## IB Physics SL - Study Guide

## Science Skills

List the seven fundamental base units and their abbreviations:

|  | Unit | Abbreviation |
| :--- | :---: | :---: |
| Length | Meter | m |
| Mass | Kilogram | kg |
| Time | Second | s |
| Electric Current | Ampere | A |
| Temperature | Kelvin | K |
| Amount of Substance | Mole | mol |
| Luminous Intensity | Candela | cd |

Metric Prefixes - List the unit prefixes in their appropriate decimal position


Dimensional Analysis
Convert the following:
$20 \mathrm{mi} \mathrm{hr}^{-1} \rightarrow \mathrm{~m} \mathrm{~s}^{-1}$

$$
\frac{20 \mathrm{mi}}{1 \mathrm{hr}} \times \frac{1609 \mathrm{~m}}{1 \mathrm{mi}} \times \frac{1 \mathrm{hr}}{60 \mathrm{~min}} \times \frac{1 \mathrm{~min}}{60 \mathrm{~s}}=8.9 \frac{\mathrm{~m}}{\mathrm{~s}}=8.9 \mathrm{~m} \mathrm{~s}^{-1}
$$

$0.0007 \mathrm{~km}^{2} \rightarrow \mathrm{~m}^{2}$

$$
0.0007 \mathrm{~km}^{2} \times \frac{1000 \mathrm{~m}}{1 \mathrm{~km}} \times \frac{1000 \mathrm{~m}}{1 \mathrm{~km}}=700 \mathrm{~m}^{2} \quad \text { or } \quad 0.0007 \mathrm{~km}^{2} \times\left(\frac{1000 \mathrm{~m}}{1 \mathrm{~km}}\right)^{2}=700 \mathrm{~m}^{2}
$$

Determine the units for Q :

$$
Q=(k g)\left(J k g^{-1} K^{-1}\right)(K)=\frac{(k g)(J)(K)}{k g K}=J
$$

| $m$ (mass) | kg |
| :---: | :---: |
| c (specific heat) | $\mathrm{J} \mathrm{kg}^{-1} \mathrm{~K}^{-1}$ |
| $\Delta T$ (change in temp) | K |


|  | Scalar | Vector |
| :---: | :---: | :---: |
| How far $(\mathrm{m})$ | Distance | Displacement |
| How fast $\left(\mathrm{m} \mathrm{s}^{-1}\right)$ | Speed | Velocity |


|  | Displacement vs Time | Velocity vs Time | Acceleration vs Time |
| :---: | :---: | :---: | :---: |
| Meaning of the Graph | Slope: $\quad$ Velocity | Slope: Acceleration <br> Area under the Curve: Displacement | Area under the Curve: <br> Velocity |
| Constant Displacement | $\xrightarrow{\square}$ |  |  |
| Constant Positive Velocity |  | $\xrightarrow{\square}$ |  |
| Constant Negative Velocity |  |  |  |
| Constant Positive Acceleration (speeding up) |  |  | $\xrightarrow{\square}$ |
| Constant Negative Acceleration (slowing down) |  |  |  |


|  | Variable <br> Symbol | Unit |
| :---: | :---: | :---: |
| Displacement | s | m |
| Initial Velocity | u | $\mathrm{m} \mathrm{s}^{-1}$ |
| Final Velocity | V | $\mathrm{m} \mathrm{s}^{-1}$ |
| Acceleration | a | $\mathrm{m} \mathrm{s}^{-2}$ |
| Time | t | s |


| Kinematic Equations | S | u | V | a | t |
| :---: | :---: | :---: | :---: | :---: | :---: |
| $v=u+a t$ |  | $\checkmark$ | $\checkmark$ | $\checkmark$ | $\checkmark$ |
| $s=u t+\frac{1}{2} a t^{2}$ | $\checkmark$ | $\checkmark$ |  | $\checkmark$ | $\checkmark$ |
| $v^{2}=u^{2}+2 a s$ | $\checkmark$ | $\checkmark$ | $\checkmark$ | $\checkmark$ |  |
| $S=\frac{(v+u) t}{2}$ | $\checkmark$ | $\checkmark$ | $\checkmark$ |  | $\checkmark$ |


| Horizontal Component | $A_{H}=A \cos \theta$ |  |
| :---: | :---: | :---: |
| Vertical Component | $A_{V}=A \sin \theta$ |  |


|  | Vertical |
| :---: | :---: |
| $s$ |  |
| $u$ | $0 \mathrm{~m} \mathrm{~s}^{-1}$ |
| $v$ |  |
| $a$ | $-9.81 \mathrm{~m} \mathrm{~s}^{-2}$ |
| $t$ |  |



Forces

| Type of Force | Variable | Description/Important Properties | Equation |
| :---: | :---: | ---: | ---: |
| Weight | $\mathrm{F}_{\mathrm{g}}$ | Force of gravity on an object with mass | $\mathrm{F}_{\mathrm{g}}=\mathrm{mg}$ |
| Tension | $\mathrm{F}_{\mathrm{T}}$ | Always pulls in the same direction as the <br> rope or chain providing the tension |  |
| Normal <br> Reaction | R | Always perpendicular to a surface |  |
| Friction | $\mathrm{F}_{\mathrm{f}}$ | Always opposes the motion of an object | $\mathrm{F}_{\mathrm{f}}=\mu \mathrm{R}$ |
| Air Resistance | $\mathrm{F}_{\text {air }}$ | Increases with surface area and velocity |  |

If an object has a net force of zero its motion is either:
Not moving (velocity $=0 \mathrm{~m} \mathrm{~s}^{-1}$ )
or
Moving at a constant velocity

## Newton's Laws

| Newton's <br> First Law | A object at rest remains at rest and an object in motion remains in motion <br> until and unless an external force acts upon it (Unbalanced force). |
| :---: | :--- |
| Newton's <br> Second Law | The rate of change of momentum of an object is proportional to the <br> resultant force acting on the body and is in the same direction. ( $F=\mathrm{ma}$ ) |
| Newton's <br> Third Law | All forces occur in pairs. Every action has an equal and opposite reaction |


|  | Variable <br> Symbol | Unit |
| :---: | :---: | :---: |
| Force | F | N |
| Mass | m | kg |
| Acceleration | a | $\mathrm{m} \mathrm{s}^{-1}$ |
| Normal Reaction Force | R | N |
| Coefficient of Kinetic Friction | $\mu_{\mathrm{d}}$ | -- |
| Coefficient of Static Friction | $\mu_{\mathrm{s}}$ | -- |



| Sliding to a Stop | Constant Velocity |
| :---: | :---: |
|  |  |
| $F_{\text {net }}=F_{f}$ | $F_{\text {net }}=0 \mathrm{~N} \quad \mathrm{~F}_{\text {pull }}=\mathrm{F}_{\mathrm{f}}$ |



## Forces on a Ramp

| Equilibrium |  |
| :---: | :---: |
| $R$ | $F_{\perp}$ |
| $F_{f}$ | $F_{\\| l}$ |
| $F_{\text {net }}$ | 0 N |
| $a$ | $0 \mathrm{~m} \mathrm{~s}^{-1}$ |


| Accelerating |  |
| :---: | :---: |
| $R$ | $F_{\perp}$ |
| $F_{f}$ | $\mu R$ |
| $F_{\text {net }}$ | $F_{I I}-F_{f}$ |
| $a$ | $F_{\text {net }} / m$ |

## Circular Motion

|  | Variable <br> Symbol | Unit |
| :---: | :---: | :---: |
| Distance | d | m |
| Angular Distance | $\theta$ | rad |
| Angular Velocity | $\omega$ | $\mathrm{rad} \mathrm{s}^{-1}$ |
| Linear Velocity | V | $\mathrm{m} \mathrm{s}^{-1}$ |
| Centripetal Acceleration | $a$ | $\mathrm{~m} \mathrm{~s}^{-2}$ |
| Centripetal Force | $\mathrm{F}_{\mathrm{c}}$ | N |

Draw in vectors for $v, a_{c}$, and $F_{c} \rightarrow$


Data Booklet Equations:
$v=\omega r$
$a=\frac{v^{2}}{r}=\frac{4 \pi^{2} r}{T^{2}}$
$F=\frac{m v^{2}}{r}=m \omega^{2} r$

## Defining Circular Motion

|  | Period | T | S | Angular Velocity | $\omega$ | $\mathrm{rad} \mathrm{s}^{-1}$ |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: |
|  | Time per revolution |  |  | $\omega=\frac{2 \pi}{T}$ |  |  |

## Vertical Circular Motion

|  | Bottom: |  |
| :--- | :--- | :--- |
|  | $F_{\text {net }}=F_{c}=F_{T}+F_{g}$ | Top: |

Top: $\quad$ Bottom:

Circular Motion with Friction and Angles

|  |  | Relationships between variables: |
| :--- | :--- | :--- | :--- |
|  |  | $F_{f}=F_{g}$ |





Relationships between variables:

$$
\begin{aligned}
T_{y} & =F_{g} \\
F_{c} & =T_{x}
\end{aligned}
$$

## Energy

|  | Variable <br> Symbol | Unit |
| :---: | :---: | :---: |
| Work | W | Joules [J] |
| Power | P | Watts [W] |
| Kinetic Energy | $\mathrm{E}_{\mathrm{k}}$ | J |
| Elastic Potential Energy | $\mathrm{E}_{\mathrm{p}}$ | J |
| Gravitational Potential Energy | $\Delta \mathrm{E}_{\mathrm{p}}$ | J |
| Spring Constant | k | $\mathrm{N} \mathrm{m}^{-1}$ |
| Spring Stretch | $\Delta x$ | m |

Data Booklet Equations:

$$
\begin{gathered}
W=F s \cos \theta \\
E_{k}=\frac{1}{2} m v^{2} \\
E_{p}=\frac{1}{2} k \Delta x^{2} \\
\Delta E_{p}=m g \Delta h \\
\text { power }=F v
\end{gathered}
$$

## Calculating Work

| Constant force at an angle: $W=F s \cos \theta$ |  |
| :---: | :---: |
| Varying Force: <br> Area under the curve |  |
| Examples of no work being done for an object in motion: <br> - Pushing something that doesn't move (no displacement, no work) <br> - Waiter carrying a tray horizontally (force is vertical, motion is horizontal) <br> - Orbiting object (velocity is tangent to path, force is toward the center) |  |

Calculating Power

| In terms of work and time: | In terms of force and velocity: |
| :---: | :--- |
| Power $=\frac{\text { Work }}{\text { Time }}$ | Power $=$ Force $\times$ Velocity $=F v$ |

Units

|  | Standard Unit | From Equation | Fundamental SI Units |
| :---: | :---: | :---: | :---: |
| Work | J | N m | $\mathrm{kg} \mathrm{m}^{2} \mathrm{~s}^{-2}$ |
| Power | W | $\mathrm{J} \mathrm{s}^{-1}$ | $\mathrm{~kg} \mathrm{~m}^{2} \mathrm{~s}^{-3}$ |

Types of Energy

| Kinetic Energy | Elastic Potential Energy | Gravitational Potential Energy |
| :---: | :---: | :---: |
| $\frac{1}{2} m v^{2}$ | $\frac{1}{2} k \Delta x^{2}$ | $m g \Delta h$ |

Conservation of Energy

## Total Energy Before = Total Energy After

## Work-Energy Theorem

Work $\rightarrow$ Energy

$$
F s=\frac{1}{2} m v^{2}
$$

Energy $\rightarrow$ Work
$\frac{1}{2} m v^{2}=F s$

|  | Variable <br> Symbol | Unit |
| :---: | :---: | :---: |
| Momentum | p | $\mathrm{kg} \mathrm{m} \mathrm{s}^{-1}$ |
| Mass | m | kg |
| Velocity | V | $\mathrm{m} \mathrm{s}^{-1}$ |
| Time | t | s |
| Kinetic Energy | $\mathrm{E}_{\mathrm{K}}$ | J |
| Impulse | Impulse | ${\mathrm{Ns} \mathrm{or} \mathrm{kg} \mathrm{m} \mathrm{s}^{-1}}$ |

Data Booklet Equations:

$$
\begin{aligned}
p & =m v \\
F & =\frac{\Delta p}{\Delta t} \\
E_{K} & =\frac{p^{2}}{2 m}
\end{aligned}
$$

Impulse $=F \Delta t=\Delta p$

## Conservation of Energy Problems

"Explosion"

## Types of Collisions

| Elastic | Kinetic Energy is conserved (perfect hit and bounce) <br> *Typically just found in particle collisions |
| :---: | :---: |
| Inelastic | Kinetic Energy is not conserved |

## Calculating Impulse

| Constant force: <br> Force $\times$ Time F $\Delta \mathrm{t}$ |  |
| :---: | :---: |
| Varying Force: <br> Area under a Force vs Time Graph |  |

## Impulse-Momentum Equation

$$
F \Delta t=\Delta p=m \Delta v=m v-m u
$$

## Collision Safety

Explain (using impulse, force, and time) how to decrease the force acting on an object undergoing a collision:
Impulse is the same overall regardless of the impact style because the object has a set mass and impact velocity. The force can be decreased by increasing the time of the impact.

$$
\text { Impulse }=\mathrm{F}_{\Delta \mathrm{t}} \quad \text { or } \quad \text { Impulse }={ }_{F} \Delta \mathrm{t}
$$

## Waves - Sound

|  | Variable <br> Symbol | Unit |
| :---: | :---: | :---: |
| Period | $T$ | s |
| Frequency | $f$ | Hz |
| Wavelength | $\lambda$ | m |
| Amplitude | A | m |
| Wave Speed | $v$ | $\mathrm{~m} \mathrm{~s}^{-1}$ |

Data Booklet Equations:

$$
\begin{aligned}
T & =\frac{1}{f} \\
c & =f \lambda
\end{aligned}
$$

## Simple Harmonic Motion Graphs



| Types of Waves | Picture | Definition | Examples |
| :---: | :---: | :---: | :---: |
| Transverse | NVVA | Particles move perpendicular to the motion of the wave | - Light <br> - Ripples in a Pond <br> - Earthquakes |
| Longitudinal | IT | Particles move parallel to the motion of the wave | - Sound <br> - Earthquakes |

## Parts of a Wave

Label the Wave:

- Amplitude
- Wavelength
- Crest
- Trough


Harmonics

|  | Open Pipe |  | Closed Pipe |  | String |  |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| End Conditions | Antinode | Antinode | Node | Antinode | Node | Node |
| $3^{\text {dra }}$ Harmonic |  |  |  |  | $2$ |  |
|  |  |  |  | $\frac{5}{4} \lambda$ |  |  |
| $2^{\text {nd }}$ Harmonic |  |  |  |  |  |  |
|  |  | $1 \lambda$ |  | ${ }_{4}{ }^{2}$ |  |  |
| $1^{\text {st }}$ Harmonic (Fundamental) |  |  |  |  |  |  |
|  |  |  |  | $\frac{1}{4} \lambda$ |  |  |

## Interference

| Constructive | Path Difference $=n \lambda$ | Destructive | Path Difference $=(n+1 / 2) \lambda$ |
| :--- | :--- | :--- | :--- |
|  |  |  |  |

## Waves - Light

## Electromagnetic Spectrum

| A | Radiowaves |
| :---: | :--- |
| B | Microwaves |
| C | Infrared |
| D | Visible Light |
| E | Ultraviolet |
| F | X-Rays |
| G | Gamma Waves |



Index of Refraction

| Medium | Wave Speed (v) | Index of Refraction (n) | $\boldsymbol{n}_{1}$ |
| :---: | :---: | :---: | :---: |
| Vacuum | $3.00 \times 10^{8} \mathrm{~m} \mathrm{~s}^{-1}$ | 1.0000 |  |
| Air | $2.999 \times 10^{8} \mathrm{~m} \mathrm{~s}^{-1}$ | 1.0003 |  |
| Water | $2.256 \times 10^{8} \mathrm{~m} \mathrm{~s}^{-1}$ | 1.33 |  |
| Glass | $1.974 \times 10^{8} \mathrm{~m} \mathrm{~s}^{-1}$ | 1.52 |  |

## Refraction

| $\frac{n_{1}}{n_{2}}=\frac{\sin \theta_{2}}{\sin \theta_{1}}$ |  |  |
| :---: | :---: | :---: |
|  |  |  |

Critical Angle

| When $\theta_{1}=\theta_{c}$ |
| :--- | :--- | :--- |
| $\theta_{2}=90^{\circ}$ |$\quad \theta_{c}=\sin ^{-1}\left(\frac{n_{2}}{n_{1}}\right) \quad \stackrel{\text { Water }}{\square}$

## Reflection

| Law of Reflection |  |
| :---: | :---: | :---: |
| Angle of Incidence $=$ Angle of Reflection |  |

## Polarized Light

| $I=I_{0} \cos ^{2} \theta$ |  |  |
| :---: | :---: | :---: |
| I | Intensity Observed |  |
| I0 | Original Intensity |  |
| $\theta$ | Difference in Angle |  |

## Double Slit Experiment



## Electricity

Charge

| Symbol | q Unit | Coulombs [C] |  |
| :---: | :---: | :---: | :---: |
| Charge of 1 Electron |  |  | $1.6 \times 10^{-19} \mathrm{C}$ |
| \# of Electrons per Coulomb | $6.25 \times 10^{18} \mathrm{e}^{-}$ |  |  |

## Current

| Symbol | I | Unit | Amperes [A] |
| :---: | :---: | :---: | :---: |
| Unit in terms of Coulombs | $\mathrm{A}=\frac{\mathrm{C}}{\mathrm{S}}$ |  |  |

## Drift Speed

|  | Variable <br> Symbol | Unit |
| :---: | :---: | :---: |
| Current | I | A |
| \# of Electrons per $\mathrm{m}^{3}$ | n | --- |
| Cross Sectional Area | A | $\mathrm{M}^{2}$ |
| Drift Speed | v | $\mathrm{m} \mathrm{s}^{-1}$ |
| Charge | q | C |
| Cross Sectional Area: |  |  |
| A $=\pi r^{2}$ |  |  |

## Electrical Properties

| Property | What is it? | Symbol | Unit |
| :---: | :---: | :---: | :---: |
| Voltage | Potential Difference | V | Volts [V] |
| Current | The rate at which charges move through a wire | I | Amperes [A] |
| Resistance | How hard it is for a current to <br> flow through a conductor | R | Ohms [ $\Omega$ ] |

## Kirchhoff's Laws

| $\Sigma \mathrm{I}=0$ (junction) |  |  |  | $\Sigma \mathrm{V}=0$ (loop) |  |  |  |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| The total current coming into a junction must equal the total current leaving the same junction |  |  |  | The sum of the voltages (potential differences) provided must equal the voltages dissipated across components |  |  |  |
|  |  |  |  | Across resistors |  | Alway | egative |
| Entering Junction | $\rightarrow$ | $\bullet$ | Positive | Negative to Positive | $\rightarrow$ | + | Positive |
| Exiting Junction | - | $\rightarrow$ | Negative | Positive to Negative | $\rightarrow$ | -ト | Negative |


| $\frac{V}{I \times R}$ | $\begin{aligned} & V=I \times R \\ & \hline I=\frac{V}{R} \\ & R=\frac{V}{I} \end{aligned}$ | Ohmic Resistor | $\xrightarrow[\text { Potential Difference/V }]{\text { Non-Ohmic Resistor }}$ |
| :---: | :---: | :---: | :---: |

## Equivalent Resistance

|  | Drawing with $R_{1}$ and $R_{2}$ | Equation |
| :---: | :---: | :---: |
| Series | $\rightarrow$ | $R_{\text {total }}=R_{1}+R_{2}+\cdots$ |
| Parallel | $\longrightarrow$ | $\frac{1}{R_{\text {total }}}=\frac{1}{R_{1}}+\frac{1}{R_{2}}+\cdots$ |

## Measuring Circuits

|  | Ammeter | Voltmeter |
| :---: | :---: | :---: |
| Ideal Resistance | $\mathrm{R}=0 \Omega$ | $\mathrm{R}=\infty \Omega$ |
| How is it connected to the <br> component being measured? | Ammeters must be <br> connected in series | Voltmeters must be <br> connected in parallel |
| Drawing of meter measuring $\mathrm{R}_{1}$ |  |  |

## Resistivity

|  | Variable <br> Symbol | Unit |
| :---: | :---: | :---: |
| Resistivity | $\rho$ | $\Omega \mathrm{m}$ |
| Resistance | R | $\Omega$ |
| Cross Sectional Area | A | $\mathrm{m}^{2}$ |
| Length | L | m |

Data Booklet Equation:

$$
\rho=\frac{\mathrm{RA}}{\mathrm{~L}}
$$

Cross Sectional Area:
$\mathrm{A}=\pi \mathrm{r}^{2}$

## Power

| In terms of $V$ and $I$ | In terms of $I$ and $R$ | In terms of $V$ and $R$ |
| :---: | :--- | ---: |
| $P=V \times I$ | $P=I^{2} R$ | $P=\frac{V^{2}}{R}$ |

## Voltage Dividers

|  | Light-Dependent Resistor | Thermistor |
| :---: | :---: | :---: |
| Symbol |  |  |
| Relationship | Light Increases | Heat Increases |
|  | Resistance Decreases | Resistance Decreases |
| Circuit | Switch turns on in the dark: | Switch turns on in a fire: |

## Batteries

| Primary Cells | Secondary Cells |
| :---: | :---: |
| Cannot be recharged | Can be recharged by passing a current <br> through the battery in the opposite direction <br> as it would normally travel |


|  | Variable <br> Symbol | Unit |
| :---: | :---: | :---: |
| Electromotive Force (e.m.f) | $\varepsilon$ | V |
| Current | I | A |
| Circuit Resistance | R | $\Omega$ |
| Internal Resistance | r | $\Omega$ |

Data Booklet Equation:

$$
\varepsilon=\mathrm{I}(\mathrm{R}+\mathrm{r})
$$

## Force Fields

Forces between objects
Coulomb's Law

|  | Variable <br> Symbol | Unit |
| :---: | :---: | :---: |
| Electrostatic Force | F | N |
| Object 1 Charge | $\mathrm{q}_{1}$ | C |
| Object 2 Charge | $\mathrm{q}_{2}$ | C |
| Separation Distance | r | M |
| Coulomb Constant | k | $\mathrm{N} \mathrm{m}^{2} \mathrm{C}^{-2}$ |
| Permittivity of Free Space | $\varepsilon_{0}$ | $\mathrm{C}^{2} \mathrm{~N}^{-1} \mathrm{~m}^{-2}$ |

Data Booklet Equations:

$$
\begin{aligned}
F & =k \frac{q_{1} q_{2}}{r^{2}} \\
& =\frac{1}{4 \pi \varepsilon_{0}}
\end{aligned}
$$

$$
\begin{aligned}
& \mathrm{k}=8.99 \times 10^{9} \mathrm{~N} \mathrm{~m}^{2} \mathrm{C}^{-2} \\
& \varepsilon_{0}=8.85 \times 10^{-12} \mathrm{C}^{2} \mathrm{~N}^{-1} \mathrm{~m}^{-2}
\end{aligned}
$$

## Universal Law of Gravitation

|  | Variable <br> Symbol | Unit |
| :---: | :---: | :---: |
| Gravitational Force | F | N |
| Object 1 Mass | M | kg |
| Object 2 Mass | m | kg |
| Separation Distance | r | m |
| Gravitational Constant | G | $\mathrm{N} \mathrm{m}^{2} \mathrm{~kg}^{-2}$ |

## Force Fields

| Electric Field |  |  | Gravitational Field |  |  |
| :---: | :---: | :---: | :---: | :---: | :---: |
| Symbol | E | Data Booklet Equation: |  |  |  |
| Units | $\mathrm{FC}^{-1}$ |  | Symbol | F |  |
|  |  | Units | $\mathrm{F} \mathrm{kg}^{-1}$ | $g=\frac{F}{m}$ | $g=G \frac{M}{r^{2}}$ |



## Right Hand Rule:

| Right Hand Rule \#1 |  |  | Right Hand Rule \#2 |  |  | Right Hand Rule \#3 |  |  |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| Magnetic field around a current carrying wire |  |  | Pole orientation for a coil of wire (electromagnet, solenoid, etc.) |  |  | Electromagnetic force direction on a wire or moving particle |  |  |
| Thumb | Current |  | Thumb | North Pole |  | Thumb |  | Current |
| Fingers | Magnetic Field |  | Fingers | Current |  | Fingers | Magnetic Field |  |
|  | $\begin{aligned} & ((\mathrm{I} \stackrel{* *}{\rightleftarrows} \\ & \stackrel{* * *}{\stackrel{*}{\rightleftarrows}})) \end{aligned}$ |  | $\text { s. }\left\\|\left\\|\left\\|\left\\|\left\\|\left\\|\\|_{\uparrow}^{\mathrm{N}}\right.\right.\right.\right.\right.\right.$ |  | $\text { N }\left\|\left\|\left\|\left\|\left\|\left\|\left\|\left\|\left\|\left\|\left.\right\|_{\dagger} \mathrm{s}\right.\right.\right.\right.\right.\right.\right.\right.\right.\right.$ |  |  |  |
|  |  |  |  |  | ${ }_{s}{ }^{\dagger}\left(\|1\|(1\|1\| 1)^{\dagger}{ }_{\mathrm{N}}\right.$ |  |  |  |

## Electromagnetic Force

|  | Variable <br> Symbol | Unit |
| :---: | :---: | :---: |
| Magnetic Force | F | N |
| Magnetic Field Strength | B | T |
| Current | I | A |
| Wire Length | L | m |
| Angle to Field | $\theta$ | $\circ$ |
| Particle Charge | q | C |
| Particle Velocity | V | $\mathrm{m} \mathrm{s}^{-1}$ |

Data Booklet Equations:

Wire:

$$
F=B I L \sin \theta
$$

Particle:

$$
F=q v B \sin \theta
$$

Charged Particles Moving through a Magnetic Field


## Thermal Physics



Specific Heat Capacity and Specific Latent Heat

|  | Variable Symbol | Unit |
| :---: | :---: | :---: |
| Heat Energy | Q | J |
| Mass | m | kg |
| Specific Heat Capacity | C | $\mathrm{Jkg}^{-1} \mathrm{~K}^{-1}$ |
| Change in Temperature | $\Delta \mathrm{T}$ | K |
| Specific Latent Heat | L | J kg |$\quad$|  |
| :---: |
| $Q=m c \Delta T$ |
| $Q=m L$ |
|  |


| $\mathrm{E}_{\mathrm{K}}$ | Kinetic Energy $\rightarrow$ Temperature |
| :--- | :--- |

Ep Potential Energy $\rightarrow$ Phase Change

## Heating Curves



## Pressure

|  | Variable Symbol | Unit |  |
| :---: | :---: | :---: | :---: |
| Force | F | N |  |
| Area | A | $\mathrm{m}^{3}$ |  |
| Pressure | p | $\mathrm{N} \mathrm{m}^{-3}$ |  |

## Data Booklet Equation:

$$
p=\frac{F}{A}
$$

## Kinetic Theory and Temperature

|  | Variable Symbol | Unit |
| :---: | :---: | :---: |
| Average Kinetic Energy | $\overline{\mathrm{E}}_{\mathrm{k}}$ | J |
| Absolute Temperature | T | K |
| Boltzmann's Constant | $\mathrm{k}_{\mathrm{b}}$ | $\mathrm{J} \mathrm{K}^{-1}$ |

## Data Booklet Equation:

$$
\begin{aligned}
& \bar{E}_{K}=\frac{3}{2} k_{B} T=\frac{3}{2} \frac{R}{N_{A}} T \\
& k_{B}=1.38 \times 10^{-23} \mathrm{~J} \mathrm{~K}^{-1}
\end{aligned}
$$

| Avogadro's Number | $\mathrm{N}_{\mathrm{A}}$ | $6.02 \times 10^{23}$ |
| :--- | :--- | :--- |

## Ideal Gas Law

|  | Variable Symbol | Unit |
| :---: | :---: | :---: |
| Pressure | p | Pa |
| Volume | V | $\mathrm{m}^{-3}$ |
| Number of Molecules | n | mol |
| Gas Constant | R | $\mathrm{J} \mathrm{K}^{-1} \mathrm{~mol}^{-1}$ |
| Temperature | T | K |

Data Booklet Equations:
$p V=n R T \quad R=8.31 \mathrm{~J} \mathrm{~K}^{-1} \mathrm{~mol}^{-1}$
Conditions for Ideal Gases:

## Ideal Gas Relationships





## Atomic Physics

## Types of Decay




## Half Life



## Mass-Energy Equivalence

|  | Variable Symbol | Unit |
| :---: | :---: | :---: |
| Energy | E | J |
| Mass | m | kg |
| Speed of Light | c | $\mathrm{m} \mathrm{s}^{-1}$ |

Data Booklet Equation:

$$
\begin{gathered}
E=m c^{2} \\
c=3.00 \times 10^{8} \mathrm{~m} \mathrm{~s}^{-1}
\end{gathered}
$$

| Unified Atomic Mass Unit | $u$ | $1.661 \times 10^{-27} \mathrm{~kg}$ | 1.000000 u | $931.5 \mathrm{MeV} \mathrm{c}^{-2}$ |
| :--- | :--- | :--- | :--- | :--- |


| Electron Rest Mass | $m_{\mathrm{e}}$ | $9.110 \times 10^{-31} \mathrm{~kg}$ | 0.000549 u | $0.511 \mathrm{MeV} \mathrm{c}^{-2}$ |
| :---: | :---: | :---: | :---: | :---: |
| Proton Rest Mass | $m_{\mathrm{p}}$ | $1.673 \times 10^{-27} \mathrm{~kg}$ | 1.007276 u | $938 \mathrm{MeV} \mathrm{c}^{-2}$ |
| Neutron Rest Mass | $m_{\mathrm{n}}$ | $1.675 \times 10^{-27} \mathrm{~kg}$ | 1.008665 u | $940 \mathrm{MeV} \mathrm{c}^{-2}$ |

Converting between Joules and Electron-Volts
$\{$ Energy in eV $\}=\frac{\{\text { Energy in } J\}}{1.60 \times 10^{-19}} \quad\{$ Energy in $J\}=\{$ Energy in eV$\} \times 1.60 \times 10^{-19}$

## Process for Calculating Binding Energy

1. Add up the "before and after" masses
2. Find the mass defect by taking the difference
3. Convert atomic mass units ( u ) into $\mathrm{MeV} \mathrm{c}^{-2}$ by using the conversion factor $1 \mathrm{u}=931.5 \mathrm{MeV} \mathrm{c}^{-2}$
4. The $c^{-2}$ cancels out when converting to energy using $E=m c^{2}$ so this is your binding energy

|  | Describe | Examples | Challenges |
| :---: | :---: | :---: | :---: |
| FiSSiOn | Lighter elements are <br> created by splitting <br> heavier elements | Nuclear Power <br> Nuclear Weapons | Proper amounts of <br> fissionable elements <br> required to maintain <br> chain reaction |
| FUSiOn | Heavier elements are <br> created by combining <br> lighter elements | The Sun/Stars | Requires high heat and <br> high pressure |

## Fundamental Particles

The following two tables are provided in the IB Physics Data Booklet

| Charge | Quarks |  |  |  |
| :---: | :---: | :---: | :---: | :---: | | Baryon |
| :---: |
| Number | \left\lvert\, | $\frac{2}{3}$ | u | c | t |
| :---: | :---: | :---: | :---: |
| $-\frac{1}{3}$ | d | s | b | | All quarks have a strangeness number of 0 except the |
| :---: |
| strange quark that has a strangeness number of -1 |\right.


| Charge | Leptons |  |  |
| :---: | :---: | :---: | :---: |
| -1 | e | $\mu$ | $\tau$ |
| 0 | $v_{e}$ | $v_{\mu}$ | $v_{\tau}$ |
| All leptons have a lepton number of 1 and <br> antileptons have a lepton number of -1 |  |  |  |


| Quarks |  |  |  |
| :---: | :---: | :---: | :---: |
| Symbol | Name | Charge | Baryon \# |
| u | Up | $+\frac{2}{3}$ | $\frac{1}{3}$ |
| d | Down | $-\frac{1}{3}$ | $\frac{1}{3}$ |
| C | Charm | $+\frac{2}{3}$ | $\frac{1}{3}$ |
| S | Strange | $-\frac{1}{3}$ | $\frac{1}{3}$ |
| t | Top | $+\frac{2}{3}$ | $\frac{1}{3}$ |
| b | Bottom | $-\frac{1}{3}$ | $\frac{1}{3}$ |


| Leptons |  |  |  |
| :---: | :---: | :---: | :---: |
| Symbol | Name | Charge | Lepton \# |
| e | Electron | -1 | 1 |
| $\mu$ | Muon | -1 | 1 |
| $\tau$ | Tau | -1 | 1 |
| $\nu_{e}$ | Electron <br> Neutrino | 0 | 1 |
| $\nu_{\mu}$ | Muon <br> Neutrino | 0 | 1 |
| $\nu_{\tau}$ | Tau Neutrino | 0 | 1 |


| Anti-Quarks |  |  |  |
| :---: | :---: | :---: | :---: |
| Symbol | Name | Charge | Baryon \# |
| $\overline{\mathrm{u}}$ | Antiup | $-\frac{2}{3}$ | $-\frac{1}{3}$ |
| $\overline{\mathrm{~d}}$ | Antidown | $+\frac{1}{3}$ | $-\frac{1}{3}$ |
| $\overline{\mathrm{c}}$ | Anticharm | $-\frac{2}{3}$ | $-\frac{1}{3}$ |
| $\overline{\mathrm{~s}}$ | Antistrange | $+\frac{1}{3}$ | $-\frac{1}{3}$ |
| $\overline{\mathrm{t}}$ | Antitop | $-\frac{2}{3}$ | $-\frac{1}{3}$ |
| $\overline{\mathrm{~b}}$ | Antibottom | $+\frac{1}{3}$ | $-\frac{1}{3}$ |


| Anti-Leptons |  |  |  |
| :---: | :---: | :---: | :---: |
| Symbol | Name | Charge | Lepton \# |
| $\overline{\mathrm{e}}$ | Antielectron <br> (positron) | +1 | -1 |
| $\bar{\mu}$ | Antimuon | +1 | -1 |
| $\bar{\tau}$ | Antitau | +1 | -1 |
| $\bar{\nu}_{e}$ | Electron <br> Antineutrino | 0 | -1 |
| $\bar{\nu}_{\mu}$ | Muon <br> Antineutrino | 0 | -1 |
| $\bar{\nu}_{\tau}$ | Tau <br> Antineutrino | 0 | -1 |

## Explain the phenomenon of Quark Confinement:

Quarks have never been observed on their own. The amount of energy required to overcome the strong nuclear force holding the quarks together gets converted into mass and forms a new quark pair.


Fundamental Forces

|  | Strength | Distance |
| :---: | :---: | :---: |
| Gravitational | Weakest | Long Range |
| Weak | Weak | Short Range |
| Electromagnetic | Strong | Very Long Range |
| Strong | Strongest | Very Short Range |

Particle Configurations

| Proton | Neutron |  |  |
| :---: | :---: | :---: | :---: |
| d | $\mathbf{u}$ |  |  |
| U $\mathbf{u}$ | $\mathbf{d} \mathbf{d}$ |  |  |
| Total Charge | +1 | Total Charge | 0 |

## Feynman Diagrams

| You can only draw two kinds of lines $\qquad$ WM | You can only connect these lines if you have two lines with arrows meeting a single wiggly line | The $x$-axis represents time and is read from left to right. Everything left of the vertex is the "before" condition. |
| :---: | :---: | :---: |
| Beta-Negative Decay |  | Beta-Positive Decay |
|  |  | $+e^{+}+v_{e}^{w^{+}}$ |

## Energy Production

## Global Energy Usage

| Rank | Energy Source | $\%$ |
| :---: | :---: | :---: |
| 1 | Oil | $32 \%$ |
| 2 | Coal | $28 \%$ |
| 3 | Natural Gas | $22 \%$ |
| 4 | Biomass | $10 \%$ |
| 5 | Nuclear | $5 \%$ |
| 6 | Hydropower | $2.5 \%$ |



## Efficiency

| Efficiency $=\frac{\text { useful work out }}{\text { total work in }}=\frac{\text { useful power out }}{\text { total power in }}$ |  |  |
| :--- | :---: | :---: |
| Sankey Diagram Rules: <br> Width of the arrow proportional to the <br> amount of energy | Energy <br> In | Out |

## Energy Density

|  | Definition | Units |
| :---: | :---: | :---: |
| Specific Energy | Energy transferred per unit mass | $\mathrm{J} \mathrm{kg}^{-1}$ |
| Energy Density | Energy transferred per unit volume | $\mathrm{J} \mathrm{m}^{-3}$ |

## Primary and Secondary Sources

| Primary Energy Sources | Secondary Energy Sources |
| :---: | :---: |
| Energy sources found in the natural environment <br> (fossil fuels, solar, wind, nuclear, hydro, etc.) | Useful transformations of the primary sources <br> (electricity, pumped storage for hydro, etc.) |

## Fossil Fuels

| Number of years left in global reserves |  |
| :---: | :---: |
| Coal | $\sim 100-150$ years |
| Oil | $\sim 50$ years |
| Natural Gas | $\sim 50$ years |

## Describe the process of Fracking:

1. Drill hole into shale rock
2. Inject fracking fluid at high pressure to create cracks
3. Extract newly released natural gas
4. Seal fracking fluid in the hole

## Nuclear Power

|  | $\%$ of U-235 |
| :---: | :---: |
| Uranium Ore | $0.7 \%$ |
| Fuel-Grade | $3.5 \%$ |
| Weapons-Grade | $90 \%$ |

Why is the concentration of U-235 important?
Only U-235 can undergo a fission chain reaction
What is done with the nuclear waste?
Stored on-site in spent fuel pools and/or concrete dry cask storage

Slows down neutrons to be absorbed by U-235 Made from Water or Graphite (carbon)

## Control Rods

Absorbs neutrons to limit number of chain reactions Made from Boron

## Renewable Energy

|  | Variable Symbol | Unit | Data Booklet Equations: |
| :---: | :---: | :---: | :---: |
| Power | P | W | $\text { Power }=\frac{1}{2} A \rho v^{3}$ |
| Cross-Sectional Area | A | $\mathrm{m}^{2}$ |  |
| Air Density | $\rho$ | $\mathrm{kg} \mathrm{m}^{-3}$ | $\mathrm{A}=\pi r^{2}$ |
| Air Speed | V | $\mathrm{m} \mathrm{s}^{-1}$ |  |
| Photovoltaic Cells | Solar Concentrator |  | Solar Heating Panel |
| Converts solar energy directly into electricity. Useful in solar panels on top of building or solar farms connected to the energy grid | Mirrors focus sunlight onto a central tower. The high thermal energy is converted to steam and runs turbines to produce electricity |  | Sun's radiation is absorbed by black pipes that transfer thermal energy to the water flowing through them. Replaces hot water heater. |


|  | Biomass | Coal | Geothermal | Hydropower | Natural Gas | Nuclear | Petroleum | Solar | Wind |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| Renewable | $\checkmark$ |  | $\checkmark$ | $\checkmark$ |  |  |  | $\checkmark$ | $\checkmark$ |
| Produces $\mathrm{CO}_{2}$ | $\checkmark$ | $\checkmark$ |  |  | $\checkmark$ |  | $\checkmark$ |  |  |

## Thermal Energy Transfer

| Conduction | Convection | Radiation |
| :---: | :---: | :---: |
| Energy is transferred through <br> molecular collisions | Energy circulates through the <br> expansion and rising of hot fluids | Energy is transferred through <br> electromagnetic radiation. Can <br> travel through a vacuum |


|  | Emissivity | Black Body Radiation |
| :---: | :---: | :---: | :---: | :---: |
| Sun | $\sim 1$ |  |
| Earth | $\sim 0.6$ |  |
| An idealized object that |  |  |
| Black-Body | 1 |  |
| absorbs all the |  |  |
| electromagnetic radiation |  |  |
| the falls on it |  |  |


| Power Emissivity | Variable Symbol | Unit |
| :---: | :---: | :---: |
| Power | P | W |
| Emissivity | e | --- |
| Surface Area | A | $\mathrm{m}^{2}$ |
| Temperature | T | K |
| Max Wavelength | $\lambda_{\max }$ | m |

Data Booklet Equations:

$$
\begin{gathered}
P=e \sigma A T^{4} \\
\lambda_{\max }=\frac{2.90 \times 10^{-3}}{T} \\
\sigma=5.67 \times 10^{-8} \mathrm{~W} \mathrm{~m}^{-2} \mathrm{~K}^{-4}
\end{gathered}
$$

## Solar Radiation and Climate Change

| Intensity | Variable Symbol | Unit | Data Booklet Equations: |
| :---: | :---: | :---: | :---: |
| Intensity | I | W m ${ }^{-2}$ | $I=\text { power }$ |
| Power | P | W |  |
| Area | A | $\mathrm{m}^{2}$ | $A_{\text {sphere }}=4 \pi r^{2}$ |
| Greenhouse Gases | Positive Feedback Loop |  | Negative Feedback Loop |
| Water Vapor ( $\mathrm{H}_{2} \mathrm{O}$ ) | Melting ice (decreases albedo) |  | Cloud formation (increases albedo) |
| Carbon Dioxide ( $\mathrm{CO}_{2}$ ) | Melting permafrost (releases methane) |  | Increased photosynthesis (uses $\mathrm{CO}_{2}$ ) |
| Methane ( $\mathrm{CH}_{4}$ ) | Rising ocean temp releases methane |  | Climate Change leads to renewables |

## Astrophysics

## The Scale of Astrophysics

| Unit Conversion | Definition |
| :--- | :---: |
| 1 light year $(\mathrm{ly})=9.46 \times 10^{15} \mathrm{~m}$ | The distance the light travels in an earth year |
| $1 \operatorname{parsec}(\mathrm{pc})=3.26 \mathrm{ly}$ | The average distance between the <br> earth and the sun |
| 1 astronomical unit $(\mathrm{AU})=1.50 \times 10^{11} \mathrm{~m}$ | The distance at which the mean radius of the <br> earth's orbit subtends an angle of 1 arc second |

## Stellar Quantities

| Brightness | Luminosity |
| :---: | :---: |
| Star intensity to an observer on earth | How much total power a star emits |
| Units: $\mathrm{W} \mathrm{m}^{-2}$ | Units: W |


|  | Variable Symbol | Unit | Data Booklet Equations: |
| :---: | :---: | :---: | :---: |
| Distance | d | pc |  |
| Parallax Angle | p | sec | (arc-second $)$ |
| Brightness | b | W m ${ }^{-2}$ | $=\frac{L}{}$ |
| Luminosity | L | W |  |
| Max Wavelength | $\lambda_{\text {max }}$ | m | $\lambda_{\text {max }} T=2.9 \times 10^{-3} \mathrm{~m} \mathrm{~K}$ |
| Temperature | T | K | $L=\sigma A T^{4}$ |
| Surface Area | A | $\mathrm{m}^{2}$ | $\sigma=5.67 \times 10^{-8} \mathrm{~W} \mathrm{~m}^{-2} \mathrm{~K}^{-4}$ |

Describe the process of Stellar Parallax: Observe how far a star moves relative to distant stars six months apart so that earth has its maximum displacement and an angle can be measured


Atomic Spectra

|  | Variable Symbol | Unit |
| :---: | :---: | :---: |
| Energy | E | J or eV |
| Planck's Constant | h | J s |
| Frequency | $f$ | Hz |
| Speed of Light | c | $\mathrm{m} \mathrm{s}^{-1}$ |
| Wavelength | $\lambda$ | m |

Data Booklet Equations:

$$
\begin{gathered}
E=h f \\
\lambda=\frac{h c}{E} \\
h=6.63 \times 10^{-34} \mathrm{~J} \mathrm{~s} \\
c=3.00 \times 10^{8} \mathrm{~m} \mathrm{~s}^{-1}
\end{gathered}
$$

hc

$$
1.99 \times 10^{-25} \mathrm{~J} \mathrm{~m}
$$

$1.24 \times 10^{-6} \mathrm{eV} \mathrm{m}$

## H-R Diagrams and Life Cycle of a Star

Label the Following:

- Main Sequence
- White Dwarfs
$\square$ Red Giants
$\square$ The Sun
$\square$ Line representing the life cycle of our sun


| Chandrasekhar Limit |  | Oppenheimer-Volkhoff Limit |  |
| :---: | :---: | :---: | :---: |
| The maximum mass of a core that can become a white dwarf is 1.4 times the mass of the sun $\left(1.4 \mathrm{M}_{\odot}\right)$ |  | The maximum mass of a core that can become a neutron star is 3 times the mass of the sun ( $3 \mathrm{M}_{\odot}$ ) |  |
| Sun Like Stars (<1.5 M ${ }^{\text {) }}$ ) <br> $\downarrow$ <br> White Dwarf | $\begin{aligned} & \text { Huge Stars }\left(1.5-3 \mathrm{M}_{0}\right) \\ & \downarrow \end{aligned}$ |  | $\begin{aligned} & \text { Giant Stars }\left(>3 \mathrm{M}_{0}\right) \\ & \quad \end{aligned}$ <br> Black Hole |

## The Expanding Universe

| Standard Candles | Evidence for Expanding Universe |
| :---: | :---: |
| Objects of known luminosity that can be <br> used with the apparent brightness to <br> measure distance from earth | Hubble discovered that the farther away <br> stars and galaxies are, the more their light <br> is redshifted. |
| Cephid Variables and Type Ia Supernovas | This means, more distant objects are <br> traveling faster than nearer objects. |


|  | Variable Symbol | Unit |
| :---: | :---: | :---: |
| Redshift | z | --- |
| Change in Wavelength | $\Delta \lambda$ | m |
| Original Wavelength | $\lambda_{0}$ | m |
| Relative Velocity of Source | V | $\mathrm{m} \mathrm{s}^{-1}$ |
| Speed of Light | C | $\mathrm{m} \mathrm{s}^{-1}$ |
| Current Scale Factor | R | --- |
| Scale Factor when Emitted | $\mathrm{R}_{0}$ | --- |
| Hubble's Constant | $\mathrm{H}_{0}$ | $\mathrm{~km} \mathrm{~s}^{-1} \mathrm{Mpc}^{-1}$ |

Data Booklet Equations:

$$
\begin{gathered}
z=\frac{\Delta \lambda}{\lambda_{0}} \approx \frac{v}{c} \\
z=\frac{R}{R_{0}}-1 \\
v=H_{0} d \\
c=3.00 \times 10^{8} \mathrm{~m} \mathrm{~s}^{-1} \\
H_{0} \approx 70 \mathrm{~km} \mathrm{~s}^{-1} \mathrm{Mpc}^{-1}
\end{gathered}
$$

## The Big Bang

|  | Peak Wavelength | Temperature |
| :---: | :---: | :---: |
| Cosmic Microwave <br> Background Radiation | $\sim 0.001 \mathrm{~m}(1 \mathrm{~mm})$ | $\sim 2.9 \mathrm{~K}$ |

## Describe why the CMB is evidence of the Big Bang:

The CMB is the heat signature from the early universe. As the universe has expanded to its current size, the wavelength stretched out to the current value seen in the CMB. This radiation is fairly uniform and it can be observed in every direction.

