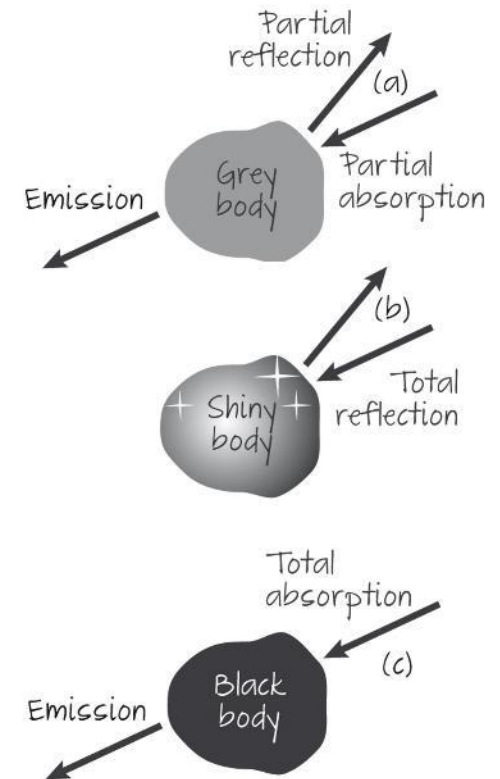
This value is always between 0 and 1



$e \approx 1$

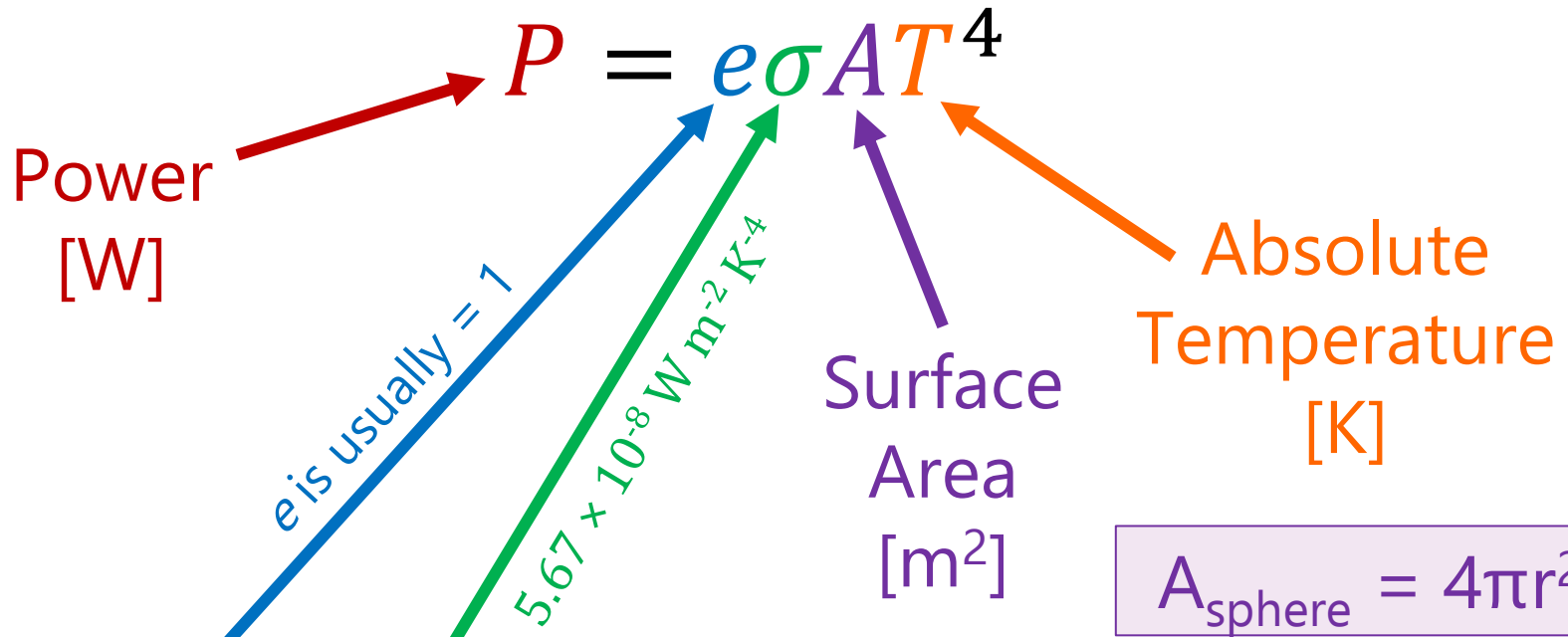


$e \approx 0.6$





# Stefan-Boltzmann Law



Emissivity	$e$	$\frac{\text{power radiated by a surface}}{\text{power radiated from a black body of the same temperature and area}}$
Stefan-Boltzmann Constant	$\sigma$	$5.67 \times 10^{-8} \text{ W m}^{-2} \text{ K}^{-4}$

# Try This

A star has a radius of  $8.3 \times 10^7$  m and a surface temperature of  $7500^\circ\text{C}$ . Calculate the power it emits.

$$e = 1$$

$$\sigma = 5.67 \times 10^{-8} \text{ W m}^{-2} \text{ K}^{-4}$$

$$A = 4\pi(8.3 \times 10^7)^2 = 8.66 \times 10^{16} \text{ m}^2$$

$$T = 7500 + 273 = 7773 \text{ K}$$

$$P = e\sigma AT^4$$

$$P = (1)(5.67 \times 10^{-8})(8.66 \times 10^{16})(7773)^4$$

$$P = 1.79 \times 10^{25} \text{ W}$$

# Proportionality

How much more heat energy is radiated from a 80°C cup of water than from a 20°C cup of water?

$$P = e\sigma AT^4$$

*\*Careful! Temperature must be converted into Kelvin*

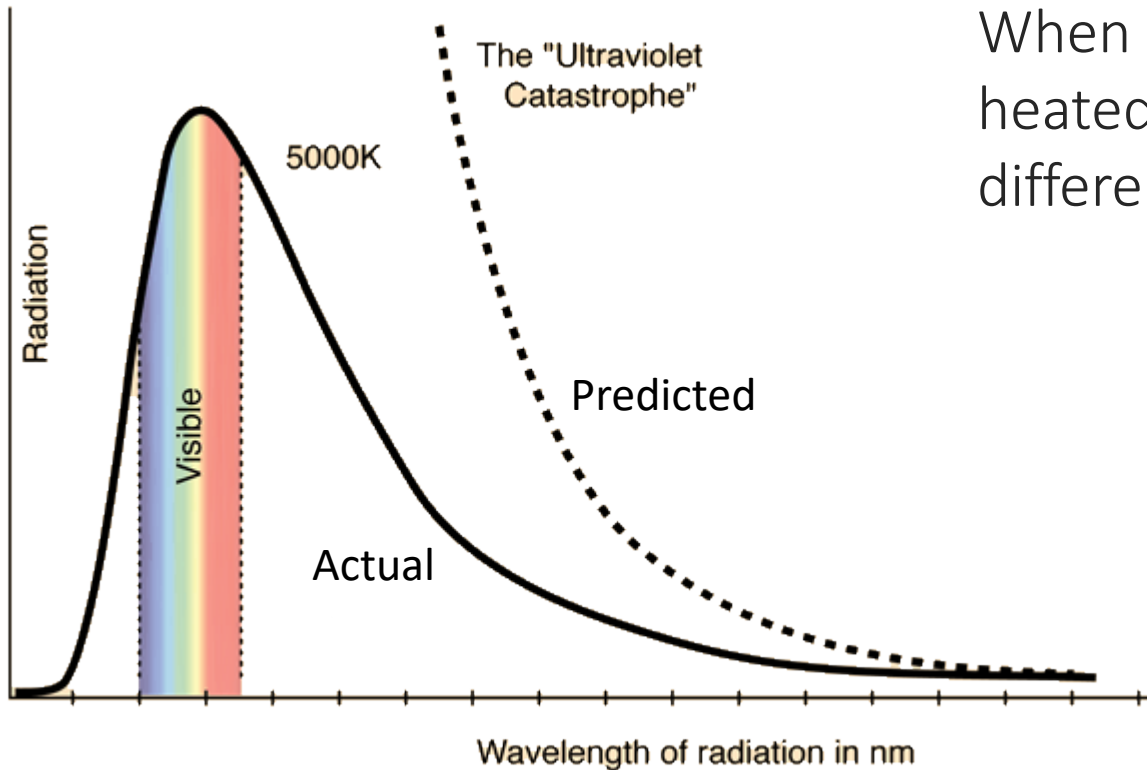
$$T_1 = 80 + 273 = 353 \text{ K}$$

$$T_2 = 20 + 273 = 293 \text{ K}$$

*e,  $\sigma$ , and A are all the same before and after...*

$$\frac{P_1}{P_2} = \frac{\cancel{e\sigma A}T_1^4}{\cancel{e\sigma A}T_2^4} = \frac{353^4}{293^4} = 2.1 \text{ times more}$$

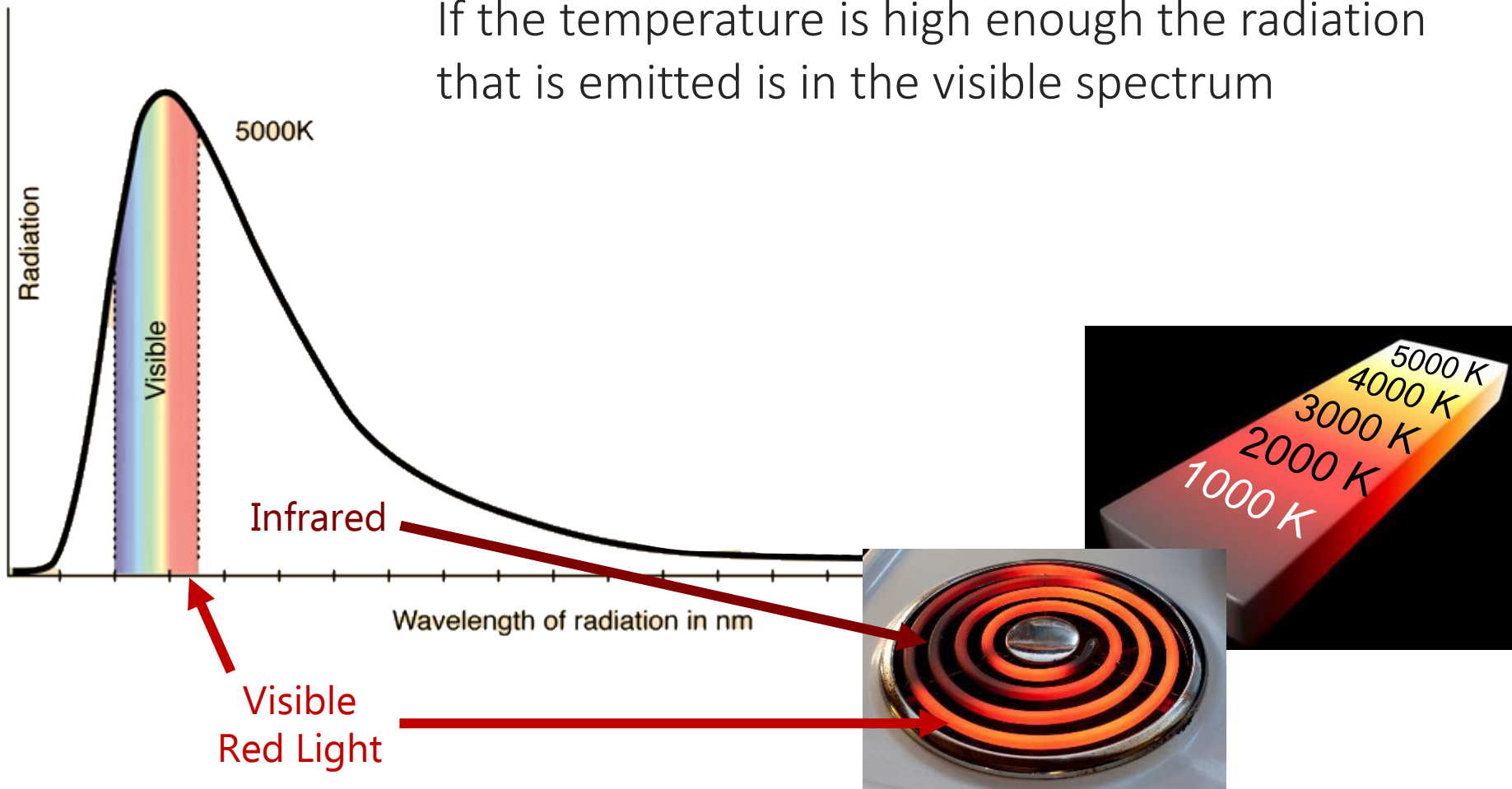
# Radiated Energy



When a black body radiator is heated up, it emits a range of different wavelengths

# Glowing Hot

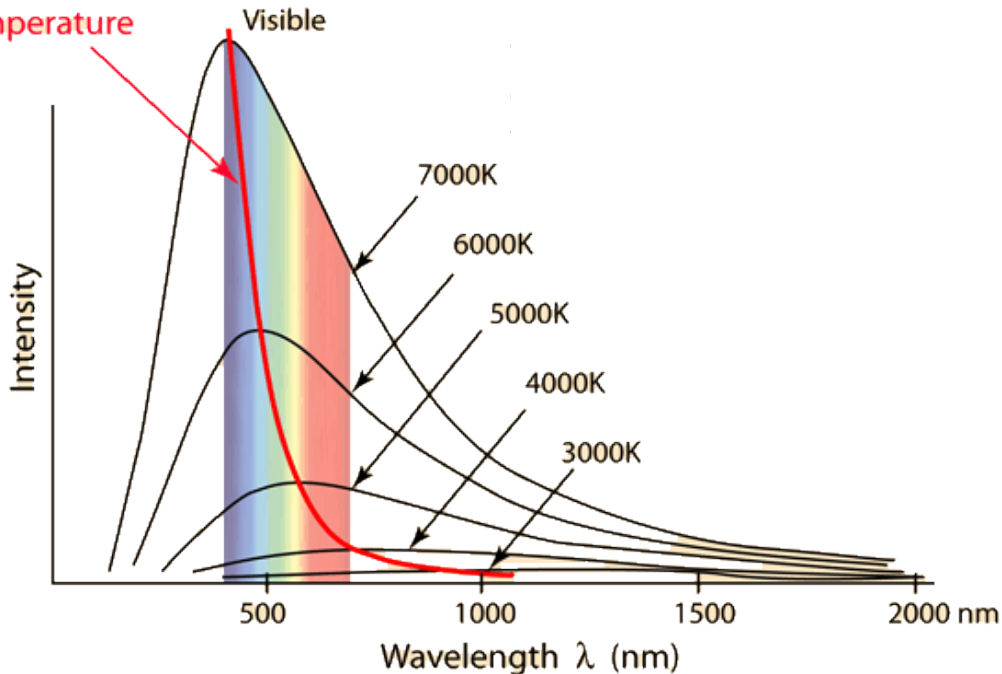
If the temperature is high enough the radiation that is emitted is in the visible spectrum



# Wien's Displacement Law

$$\lambda_{\max}(\text{metres}) = \frac{2.90 \times 10^{-3}}{T(\text{kelvin})}$$

Decrease of  $\lambda_{\text{peak}}$   
with increase in  
temperature



\*Note: This assumes perfect blackbody radiation

# Try This

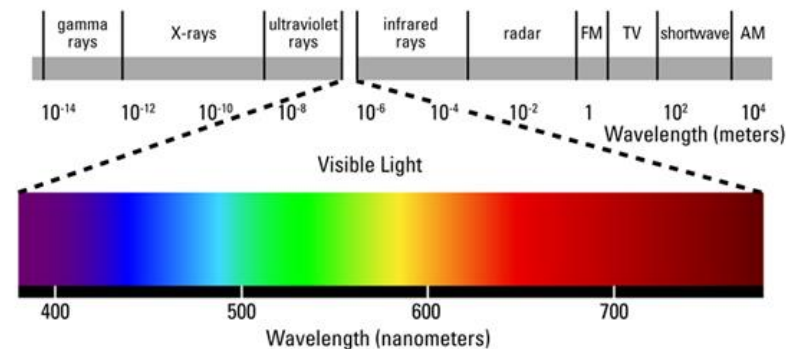
$$\lambda_{\max}(\text{metres}) = \frac{2.90 \times 10^{-3}}{T (\text{kelvin})}$$

At what wavelength is the emitted radiation of the Sun maximized if it has a surface temperature of 5780 K?

$$\lambda = \frac{2.90 \times 10^{-3}}{5780} = 5.02 \times 10^{-7} \text{ m} = \boxed{502 \text{ nm}}$$

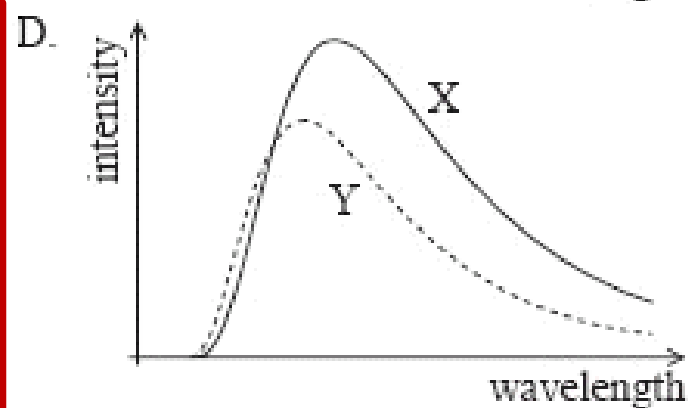
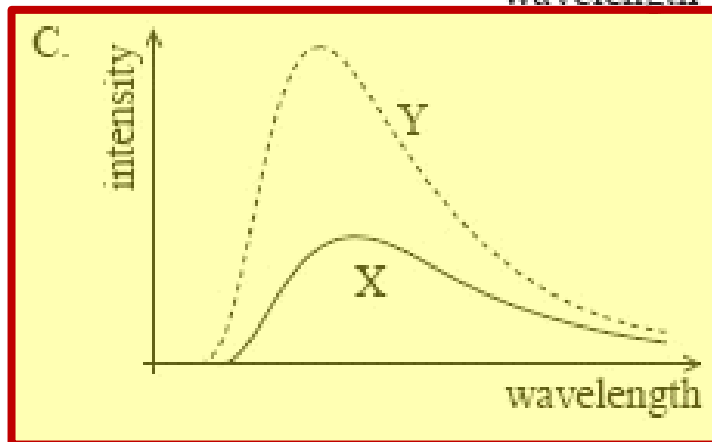
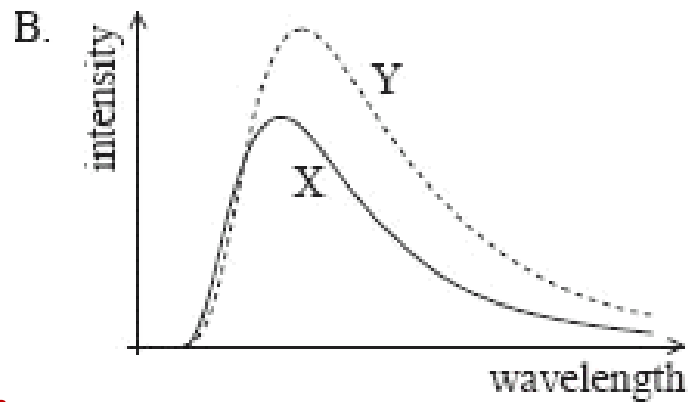
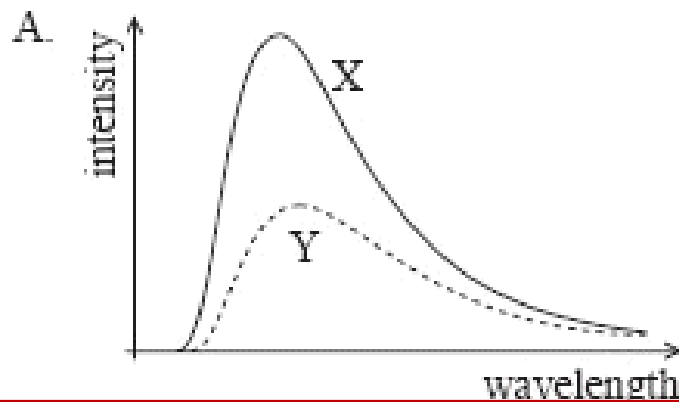
What is the most prevalent color of sunlight?

Green



# Sample IB Question

Two black bodies X and Y are at different temperatures. The temperature of body Y is higher than that of body X. Which of the following shows the black body spectra for the two bodies?





# Takeaways from Today

Know the difference between:

- Conduction
- Convection
- Radiation

Black Body Radiators

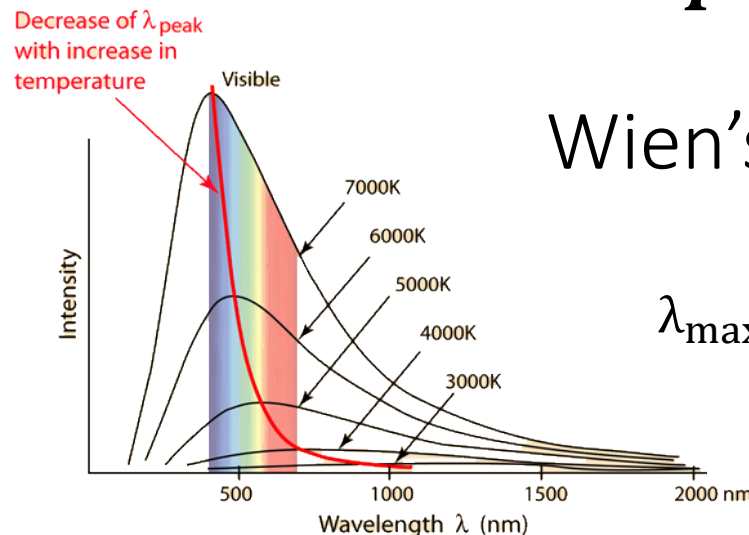
Emissivity

Stefan-Boltzmann Law

$$P = e\sigma AT^4$$

Wien's Displacement Law

$$\lambda_{\max}(\text{metres}) = \frac{2.90 \times 10^{-3}}{T(\text{kelvin})}$$



# Radiation from the Sun

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# Intensity

$$\textit{Intensity} = \frac{\textit{Power}}{A}$$

$$\textit{Units} = \frac{W}{m^2} = W m^{-2}$$

# Intensity

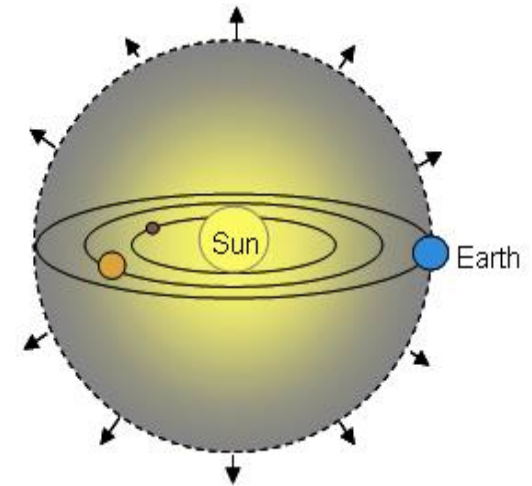
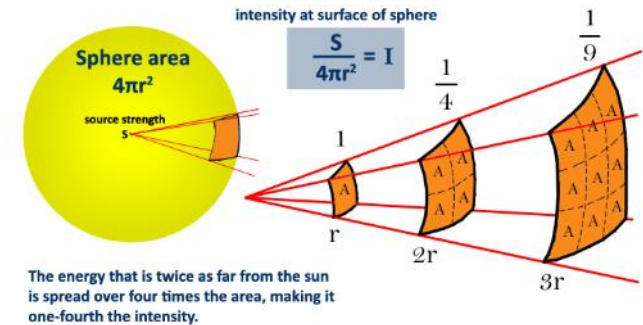
Calculate the intensity of the Sun's radiation arriving to Earth

Sun's Power =  $3.84 \times 10^{26}$  W

Earth's Distance from Sun =  $150 \times 10^6$  km

$$I = \frac{P}{A} = \frac{3.84 \times 10^{26}}{4\pi(150 \times 10^9)^2}$$

$$I = 1358 \text{ Wm}^{-2}$$



# Solar Constant

The average intensity falling on an area above the earth's atmosphere perpendicular to the direction traveled by the radiation

$$S = 1360 \text{ W m}^{-2} = 1.36 \times 10^3 \text{ W m}^{-2}$$

Quantity	Symbol	Approximate value
Elementary charge	$e$	$1.60 \times 10^{-19} \text{ C}$
Electron rest mass	$m_e$	$9.110 \times 10^{-31} \text{ kg} = 0.000549 \text{ u} = 0.511 \text{ MeV c}^{-2}$
Proton rest mass	$m_p$	$1.673 \times 10^{-27} \text{ kg} = 1.007276 \text{ u} = 938 \text{ MeV c}^{-2}$
Neutron rest mass	$m_n$	$1.675 \times 10^{-27} \text{ kg} = 1.008665 \text{ u} = 940 \text{ MeV c}^{-2}$
Unified atomic mass unit	$u$	$1.661 \times 10^{-27} \text{ kg} = 931.5 \text{ MeV c}^{-2}$
Solar constant	$S$	$1.36 \times 10^3 \text{ W m}^{-2}$
Fermi radius	$R_0$	$1.20 \times 10^{-15} \text{ m}$

# Average Solar Intensity on Earth

Earth's Radius =  $6.37 \times 10^6$  m

Area of sun power captured:

$$\pi r^2 = \pi(6.37 \times 10^6)^2 =$$

$$1.27 \times 10^{14} \text{ m}^2$$

1360 W m<sup>-2</sup>



Total sun power captured:

$$1.27 \times 10^{14} \text{ m}^2 \times \frac{1360 \text{ W}}{1 \text{ m}^2} =$$

$$1.7 \times 10^{17} \text{ W}$$

Total Power Received by the Earth

$$1.7 \times 10^{17} \text{ W}$$

Average spread out across Earth's surface:

$$\frac{P}{A_{sphere}} = \frac{1.7 \times 10^{17} \text{ W}}{4\pi(6.37 \times 10^6)^2} =$$

$$340 \text{ W m}^2$$

Average Solar Intensity on Earth

$$340 \text{ W m}^2$$

# Albedo vs. Emissivity

Albedo

$$\frac{\text{power scattered by a body}}{\text{incident power}}$$

% Reflected

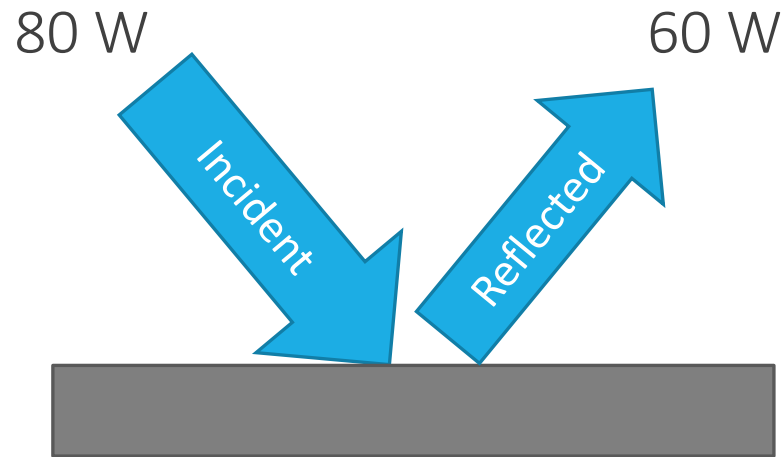
Emissivity

$e$

$$\frac{\text{power radiated by a surface}}{\text{power radiated from a black body}}$$

% Absorbed

# Albedo vs. Emissivity



Albedo

$$\frac{60}{80} = 0.75$$

Emissivity

$$\frac{20}{80} = 0.25$$



# Albedo of Earth



$$\textit{Albedo} = \frac{102}{340} = 0.3$$

# Albedo of Earth

Highest Albedo?

**0.66**

**Snow**

Lowest Albedo?

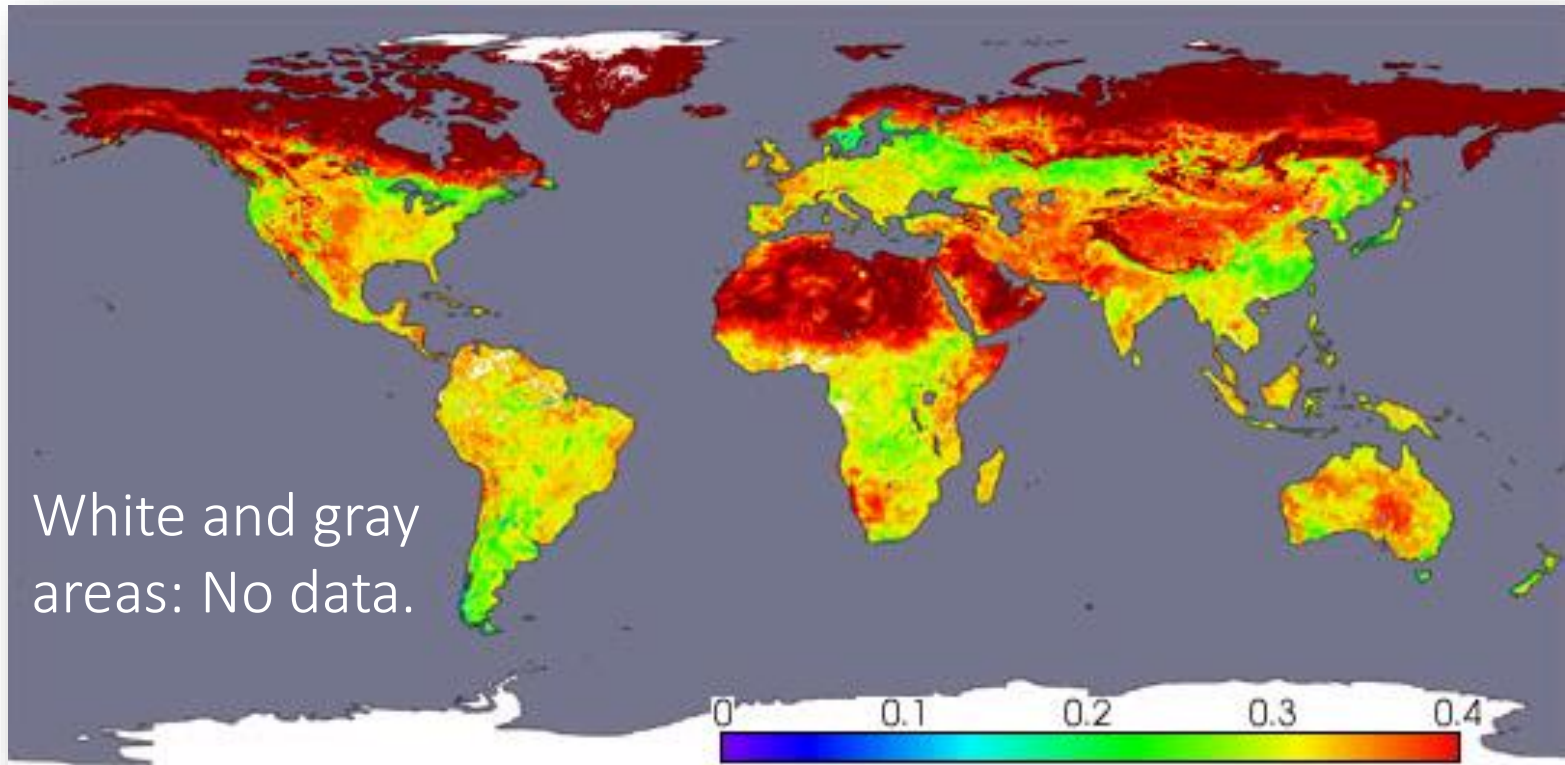
**0.07**

**Ocean**

Surface	Albedo <sup>a</sup>	
	Summer	Winter
Mixed farming, tall grass	0.16	0.18
Tall/medium grassland, evergreen shrubland	0.20	0.21
Short, grassland, meadow and shrubland	0.21	0.20
Evergreen forest (needle leaved)	0.12	0.13
Mixed deciduous, evergreen forest	0.16	0.16
Deciduous forest	0.17	0.18
Tropical evergreen broadleaved forest	0.12	0.15
Medium/tall grassland, woodland	0.15	0.18
Desert	0.36	0.36
Tundra	0.17	0.17
Snow	0.66	0.66
Sea ice	0.62	0.62
Ocean	0.07	0.07

Data taken from Briegleb *et al.* (1986).

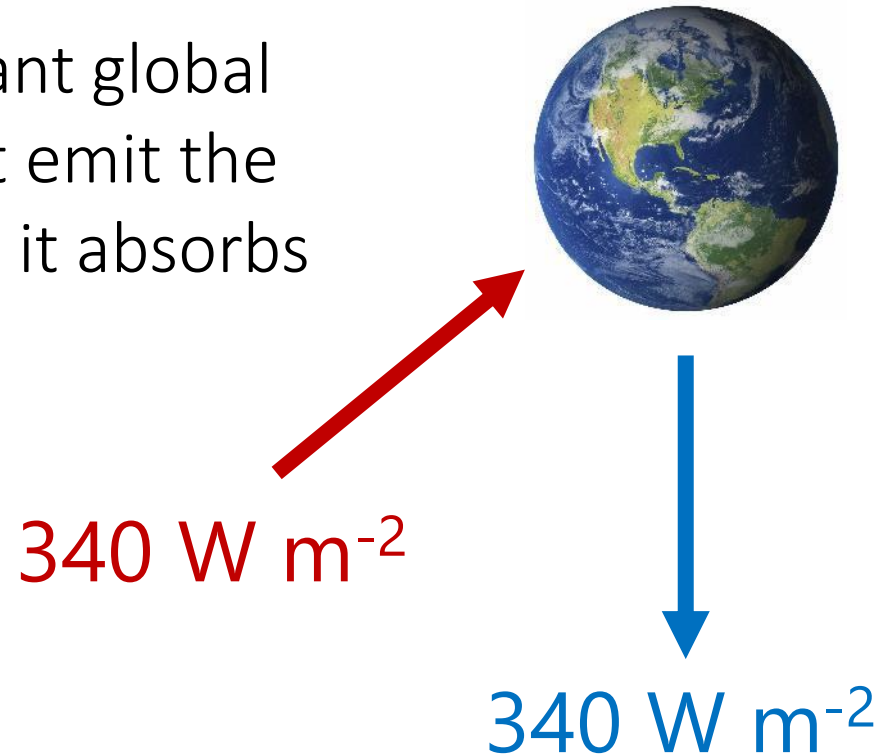
# Albedo of Earth



April, 2002, *Terra* satellite, NASA

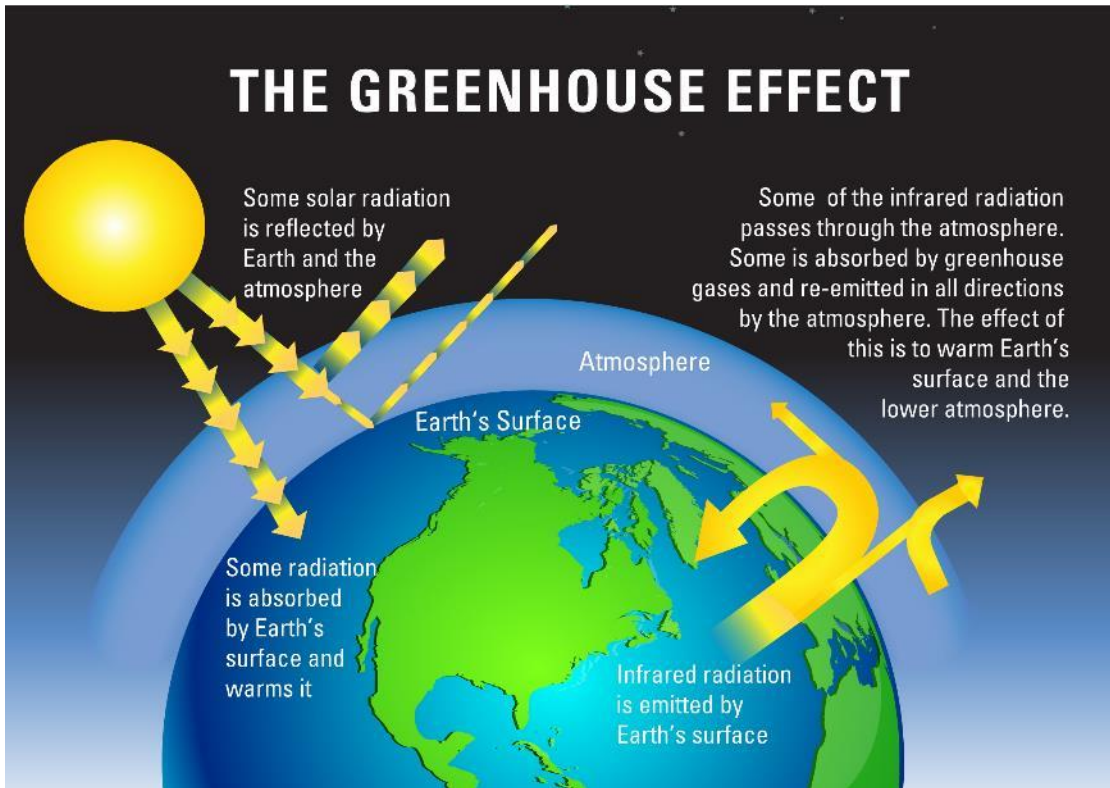
# Thermal Equilibrium

In order to maintain a constant global temperature, the Earth must emit the same amount of energy that it absorbs

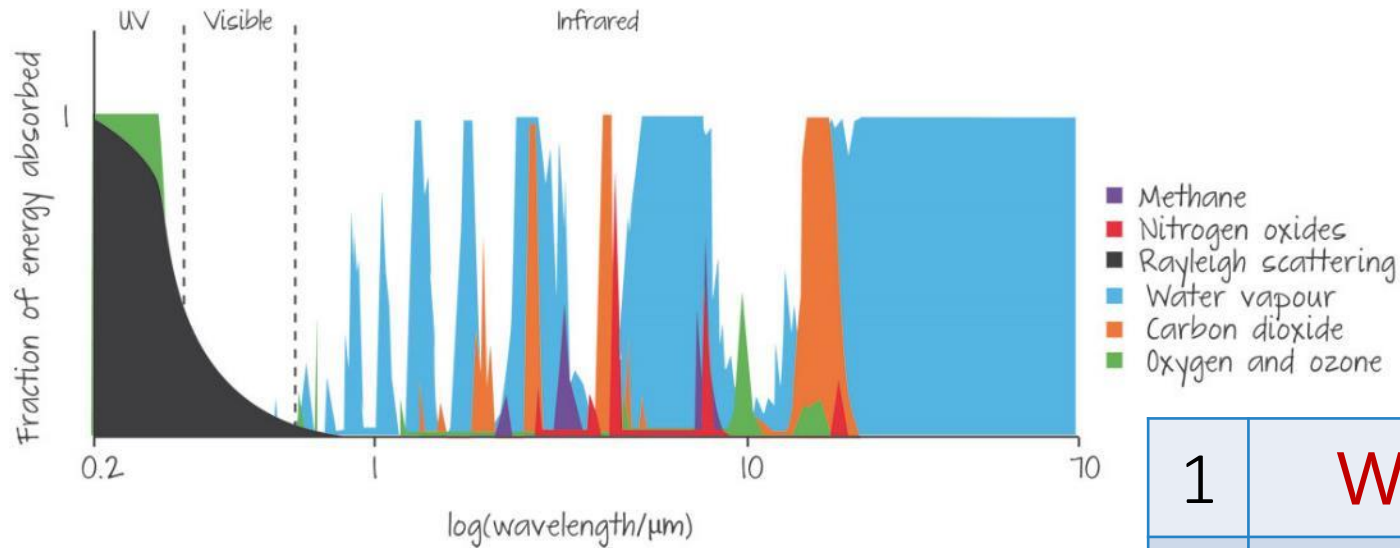


# Greenhouse Effect

If there was no atmosphere, the earth would experience a net loss of energy and reach equilibrium at an average temperature about 30°C colder than it is currently.



# Role of the Atmosphere



Rank the following Greenhouse Gases based on the amount of infrared energy they absorb



1	<b>Water Vapor</b>
2	<b>Carbon Dioxide</b>
3	<b>Methane</b>
4	<b>Nitrogen Oxides</b>
5	<b>Oxygen/Ozone</b>

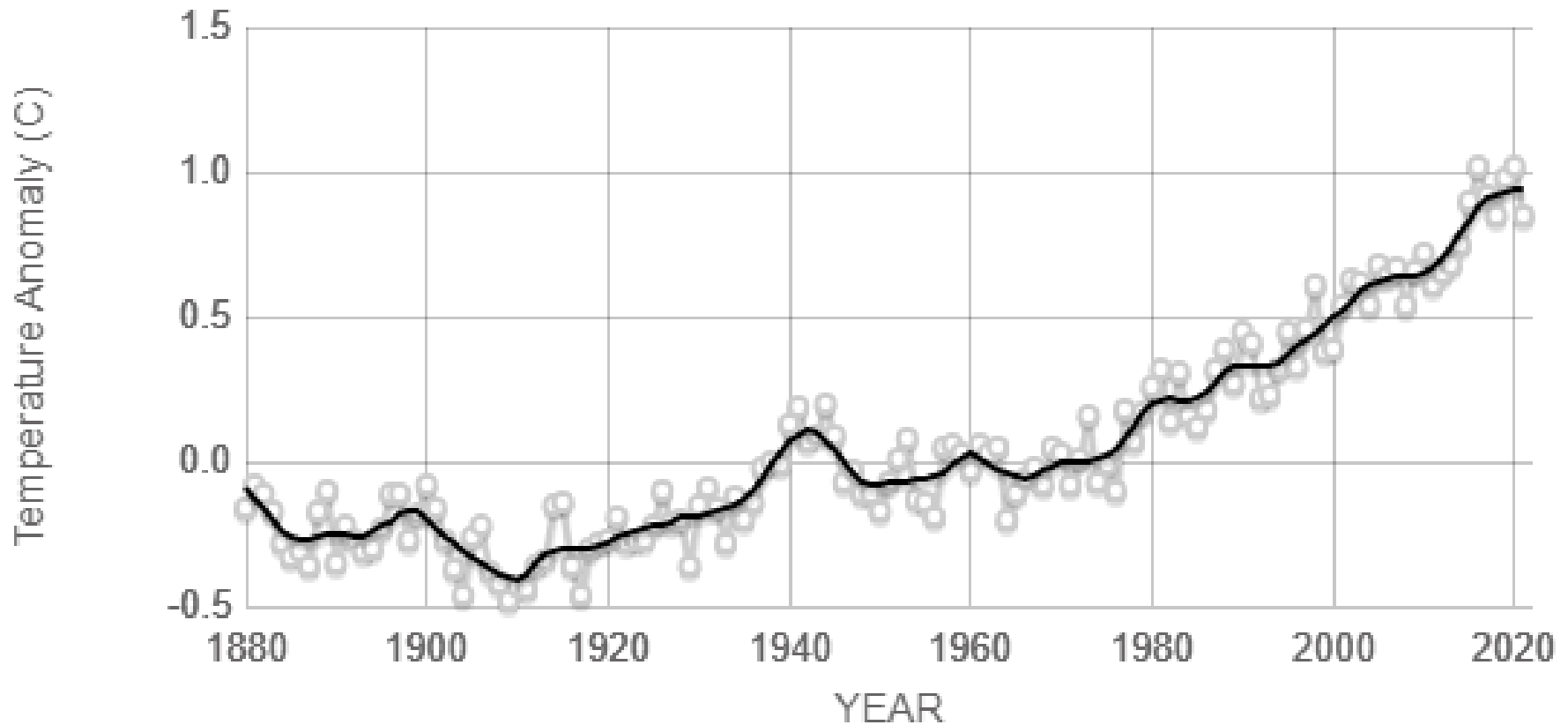
More on this later...

# Climate Change

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IB PHYSICS | ENERGY PRODUCTION

# Temperature has been Rising



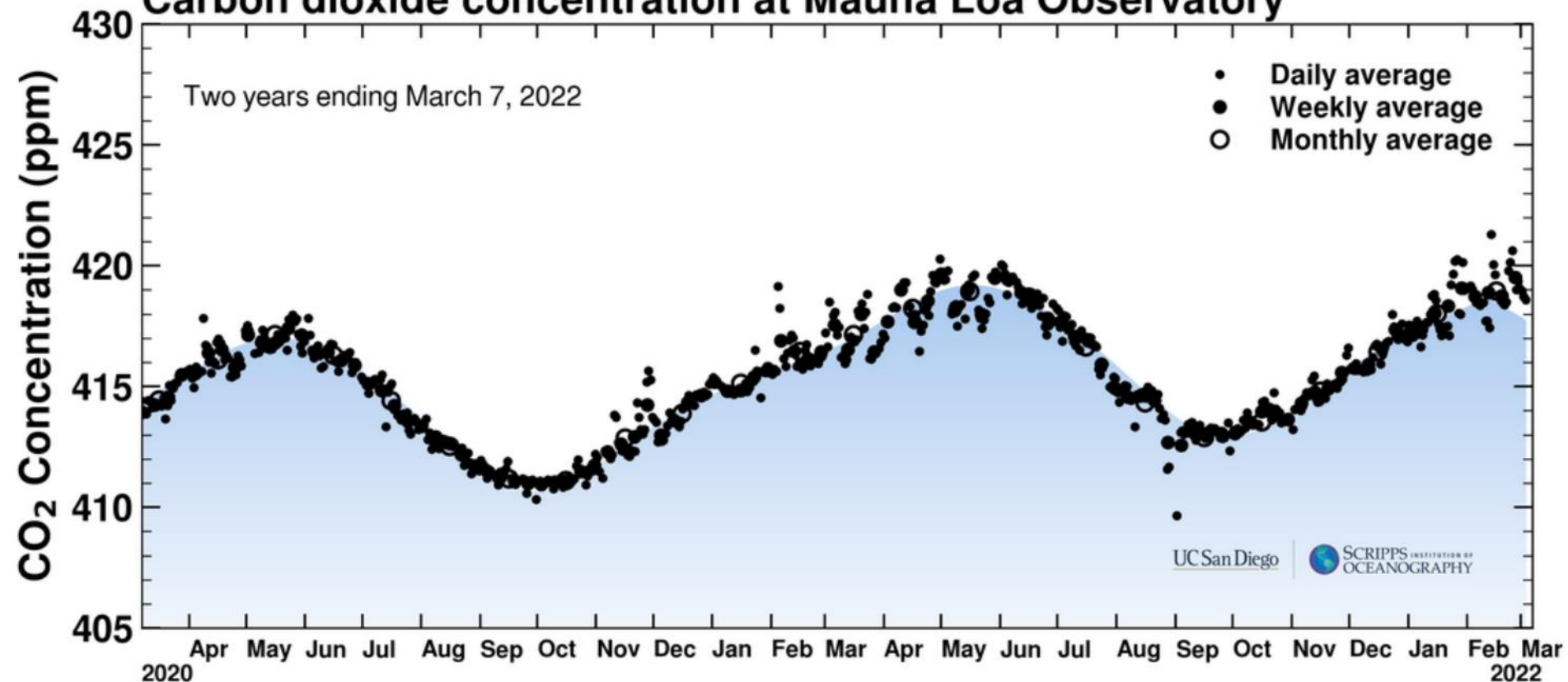
Source: [climate.nasa.gov](http://climate.nasa.gov)



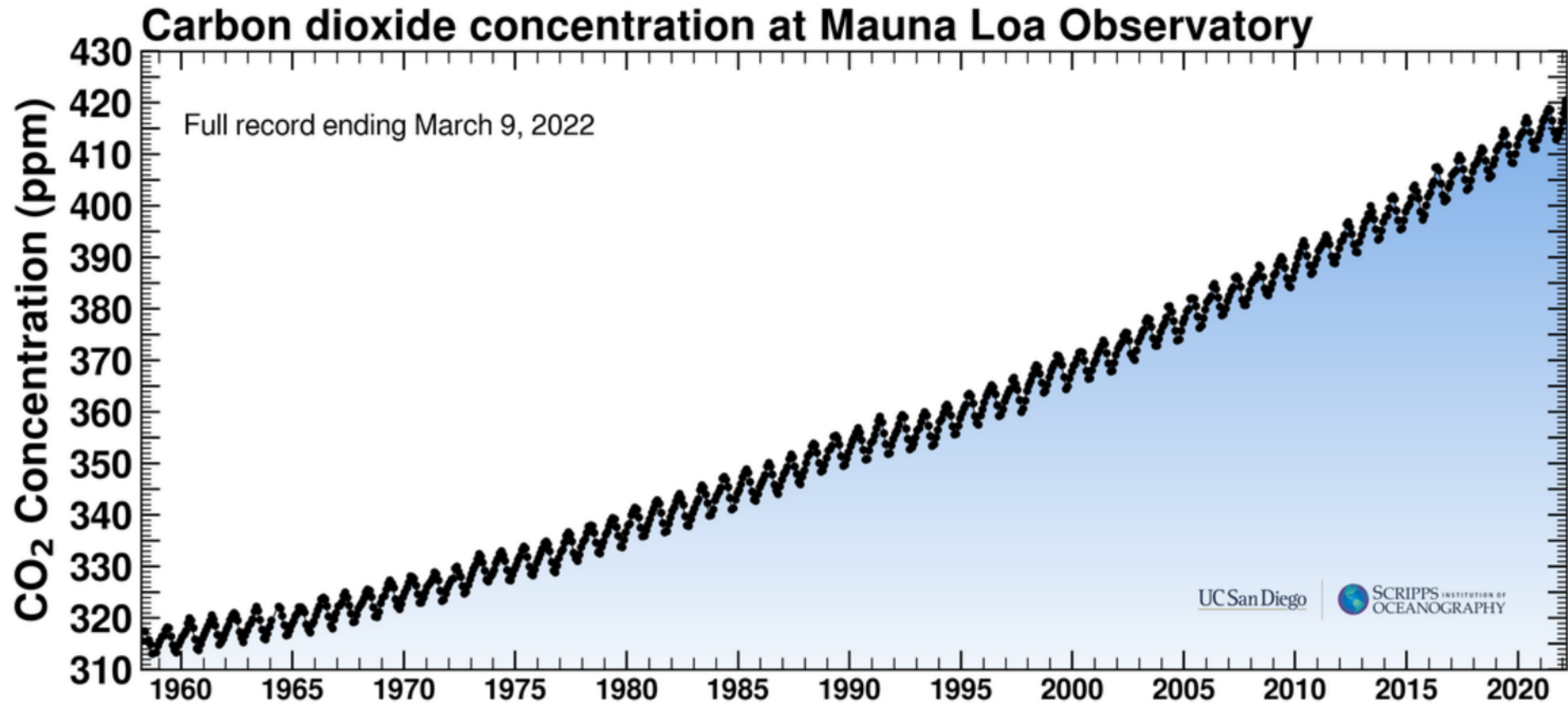
# CO<sub>2</sub> Concentration | 2 years

Latest CO<sub>2</sub> Reading: **417.88 ppm**

Carbon dioxide concentration at Mauna Loa Observatory

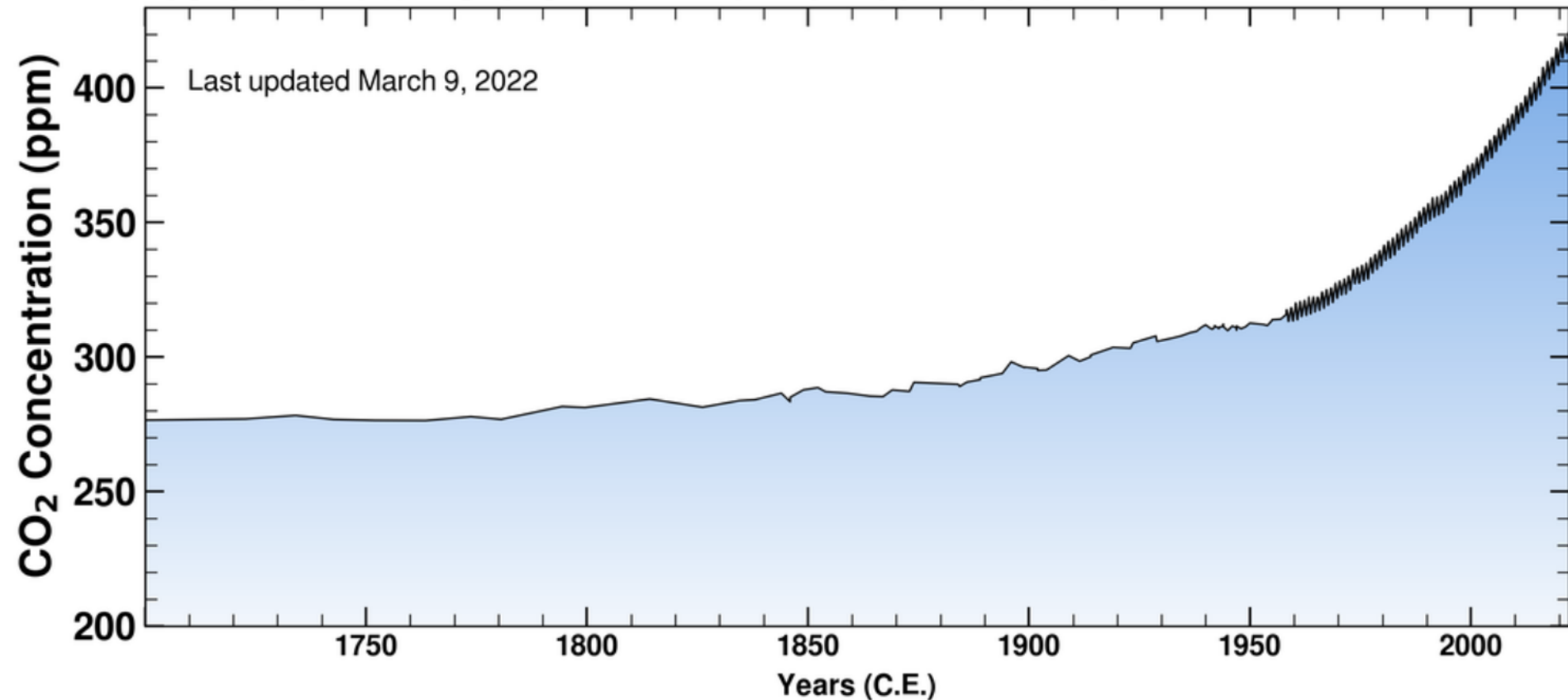


# CO<sub>2</sub> Concentration | 63 years



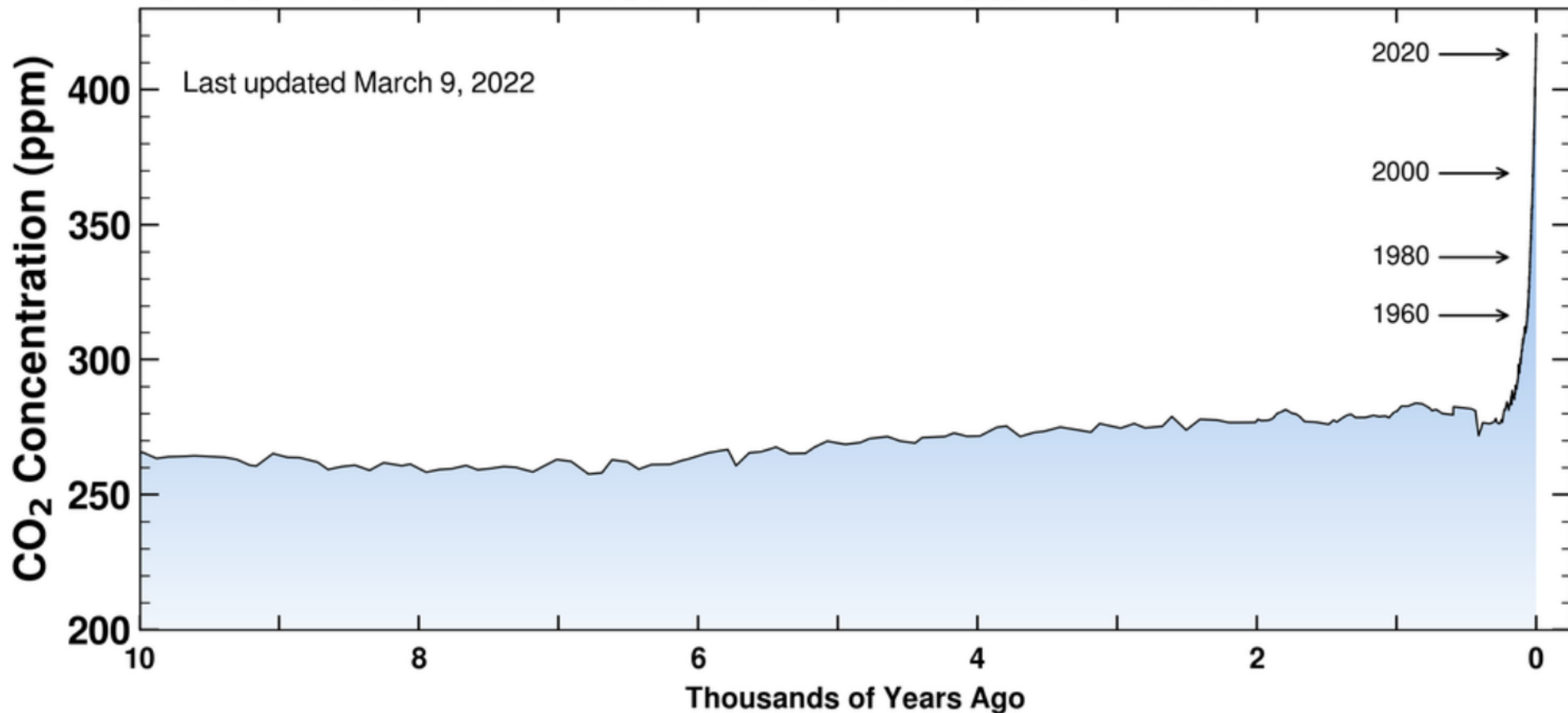
# CO<sub>2</sub> Concentration | 300 years

**Ice-core data before 1958. Mauna Loa Data after 1958.**



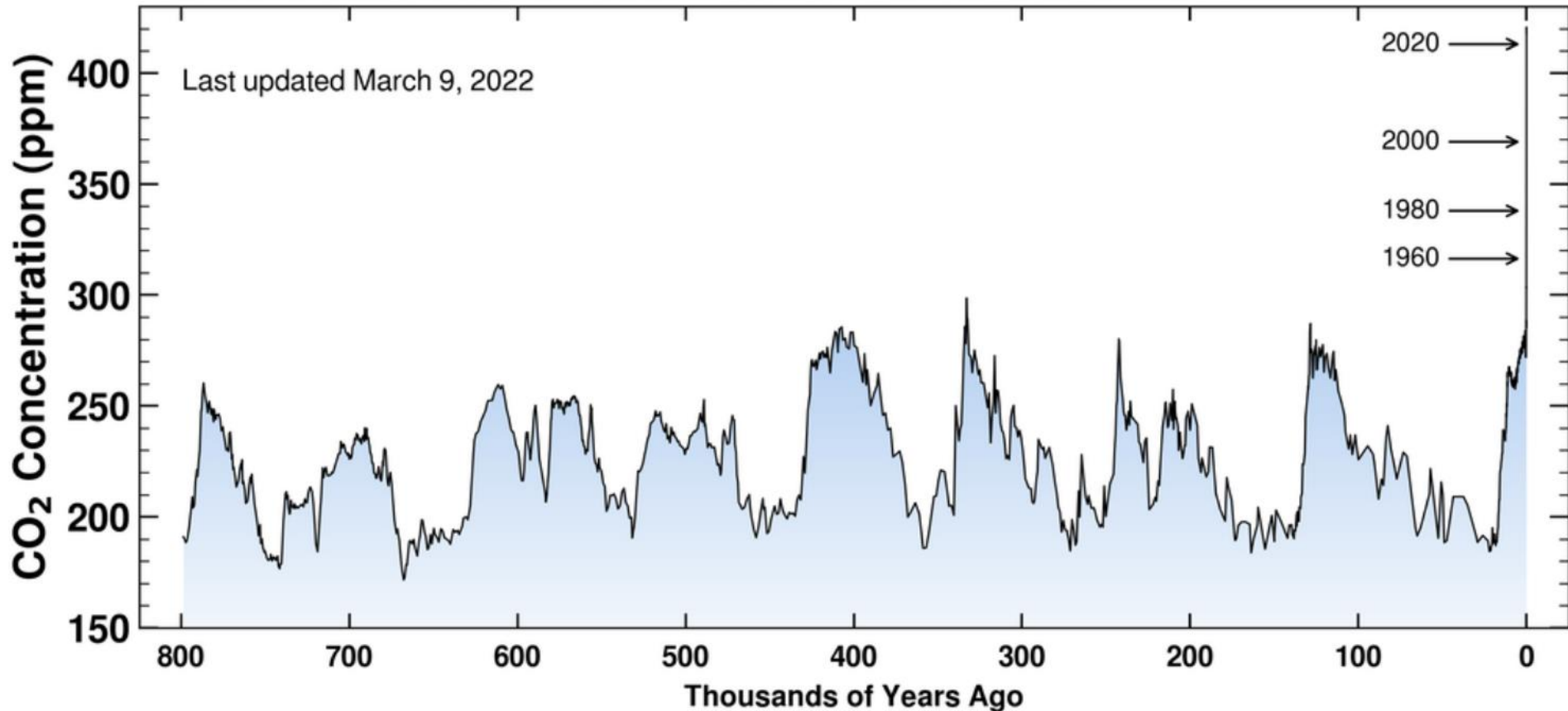
# CO<sub>2</sub> Concentration | 10,000 years

Ice-core data before 1958. Mauna Loa Data after 1958.

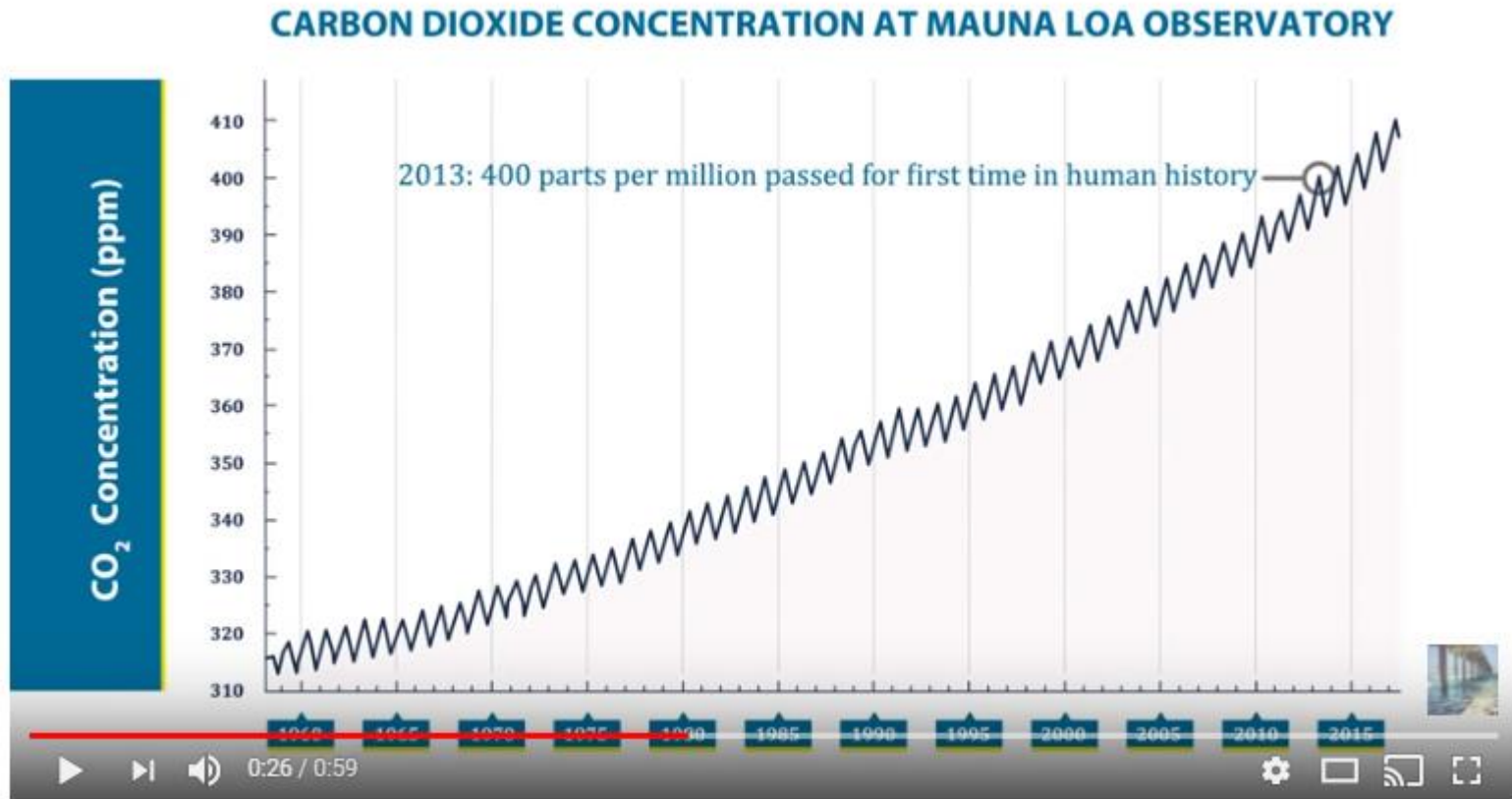


# CO<sub>2</sub> Concentration | 800,000 years

Ice-core data before 1958. Mauna Loa Data after 1958.

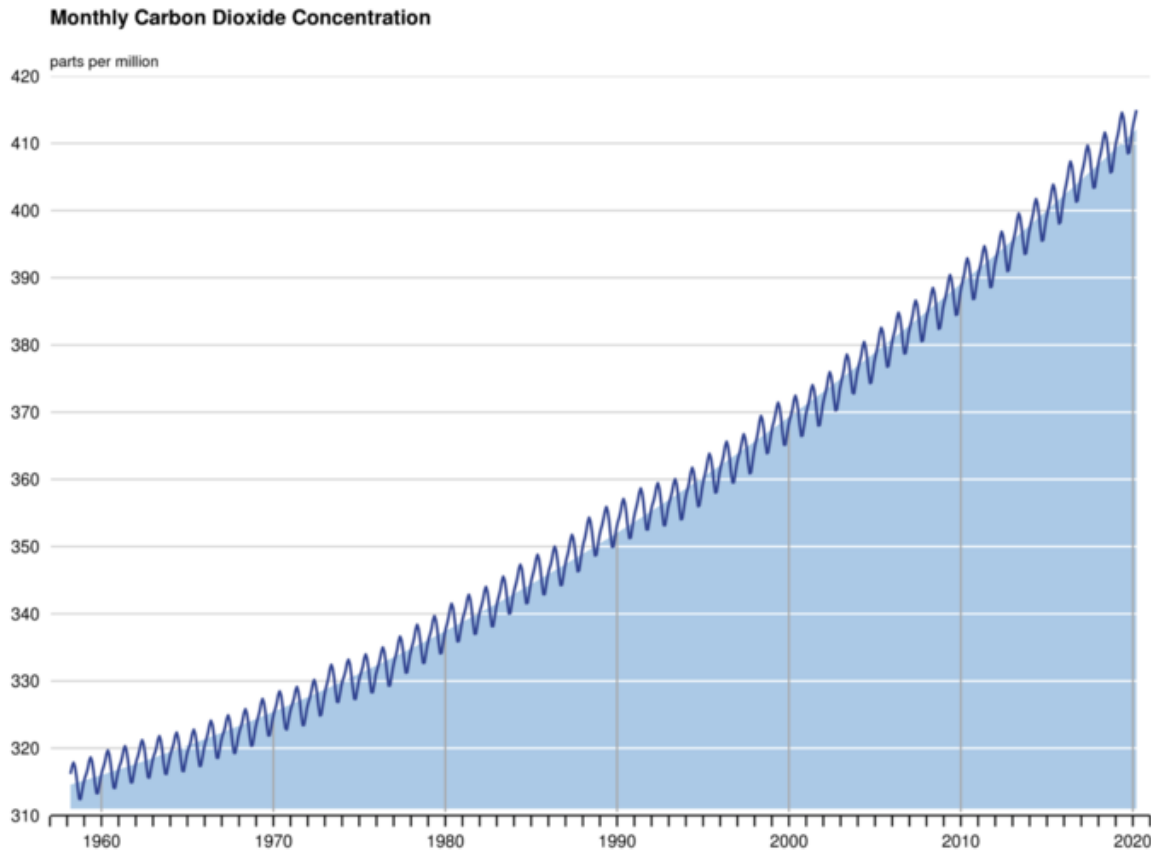


# CO<sub>2</sub> Concentration



The Keeling Curve animation

# Why does the level fluctuate yearly?



## Seasons

*\*There are more land (plants) in the Northern Hemisphere that remove CO<sub>2</sub> from the atmosphere during the summer months*



# The Greenhouse Effect

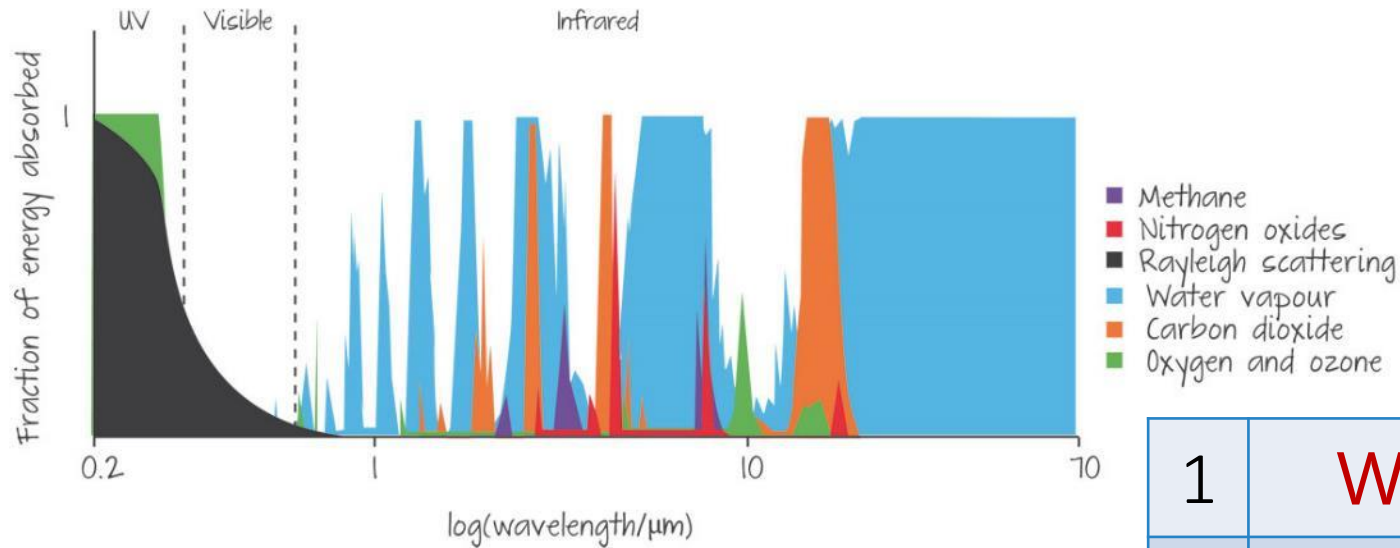


Gas molecules absorb and reemit infrared radiation

*\*This happens because the shape of these molecules means that they have natural vibration frequency that matches the frequency of infrared waves*



# The Top Greenhouse Gases

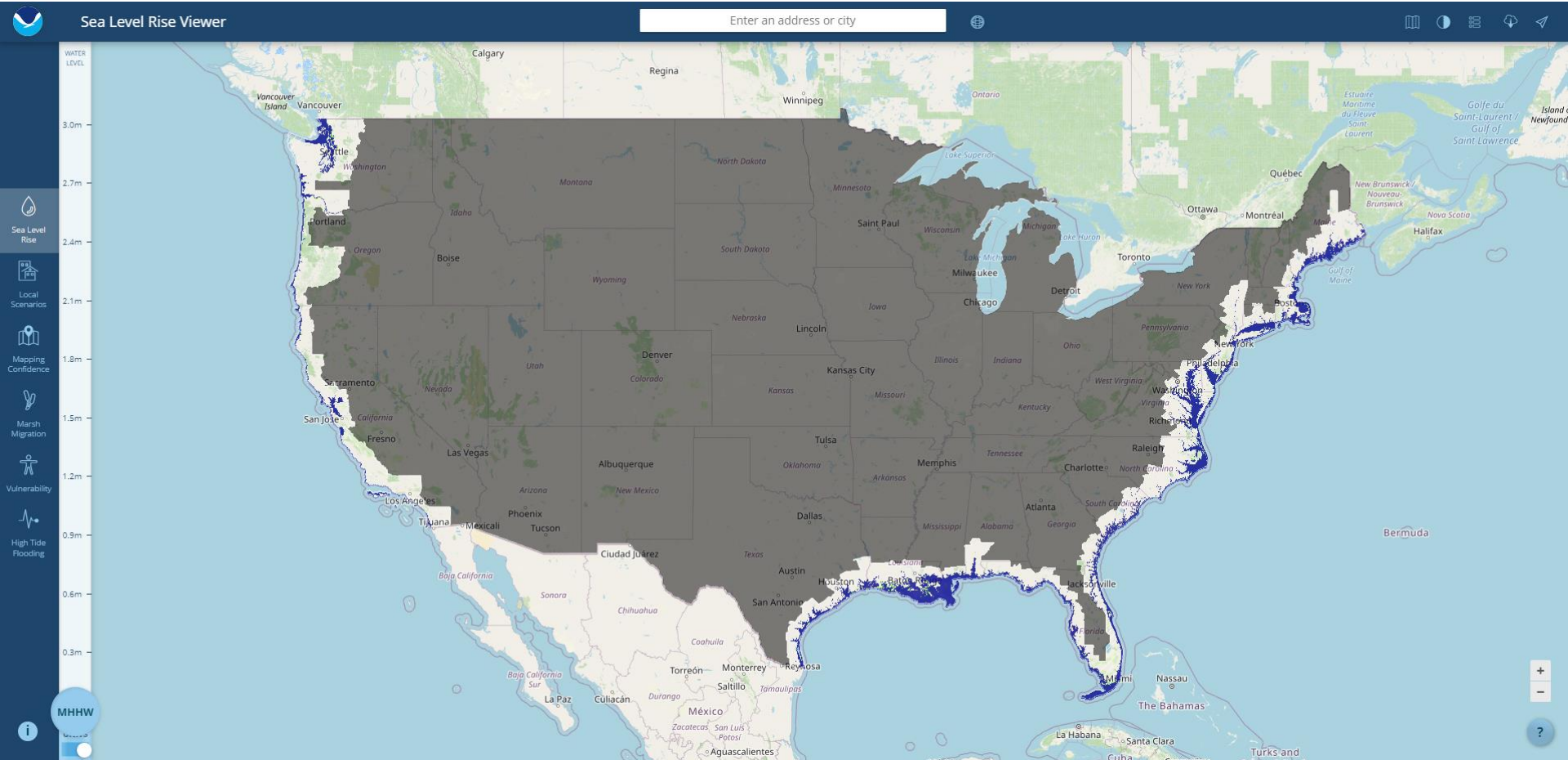


Rank the following Greenhouse Gases based on the amount of infrared energy they absorb



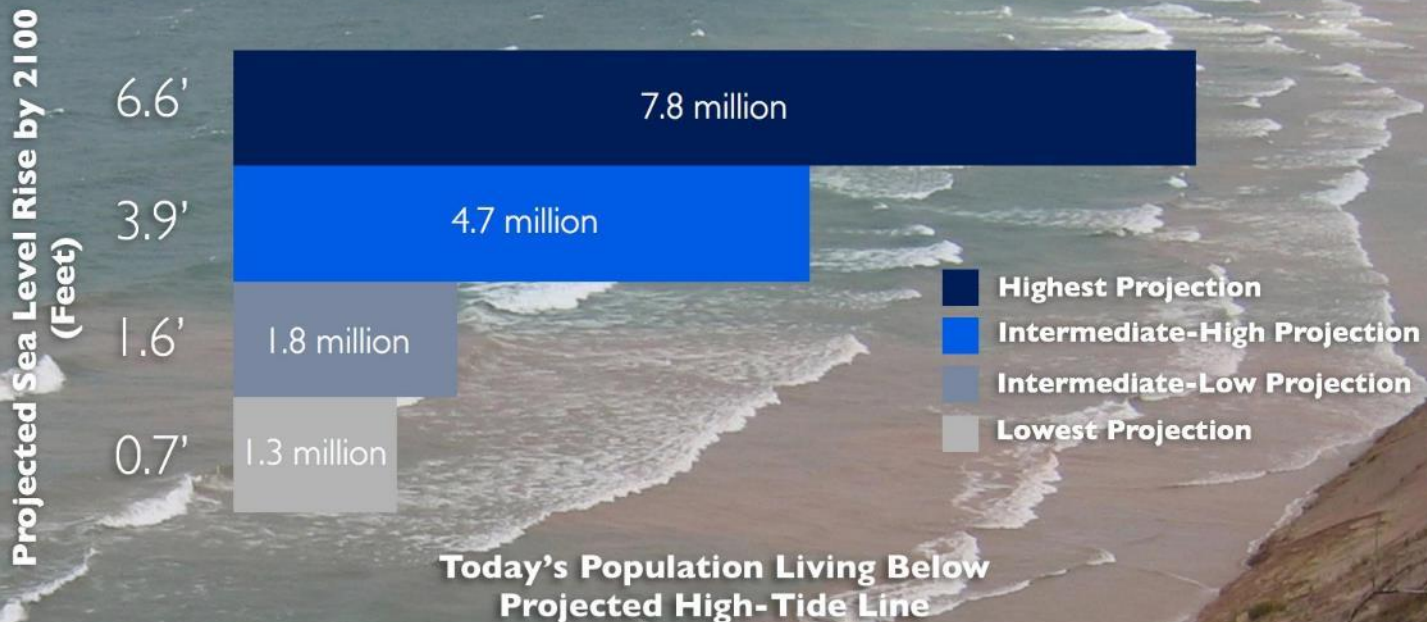
1	<b>Water Vapor</b>
2	<b>Carbon Dioxide</b>
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4	Nitrogen Oxides
5	Oxygen/Ozone

# Impacts of Climate Change



# Impacts of Climate Change

## Sea Level Rise & Population Impact





# Sea Levels Rising | Melting Ice



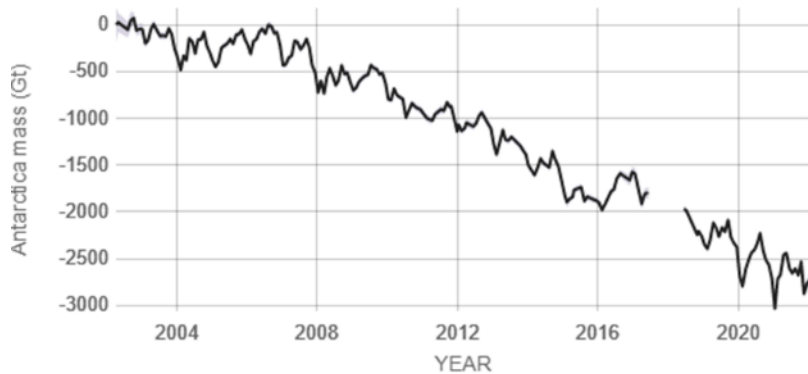
A melting iceberg does not cause a direct change in sea level



A melting glacier adds water to the ocean and causes a direct change in sea level

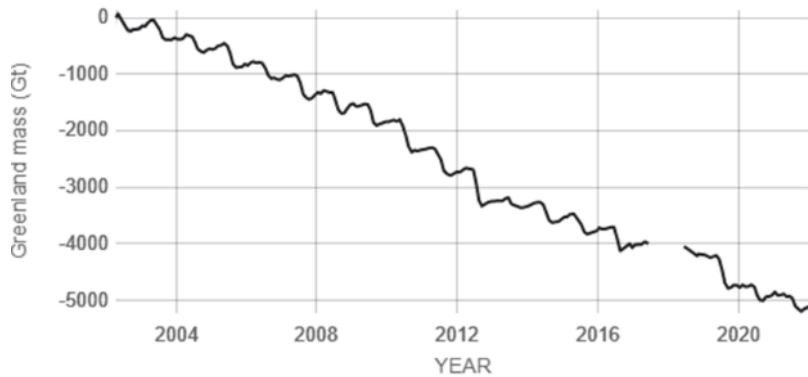
## Why?

# Sea Levels Rising | Melting Ice



Source: climate.nasa.gov

Antarctica ice mass is decreasing at a rate of 1521 billion tons per year



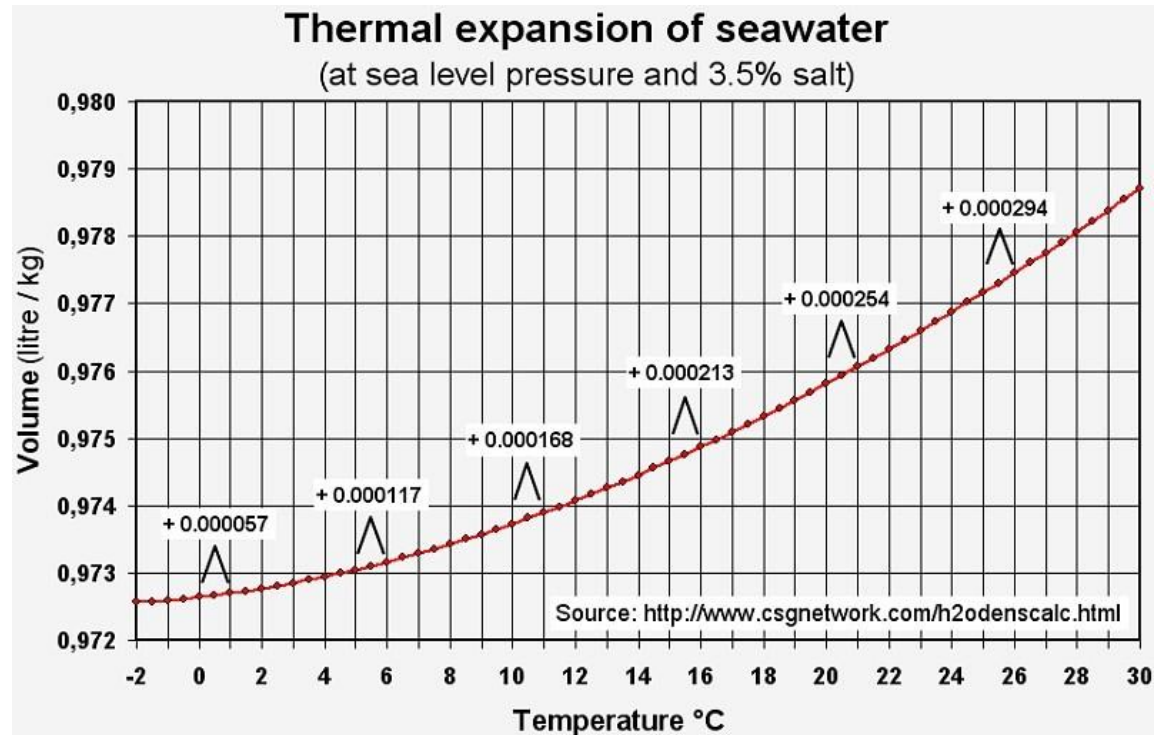
Source: climate.nasa.gov

Greenland ice mass is decreasing at a rate of 275 billion tons per year

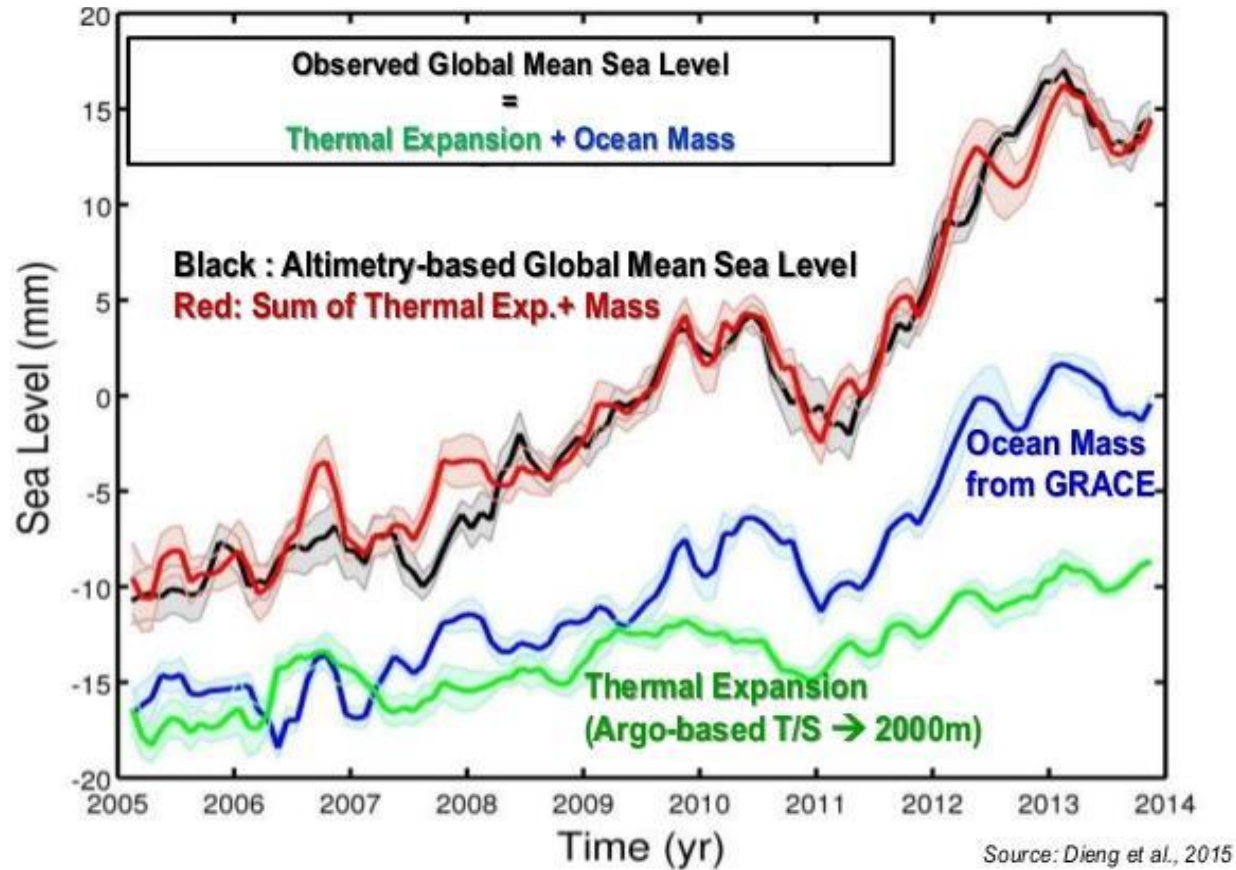
# Sea Levels Rising | Expansion



When most objects are heated they expand. Water is no different.



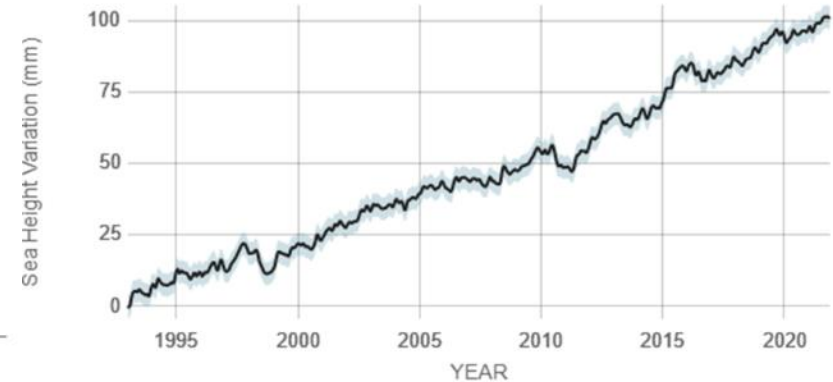
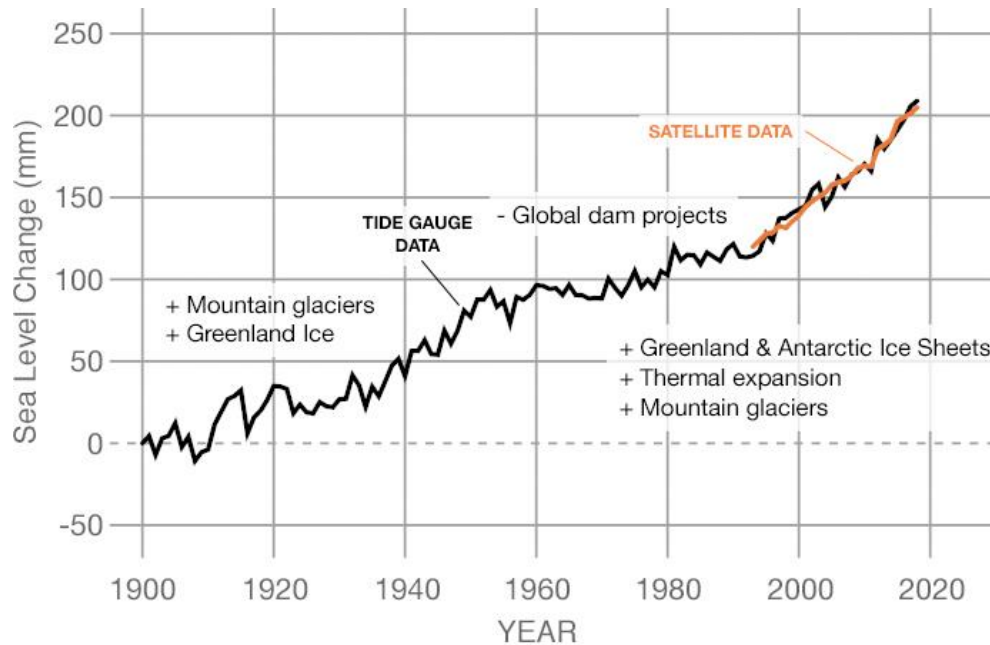
# Sea Levels Rising



# Ground Based Sea Level | 1900-Present

RATE OF CHANGE

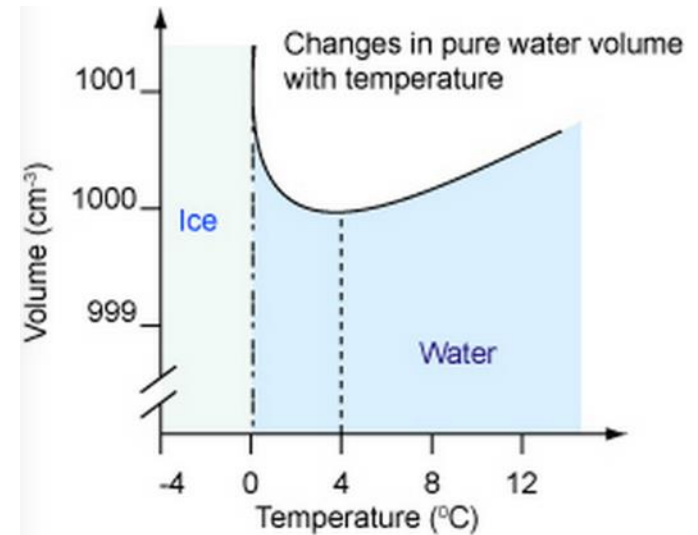
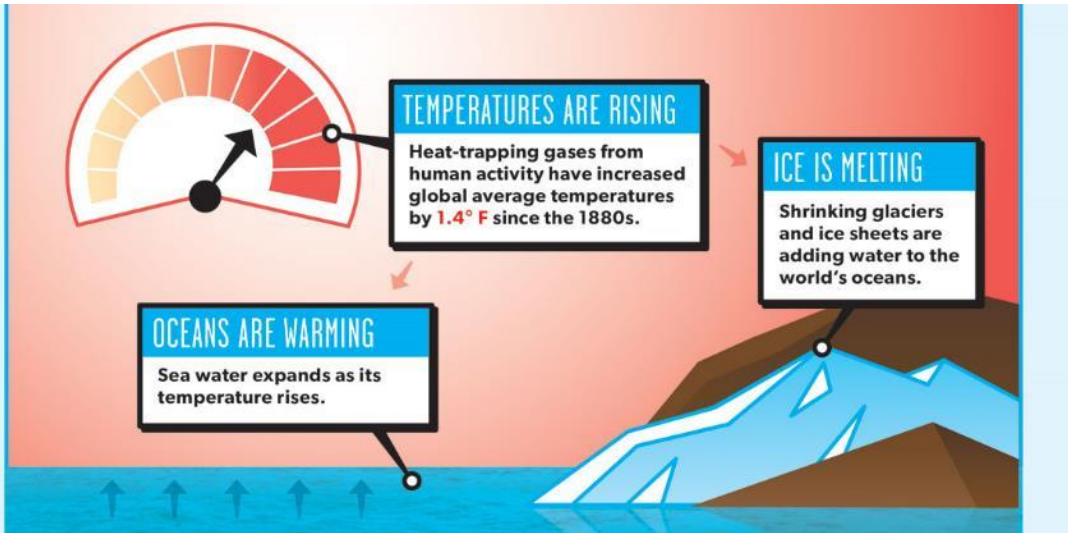
↑ 3.3  
millimeters per year



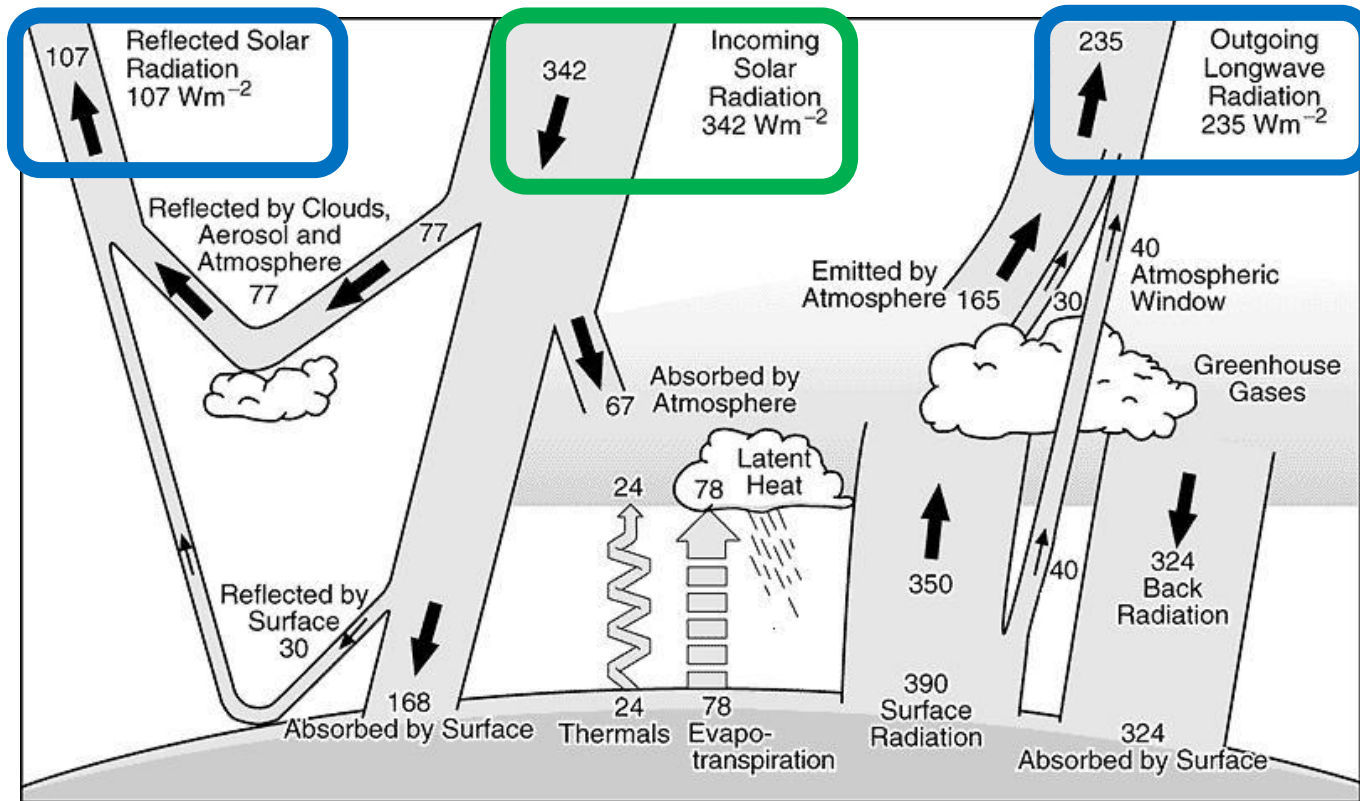
Source: climate.nasa.gov



# Sea Levels Rising

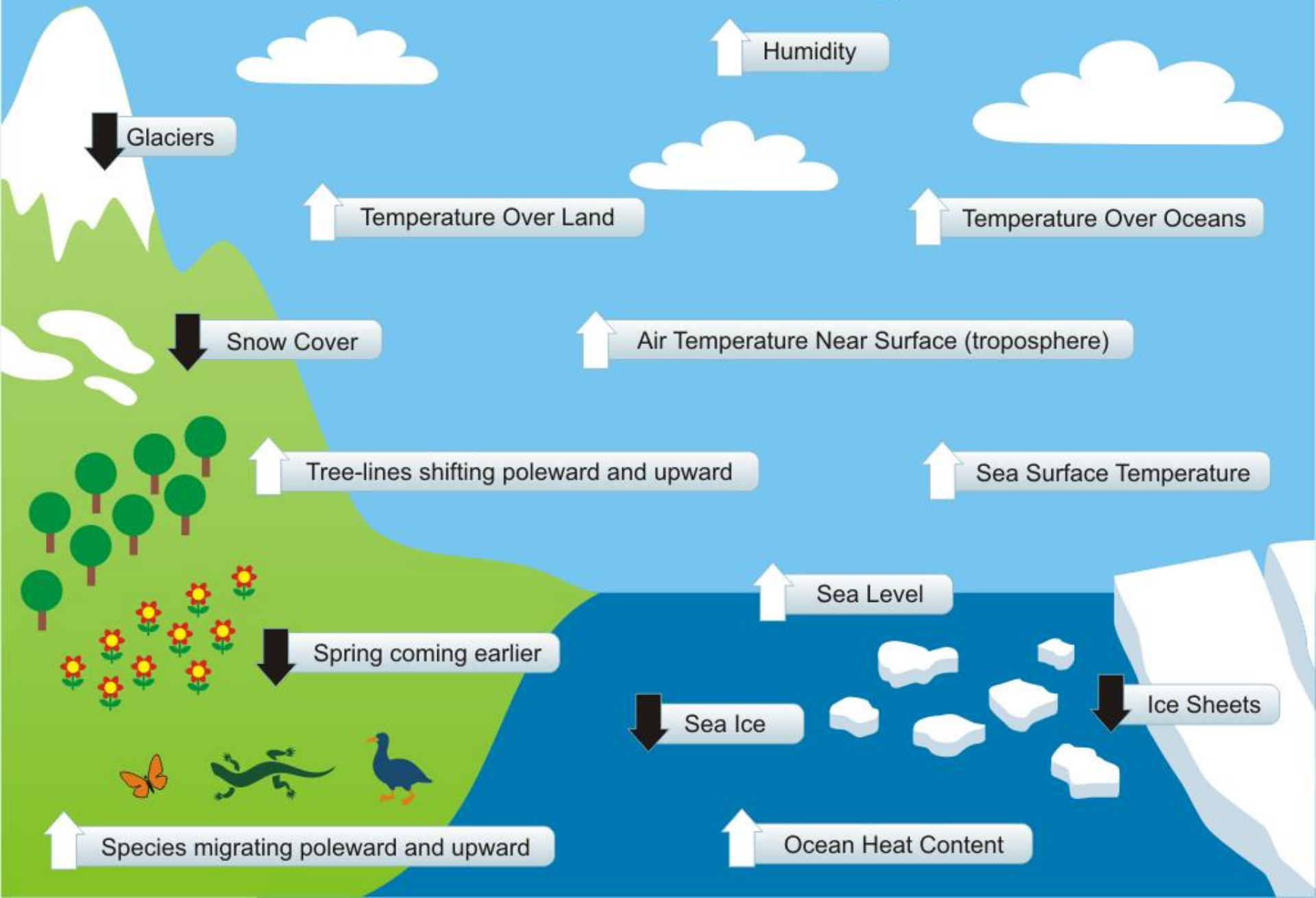


# Thermal Equilibrium



(in) **342  $\text{Wm}^{-2}$**  = **342  $\text{Wm}^{-2}$**  (out)

# Indicators of a Warming World



# Feedback Loops

## Positive Feedback Loop

Warming of Earth leads to events that further warm the Earth

- **Melting ice**
  - Higher temps decrease ice cover on the planet
  - Decreases albedo
- **Melting permafrost**
  - Releases methane
- **Methane on ocean floor**
  - Higher ocean temperatures release frozen methane deposits

## Negative Feedback Loop

Warming of Earth leads to events that start to cool the Earth

- **More Clouds**
  - Higher temps evaporate more water
  - Increase Albedo
- **Increased Photosynthesis**
  - More CO<sub>2</sub> leads to more plant life that absorbs CO<sub>2</sub>
- **Renewable Investment**
  - Higher temperatures lead to a greater urgency for change

# How we know we're causing global warming

Shrinking upper atmosphere

Less heat escaping to space

Cooling upper atmosphere

Rising tropopause

Winter warming faster than summer

More fossil fuel carbon in the air

Less oxygen in the air

More heat returning to Earth

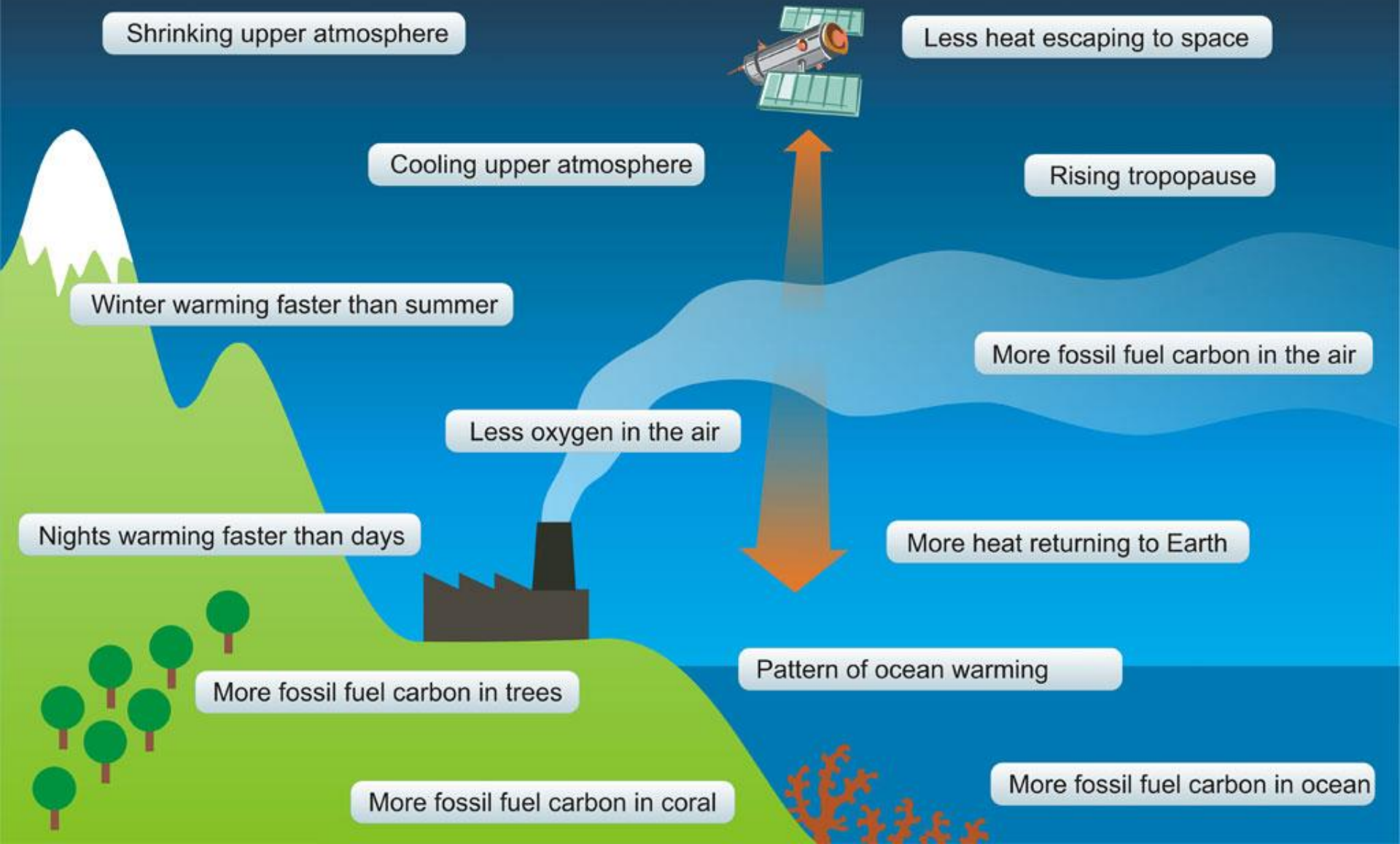
Nights warming faster than days

Pattern of ocean warming

More fossil fuel carbon in trees

More fossil fuel carbon in ocean

More fossil fuel carbon in coral





# Why Deny?

